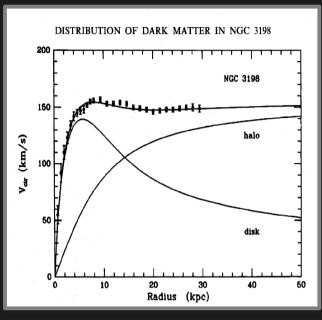
Dark matter

Maxim Laletin STAR Workshop 15/09/2017

OUTLINE

- Observational evidence
- Hypotheses
- Search strategies
- What's new?

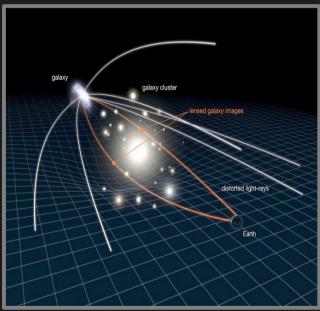
OBSERVATIONAL EVIDENCE



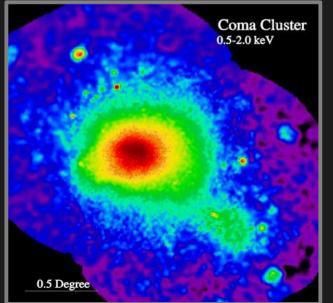
Galaxy rotation curves



Motion of galaxies
In clusters

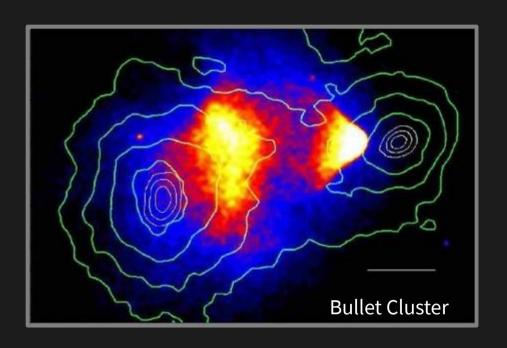


Gravitational lensing of clusters



X-ray Observations Of clusters

OBSERVATIONAL EVIDENCE



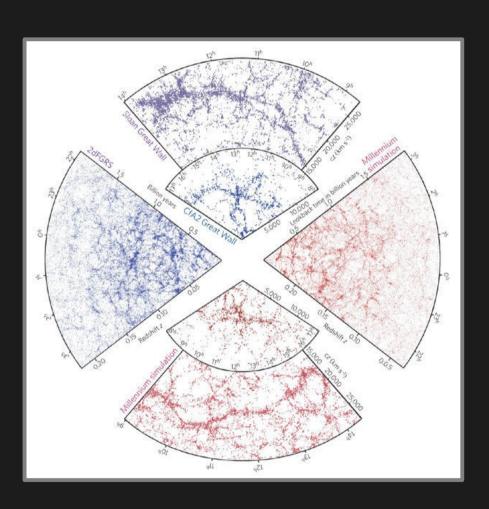




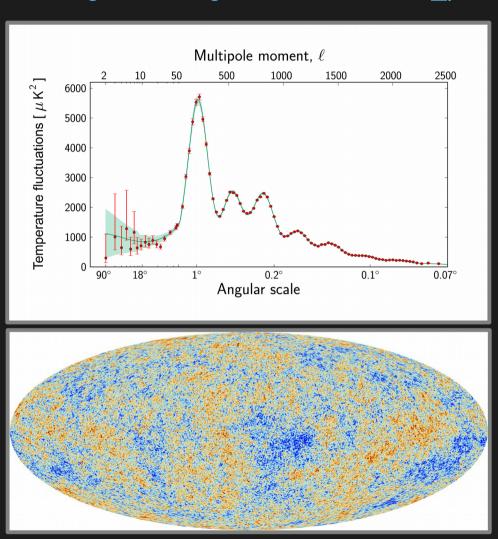
Separation of matter in cluster collisions

OBSERVATIONAL EVIDENCE

https://lambda.gsfc.nasa.gov/education/cmb_plotter/



Large scale structure of the Universe



Cosmic microwave background power spectrum

What do we know about dark matter from observations?

- It gravitates;
- It is dark (doesn't radiate much, probably doesn't even interact electromagnetically);
- It makes up ~ 26% of the energy density in the Universe;
- Cosmologically stable (or long-lived);
- It is **collisionless** (or doesn't collide much)

SOME CANDIDATES

Weakly Interacting Massive Particles (WIMPs):

The most trivial and well elaborated group of models; many beyond SM physics models predict such particles (e.g. SUSY, Inert Doublet Models, sterile neutrino, etc.)

Particles with non-trivial interactions:

Various particle physics models providing some features that WIMPs don't have, such as sizable self-interaction, complicated production mechanism or thermal evolution and so on (e.g. axions, mirror DM, composite DM, FIMPs, SIMPs, PIMPs, etc.)

Primordial black holes (PBH)

A new wave of interest in PBH as DM candidate started recently due to LIGO's observation of gravitational waves from merging black holes.

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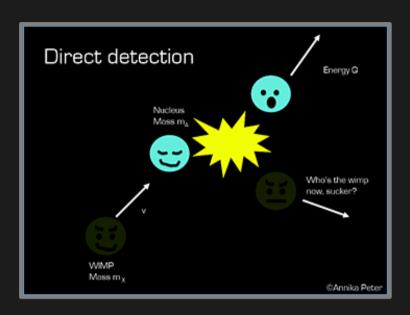
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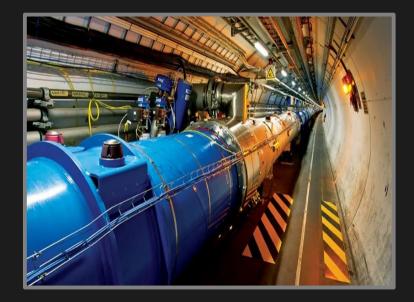
SEARCH STRATEGIES

Direct detection



Indirect detection

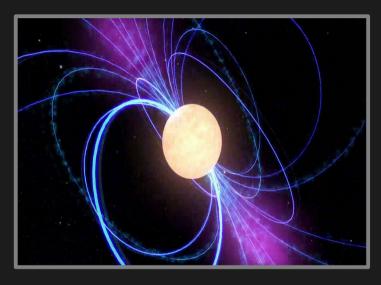




Production at colliders

SEARCH STRATEGIES

In indirect searches astrophysics is really involved...





C. Kouvaris, M. Angeles Perez-Garcia,

"Can Dark Matter explain the Braking Index of Neutron Stars?", Phys.Rev. D89 (2014) no.10, 103539, 1401.3644

F. Contenta et al.,

"Probing dark matter with star clusters: a dark matter core in the ultra-faint dwarf Eridanus II", 1705.01820

L. Gabriel Gómez et al.,

"Dark-matter dynamical friction versus gravitational-wave emission in the evolution of compact-star binaries", Phys.Rev. D96 (2017) no.6, 063001, 1706.06801

A. Khmelnitsky, V. Rubakov,

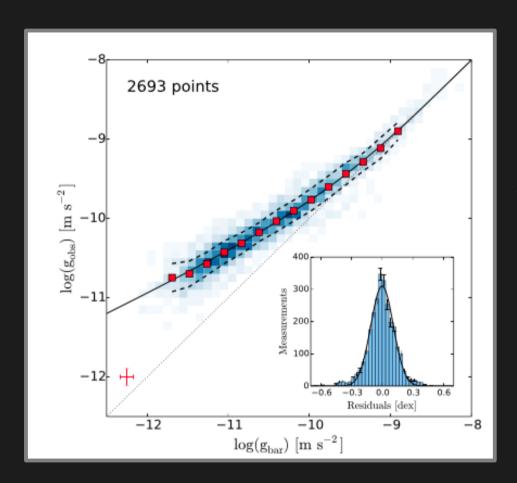
"Pulsar timing signal from ultralight scalar dark matter", JCAP 1402 (2014) 019, 1309.5888

L. Tolos, J. Schaffner-Bielich,

"Dark Compact Planets",

Phys.Rev. D92 (2015) 123002, 1507.08197

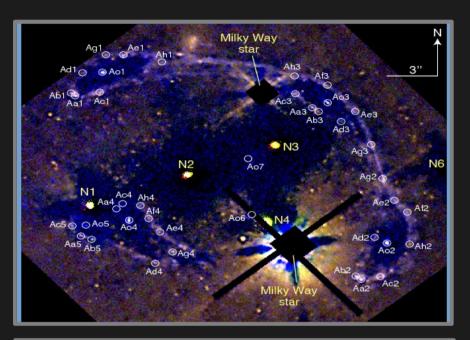
WHAT'S NEW?



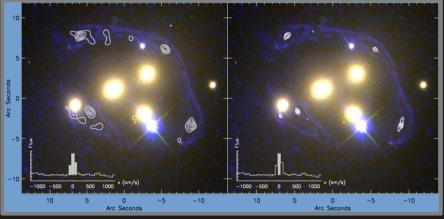
Strong relation between radial acceleration traced by rotation curves and that predicted by the observed distribution of baryons was **observed**.

(do not confuse with Tully-Fischer relation)

WHAT'S NEW?



In 2015 Massey et al. found a discrepancy between the observed mass distribution in Abell 3827 galaxy cluster and that predicted within standard cold DM model. They attributed the effect to self-interacting dark matter.



The **refined analysis** from 2017 revealed that the dynamics in this cluster **consistent** with standard cold dark matter.

D. Harvey et al., "Dark matter dynamics in Abell 3827: new data consistent with standard Cold Dark Matter", 1708.04245

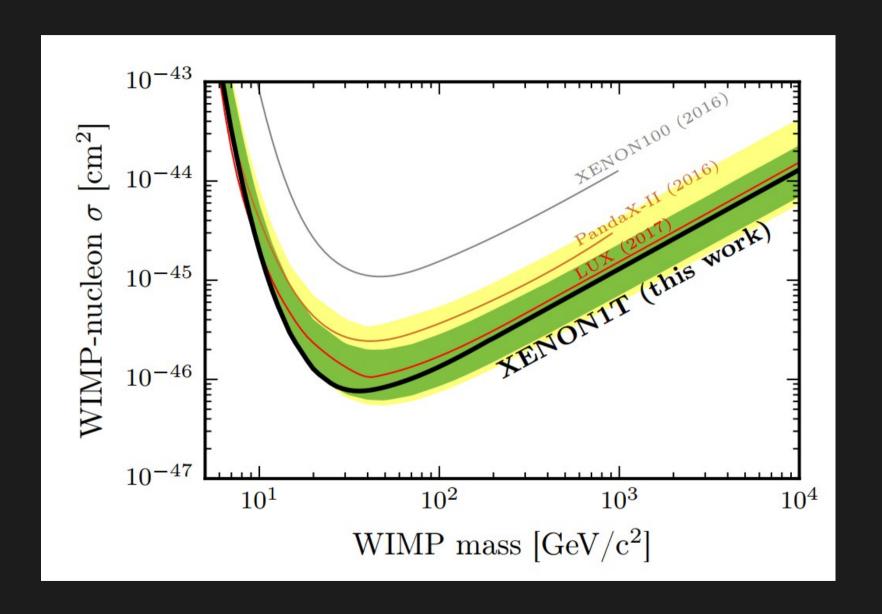
CONCLUSION

- Multiple observational **evidence** indicate the presence of new physics on different scales, which can be associated with the existence of **dark matter**.
- But we still don't know the nature of DM. Direct, indirect and collider searches give null results.
- On the other hand, we know pretty much about what dark matter is not.
- We're trying hard to reveal at least a little more **properties** of dark matter. Detailed study of **astrophysical systems** (stars, clusters, pulsars, etc.) is one of the ways to do it.

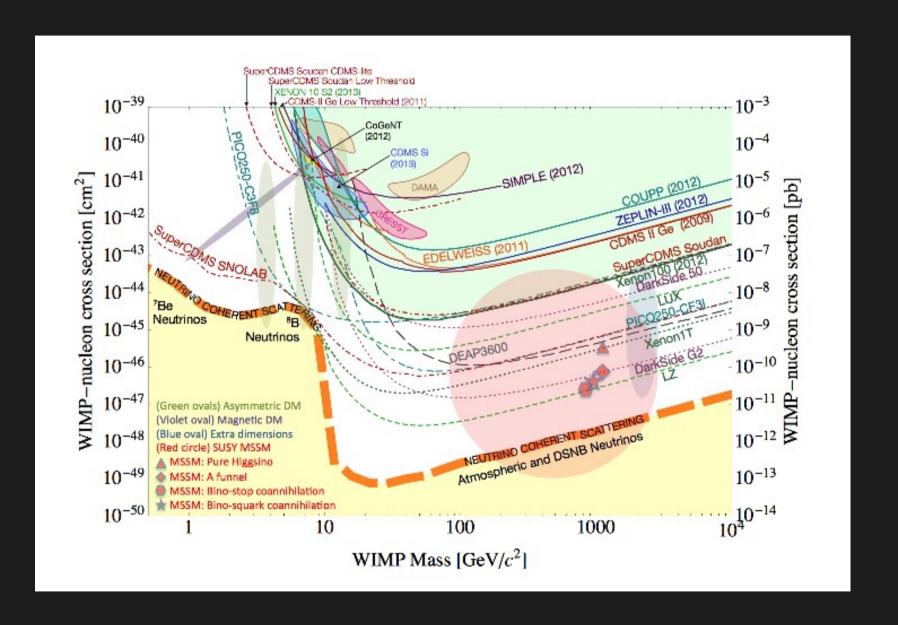
Thank you for attention!

BACKUP

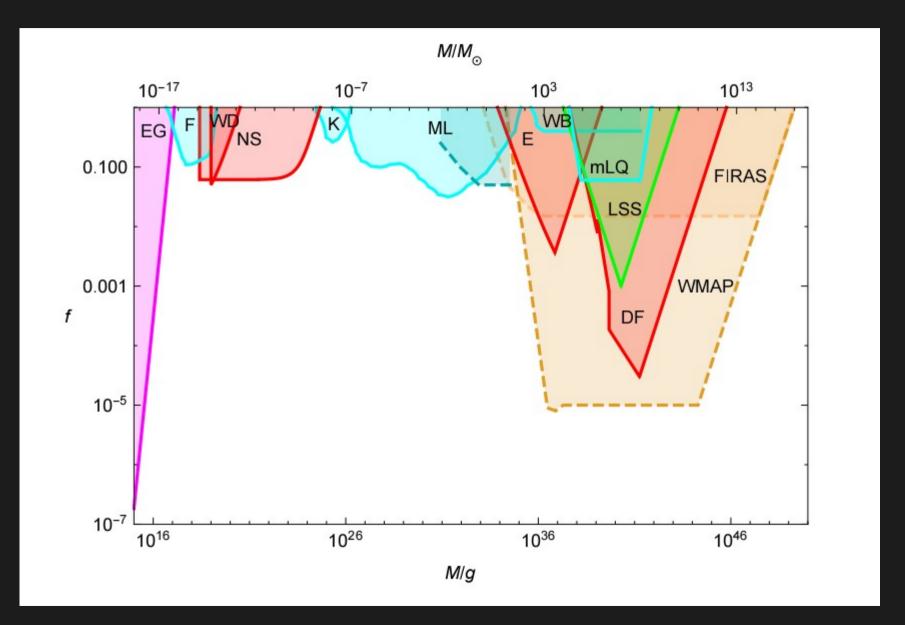
Direct detection results



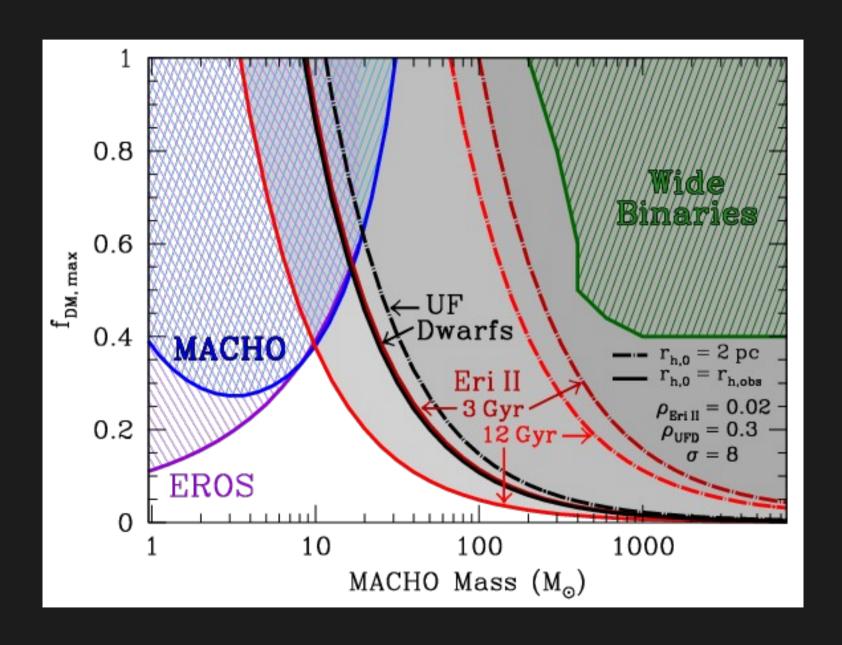
Direct detection results



Constraints on PBH



Constraints on MACHOs



Big Bang nucleosynthesis

