

LLAMA and obscuration

Local Luminous AGN with Matched Analogs

Leonard Burtscher, Ric Davies, Ming-yi Lin,
Gilles Orban de Xivry, David Rosario et al.

Max-Planck-Institut für extraterrestrische Physik
Garching

 @LeoBurtscher



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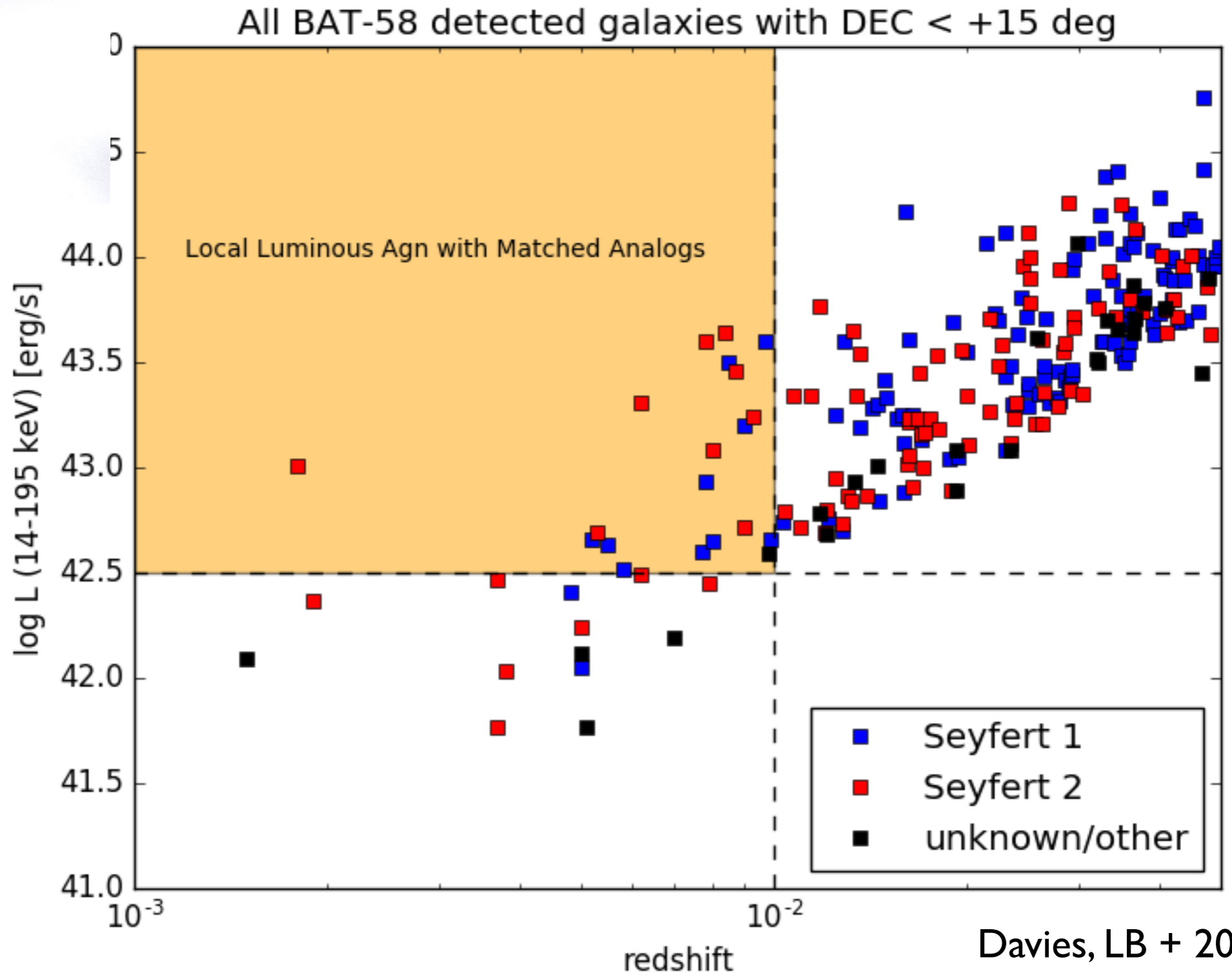
 @LeoBurtscher





Sample and rationale

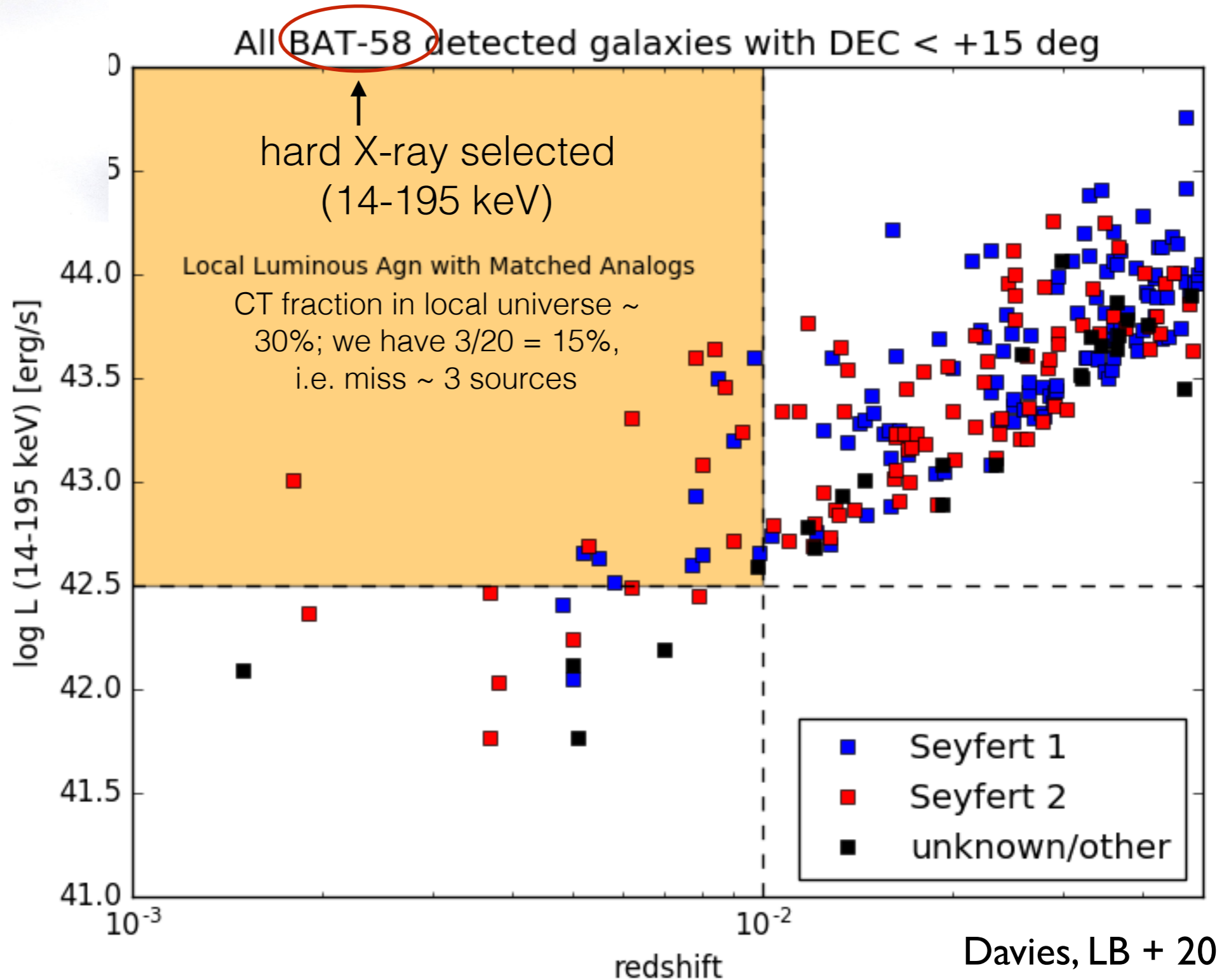
A complete, hard X-ray selected local sample





Sample and rationale

A complete, hard X-ray selected local sample

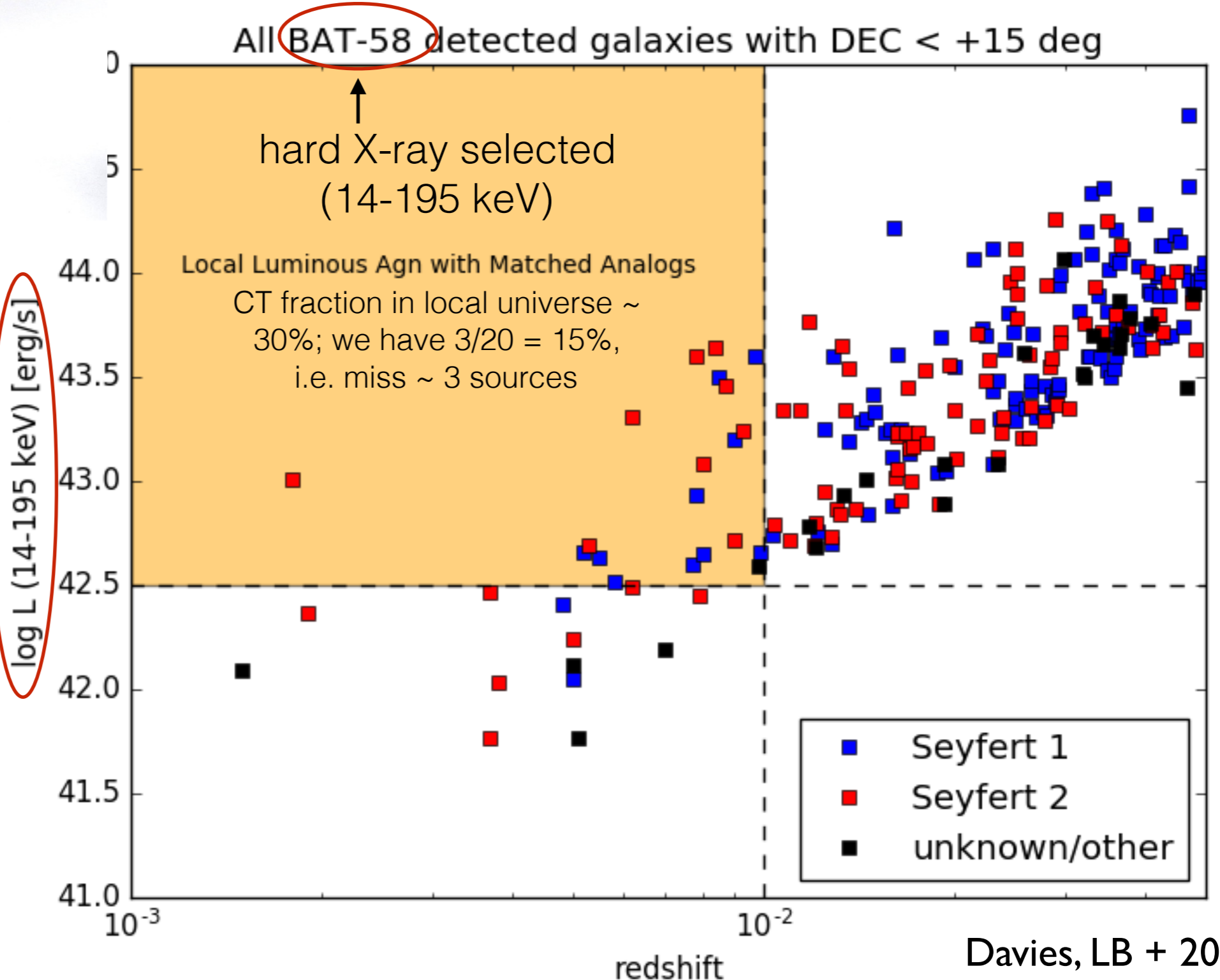




Sample and rationale

A complete, hard X-ray selected local sample

The most luminous local sources: AGN activity is more than just „weather“

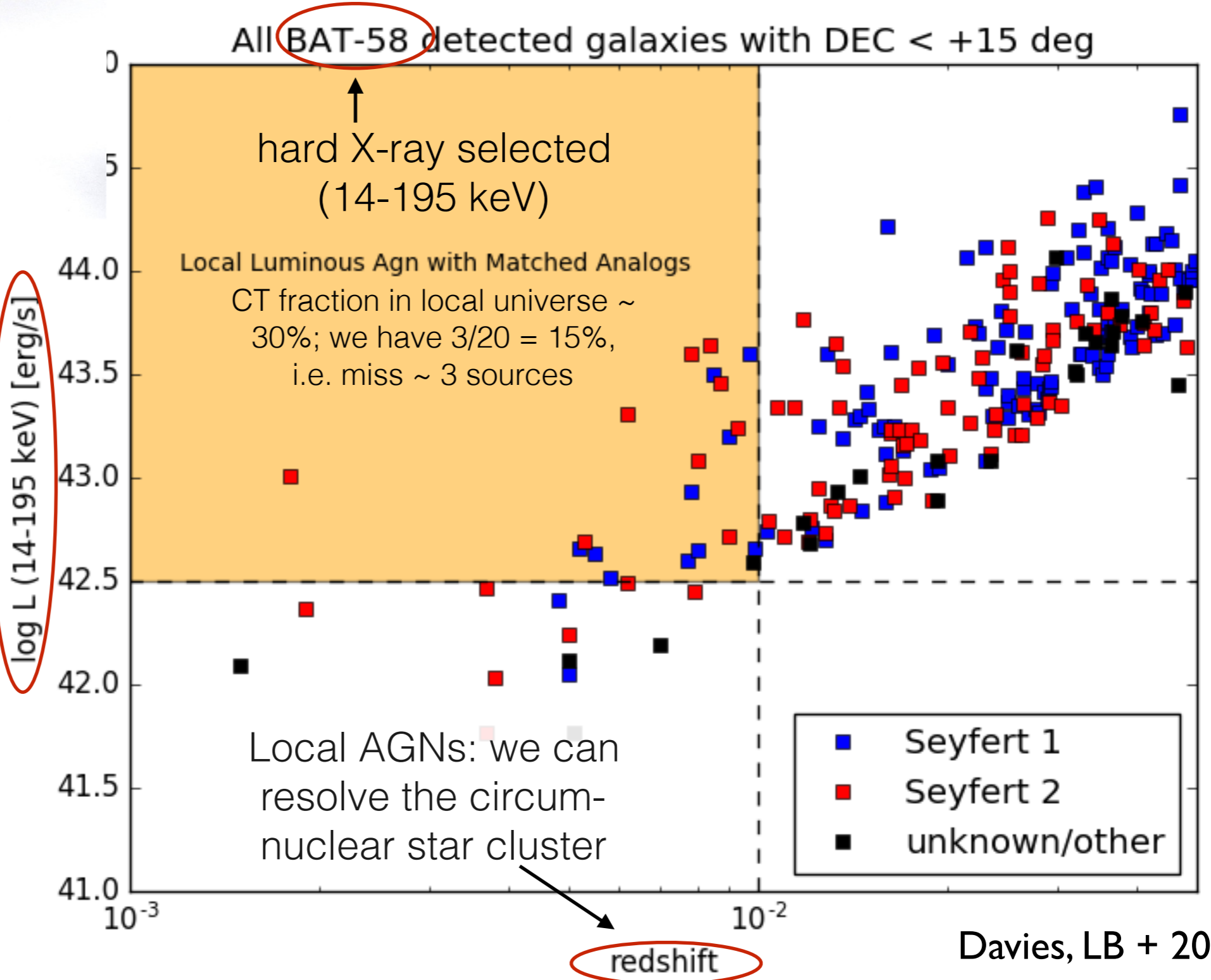




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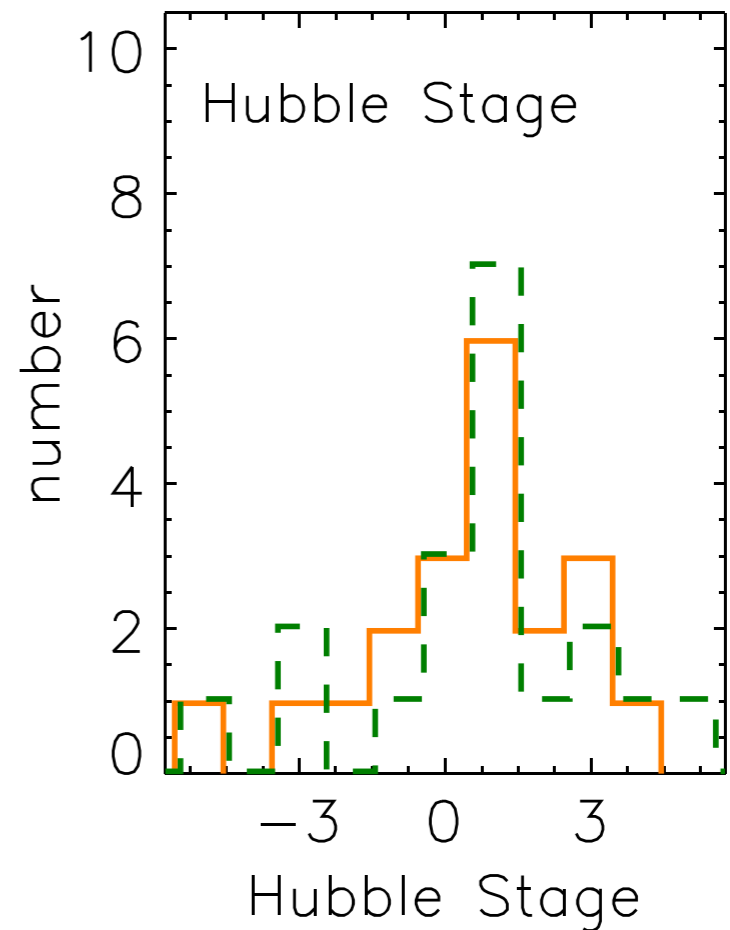
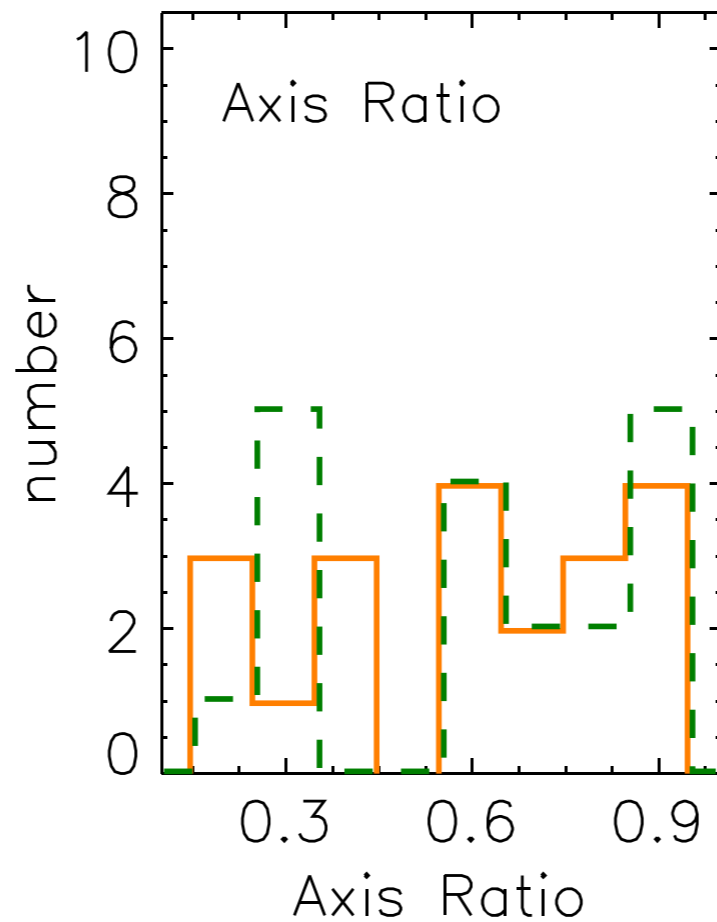
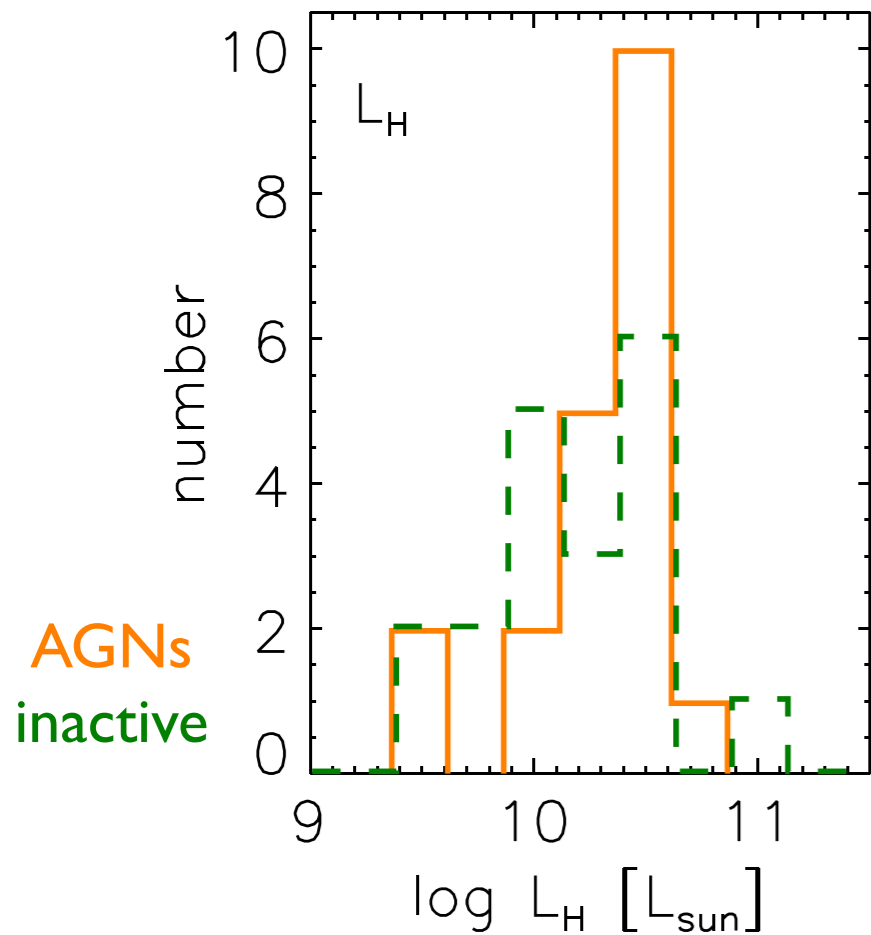




Sample and rationale

A complete, hard X-ray selected local sample

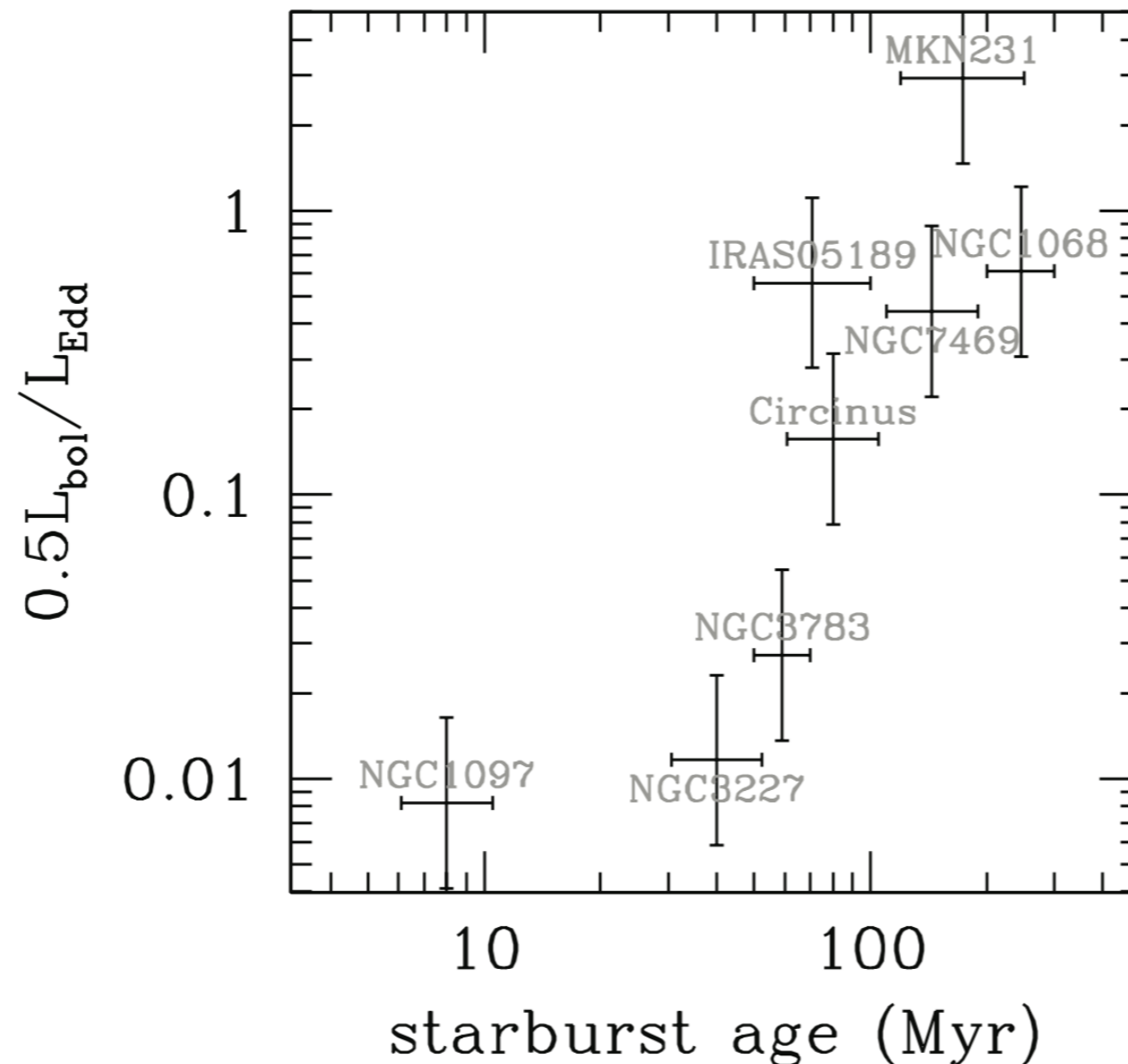
All **BAT-58** detected galaxies with DEC < +15 deg



10^{-3} 10^{-2}
 ↘
redshift

The torus: more than obscuration

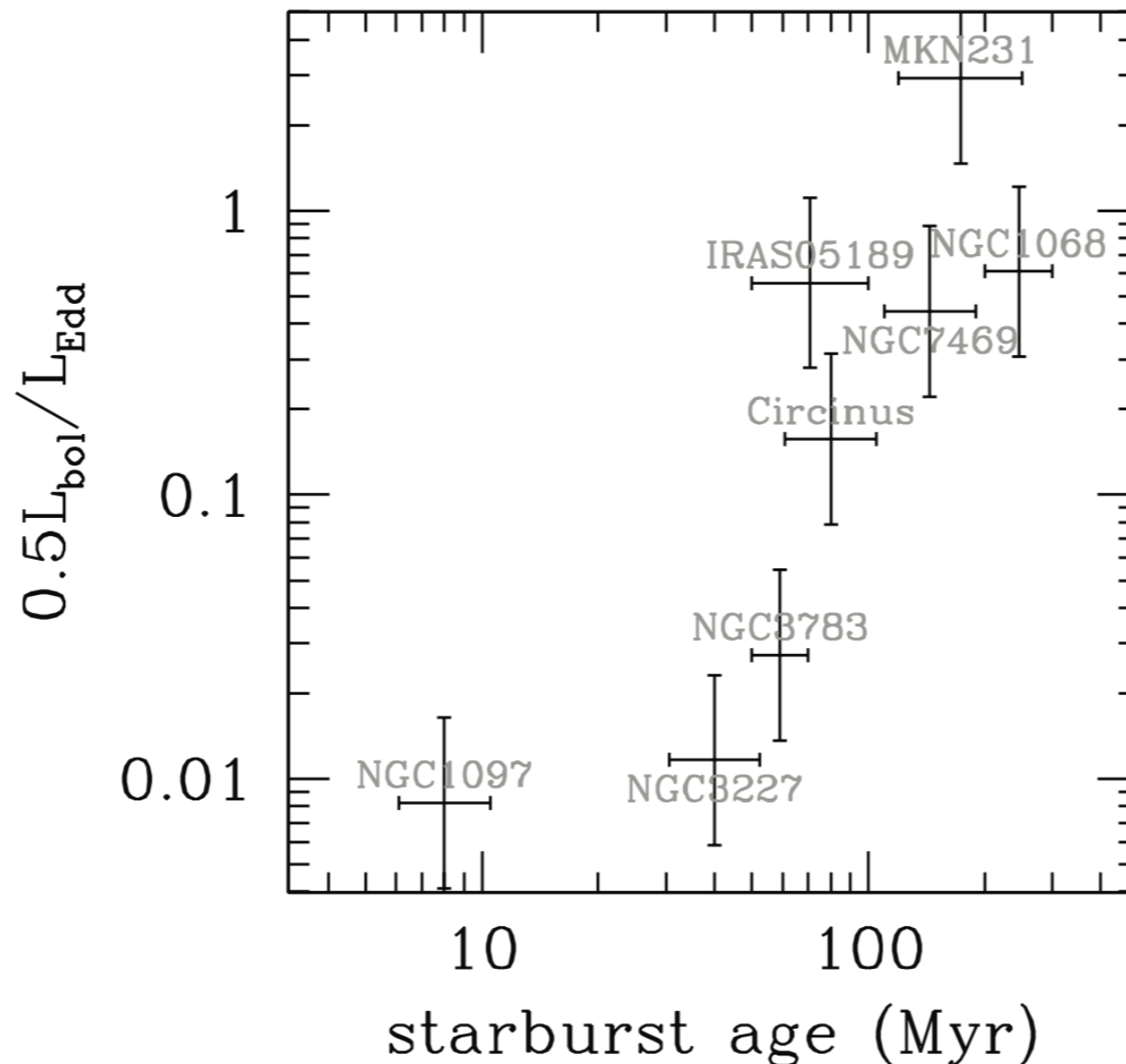
A starburst-AGN connection?



Davies+ 2007

The torus: more than obscuration

A starburst-AGN connection?



powerful AGN
activity only in
post-starburst
nuclei?

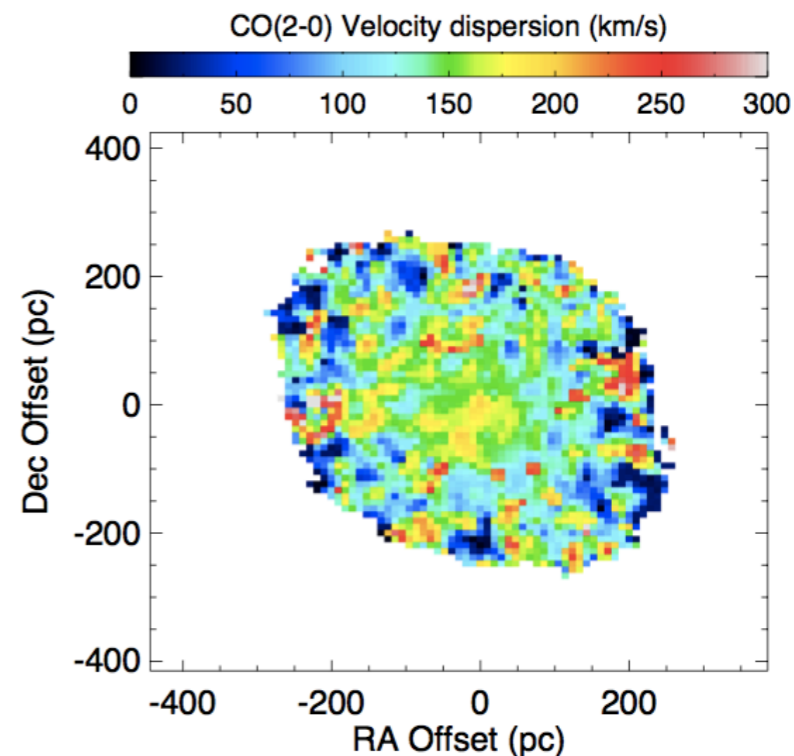
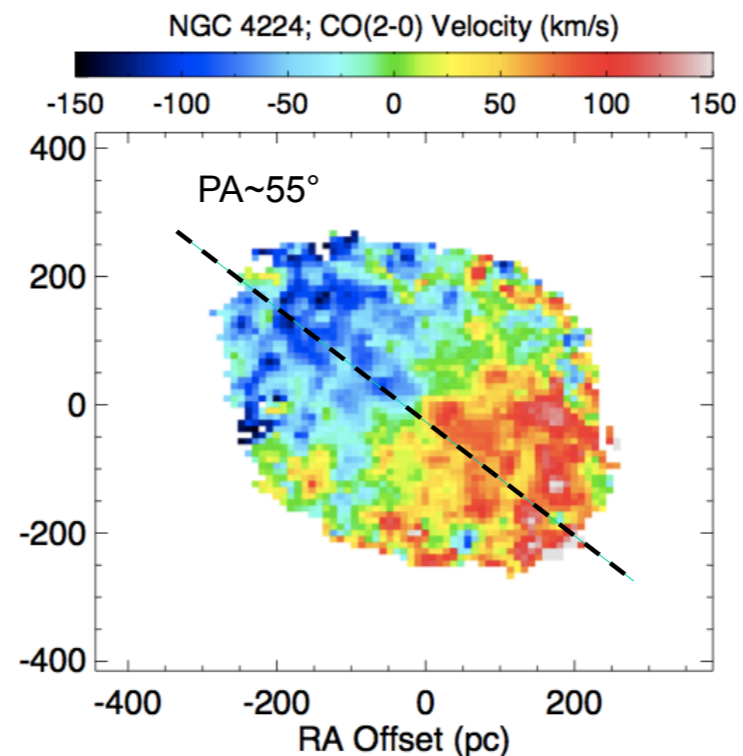
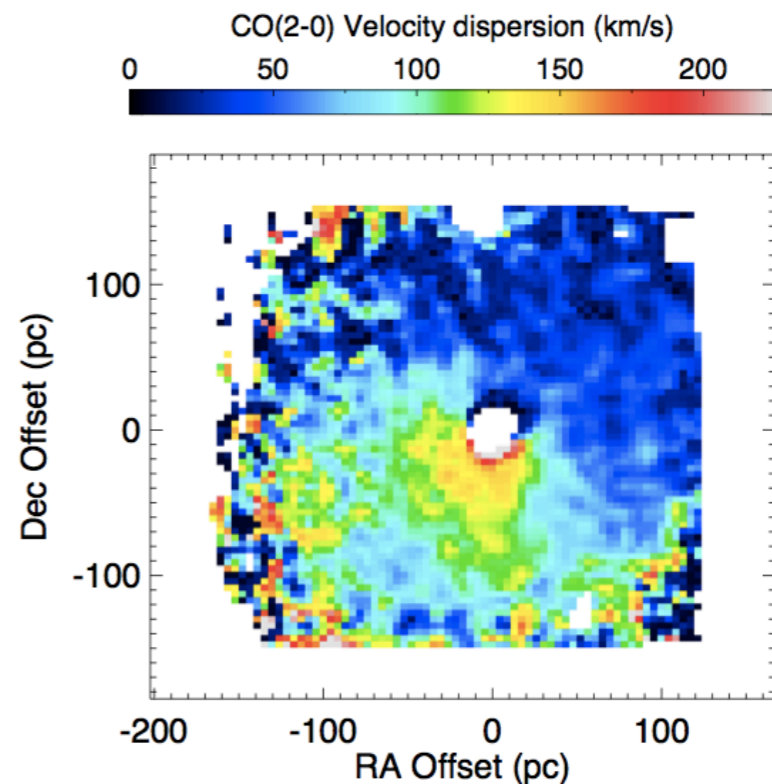
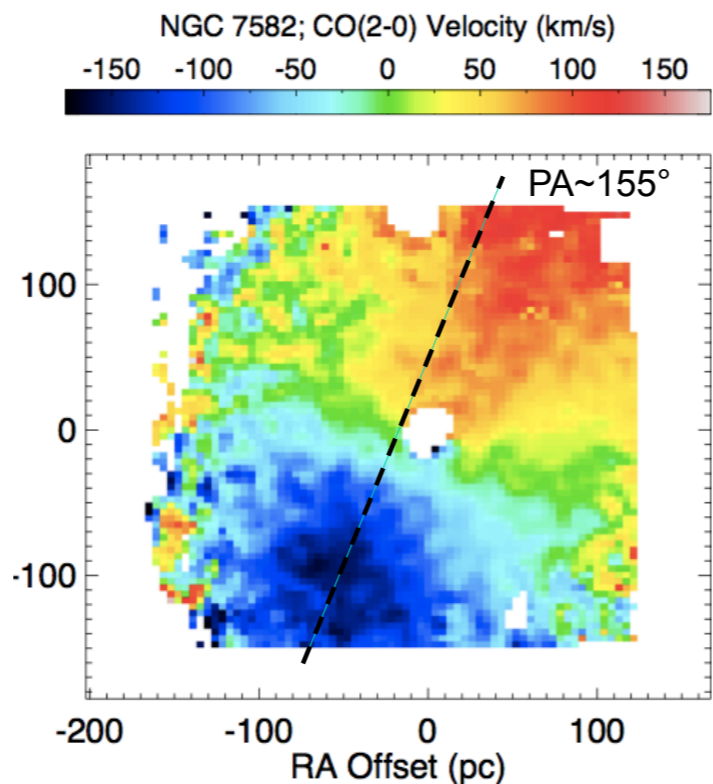
Davies+ 2007



- SINFONI IFU cubes to analyze gas inflow / outflow

The next steps

A complete, hard X-ray selected local sample



Lin+ (in prep)

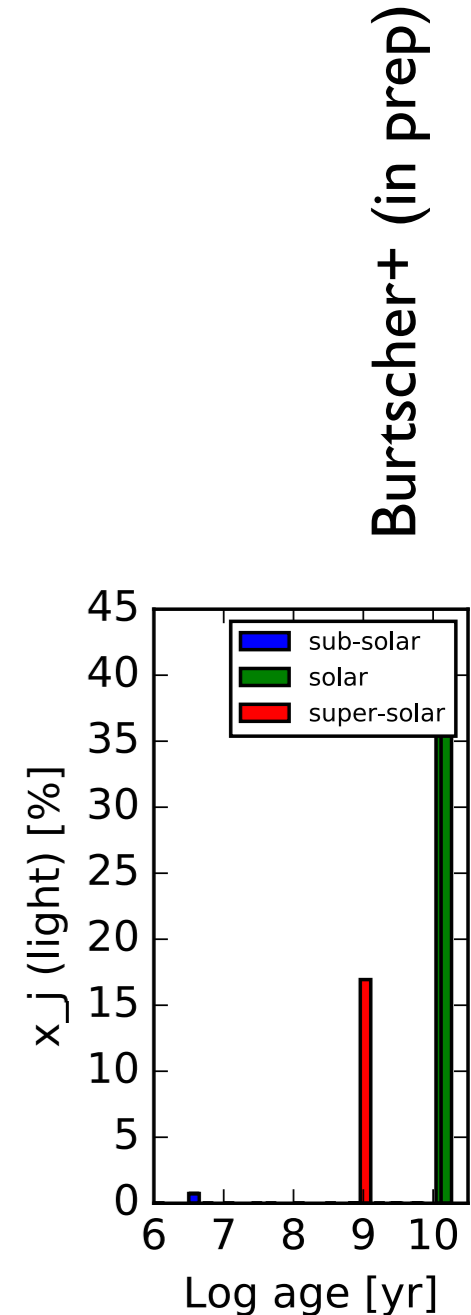
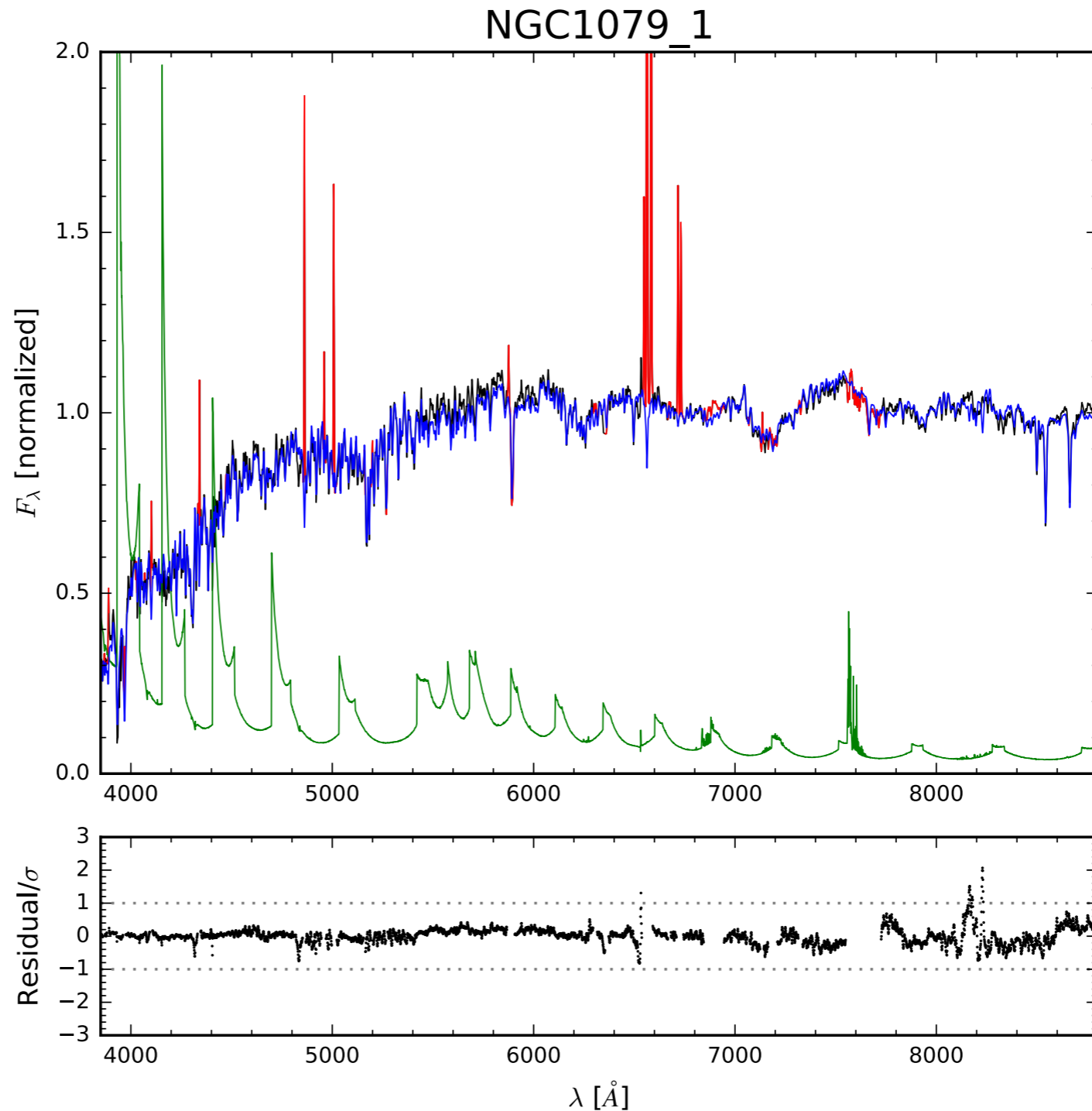
see poster B9 by Ming-Yi Lin

The next steps

A complete, hard X-ray selected local sample



- SINFONI IFU cubes to analyze gas inflow / outflow
- X-SHOOTER spectra to robustly analyze the star formation histories

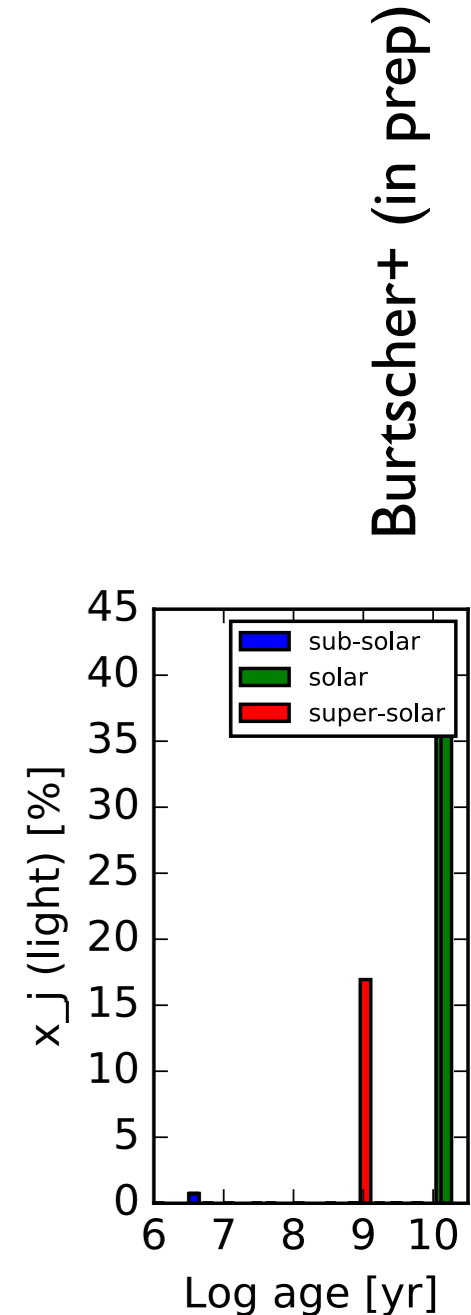
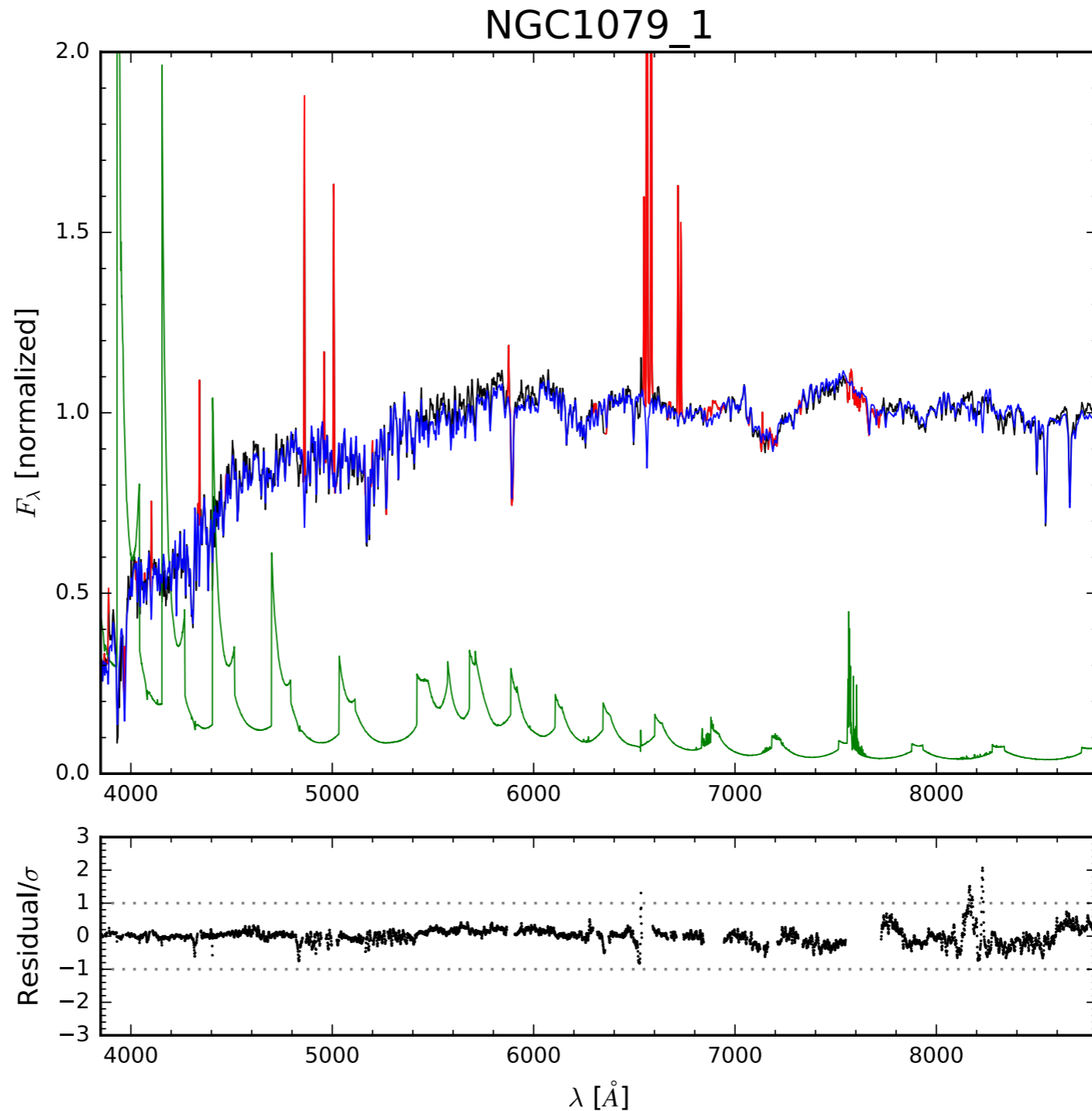


The next steps

A complete, hard X-ray selected local sample

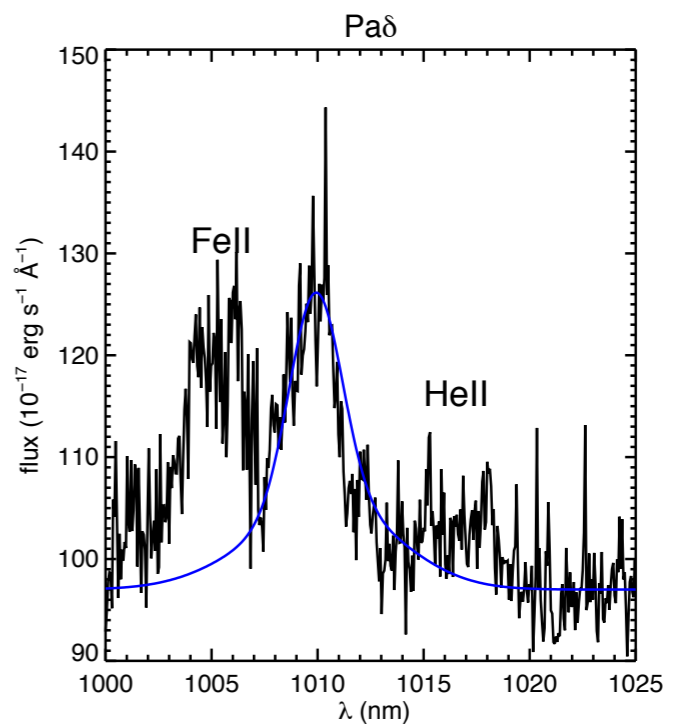
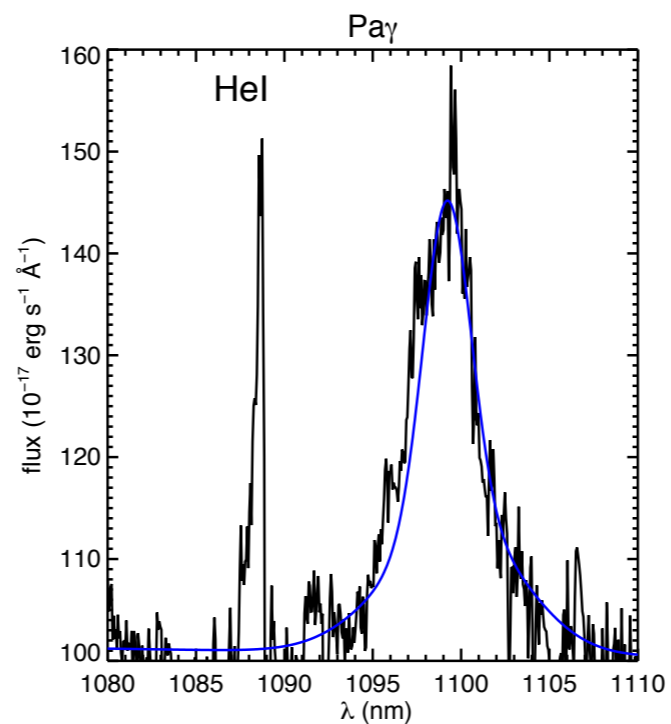
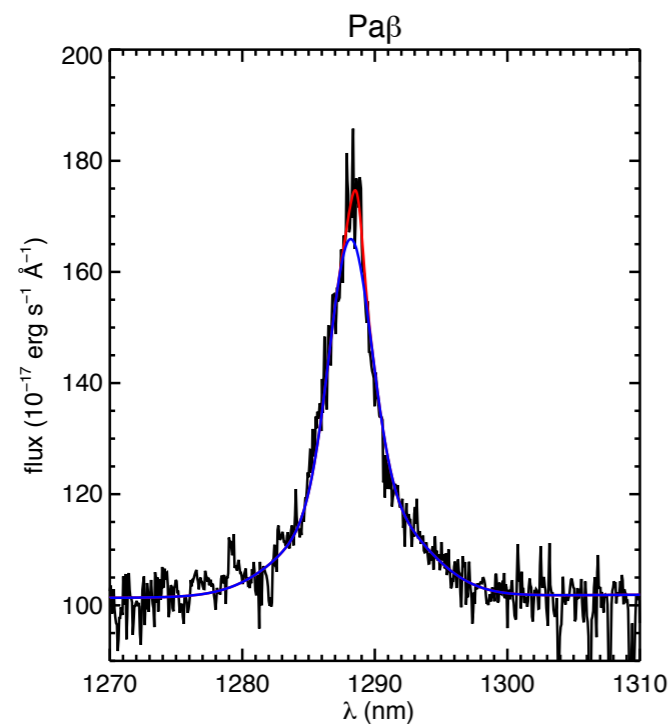
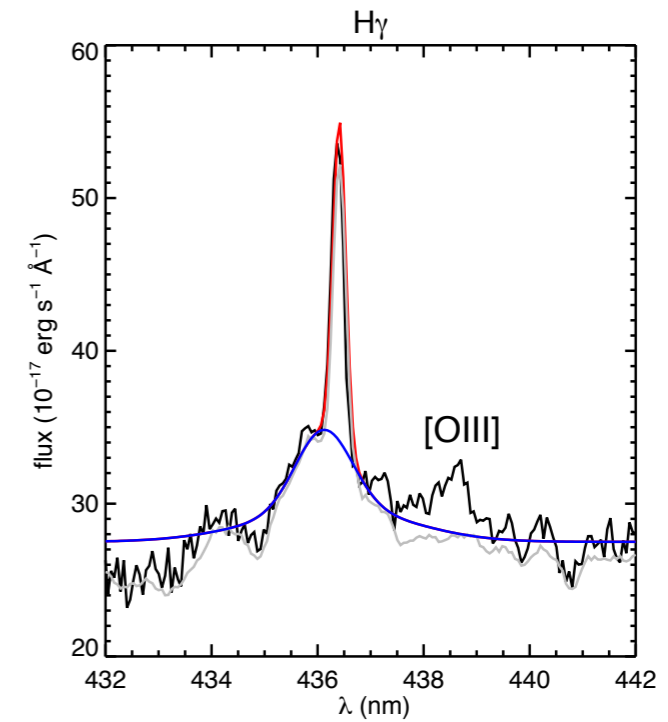
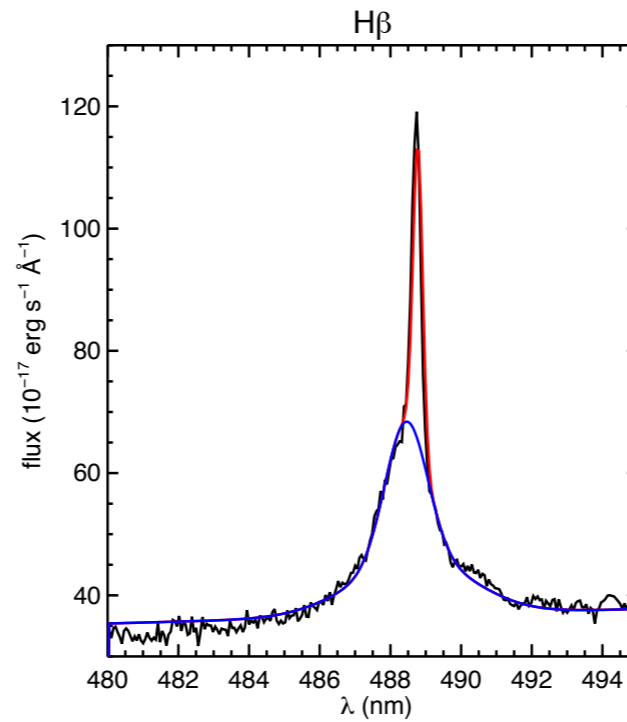
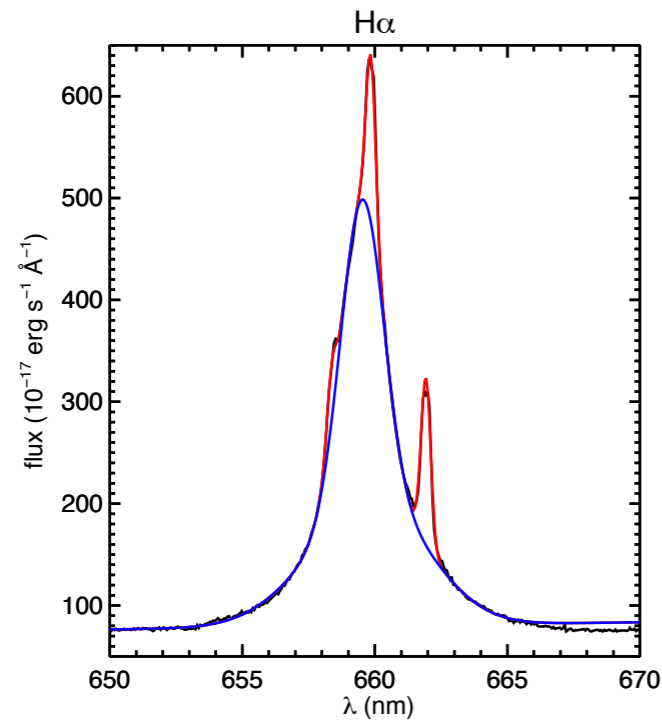


- SINFONI IFU cubes to analyze gas inflow / outflow
- X-SHOOTER spectra to robustly analyze the star formation histories
- APEX data to probe molecular inventory (+ trying to get ALMA + HST...)

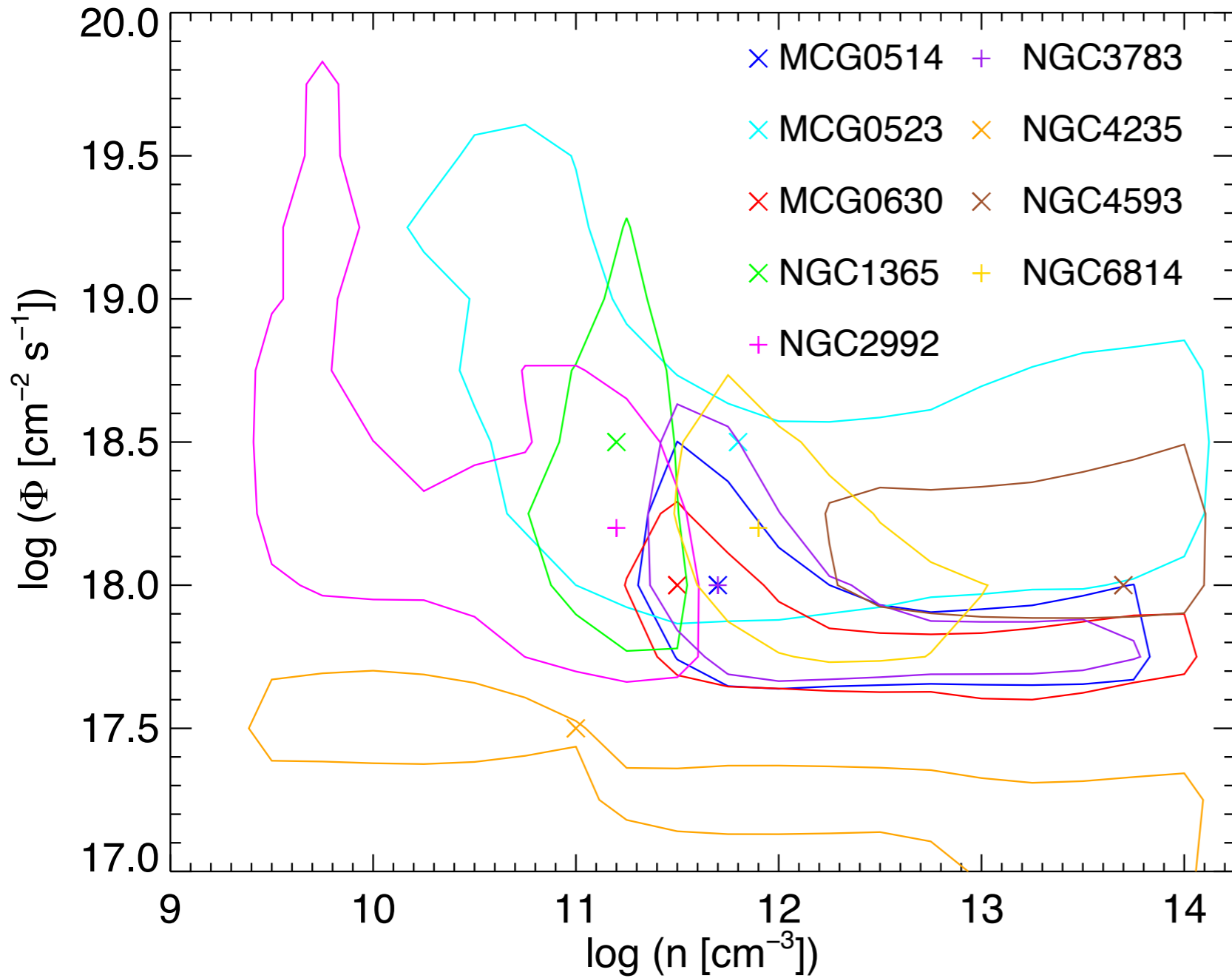


First results: BLR properties and obscuration

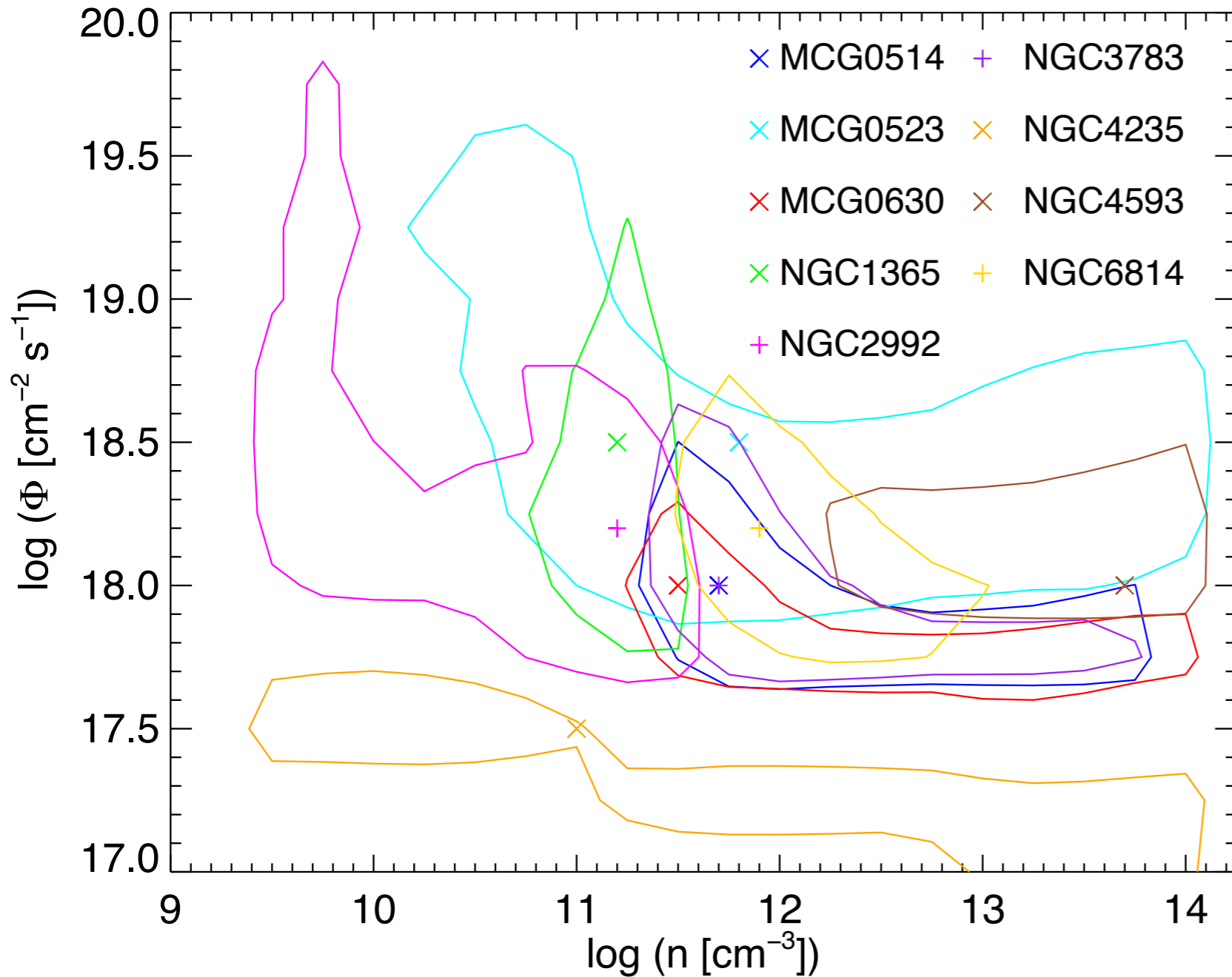
NGC1365



First results: BLR properties and obscuration



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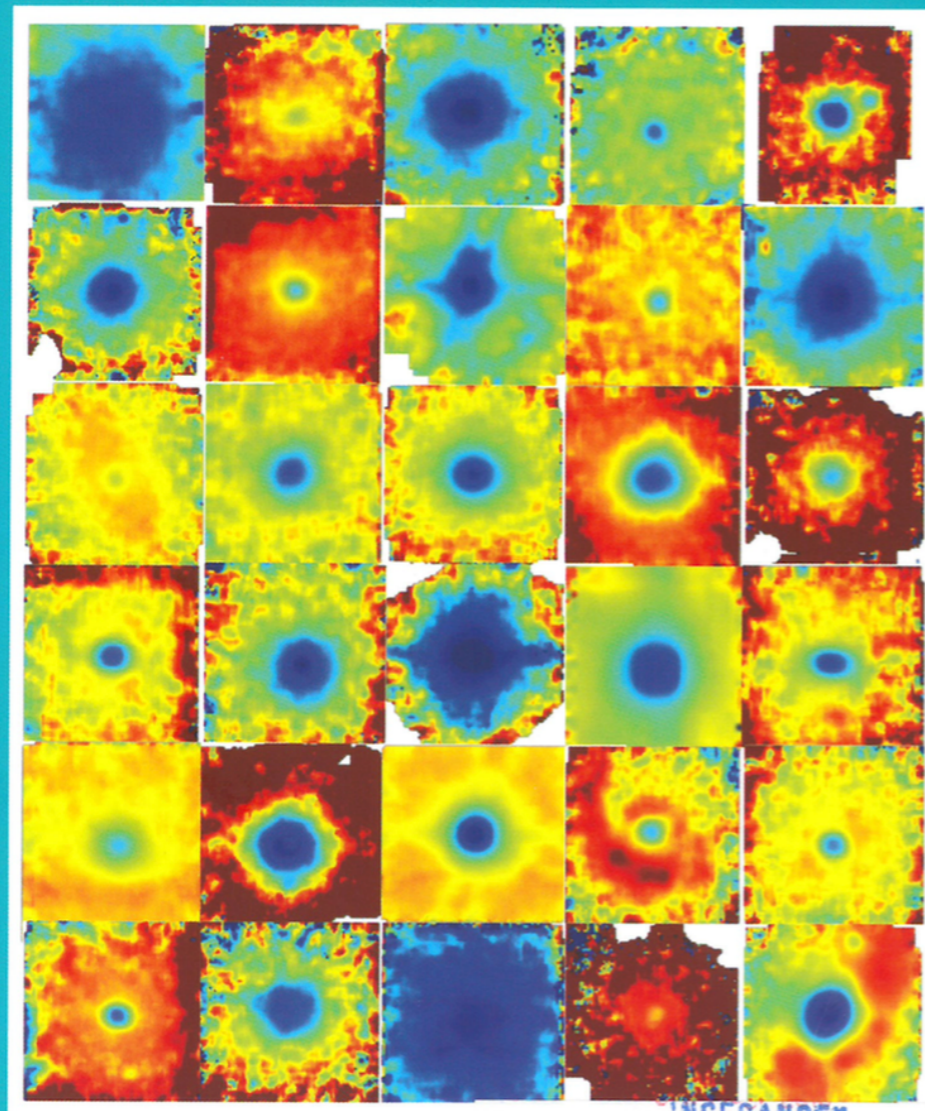


Hydrogen emitting
clouds are \sim at
sublimation radius

Object	r_b
MCG-05-14-012	1.0
MCG-05-23-16	0.5
MCG-06-30-015	1.0
NGC1365	0.7
NGC2992	0.8
NGC3783	0.8
NGC4235	1.8
NGC4593	0.9
NGC6814	0.9

Astronomy & Astrophysics

A near-IR high-resolution atlas of local AGNs



Burtscher+ 2015

Dilution of stellar light by the AGN continuum

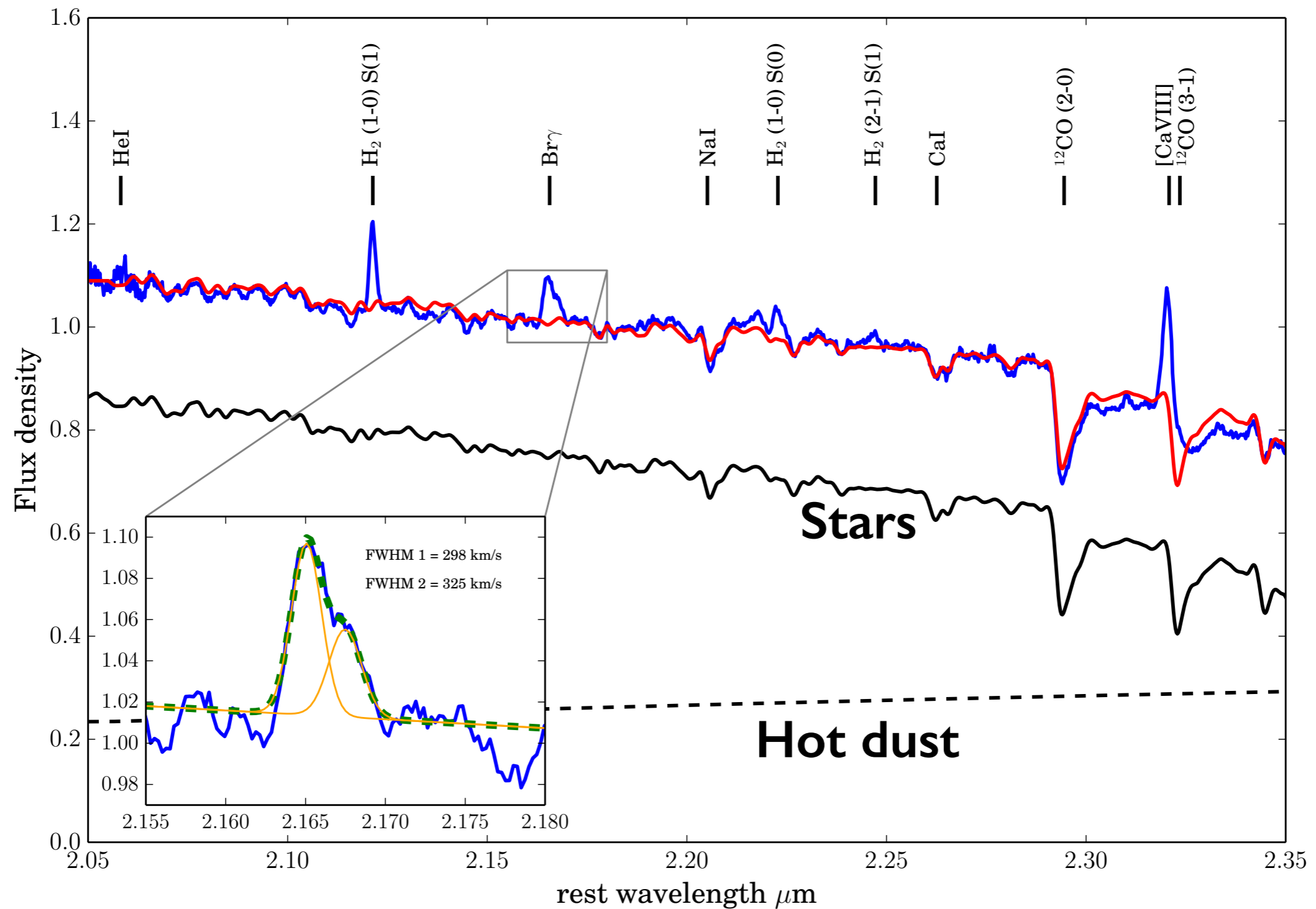
Vol. 578 - Part I
JUNE • 2015

A&A Online: <http://www.aanda.org> <http://cds.aanda.org>

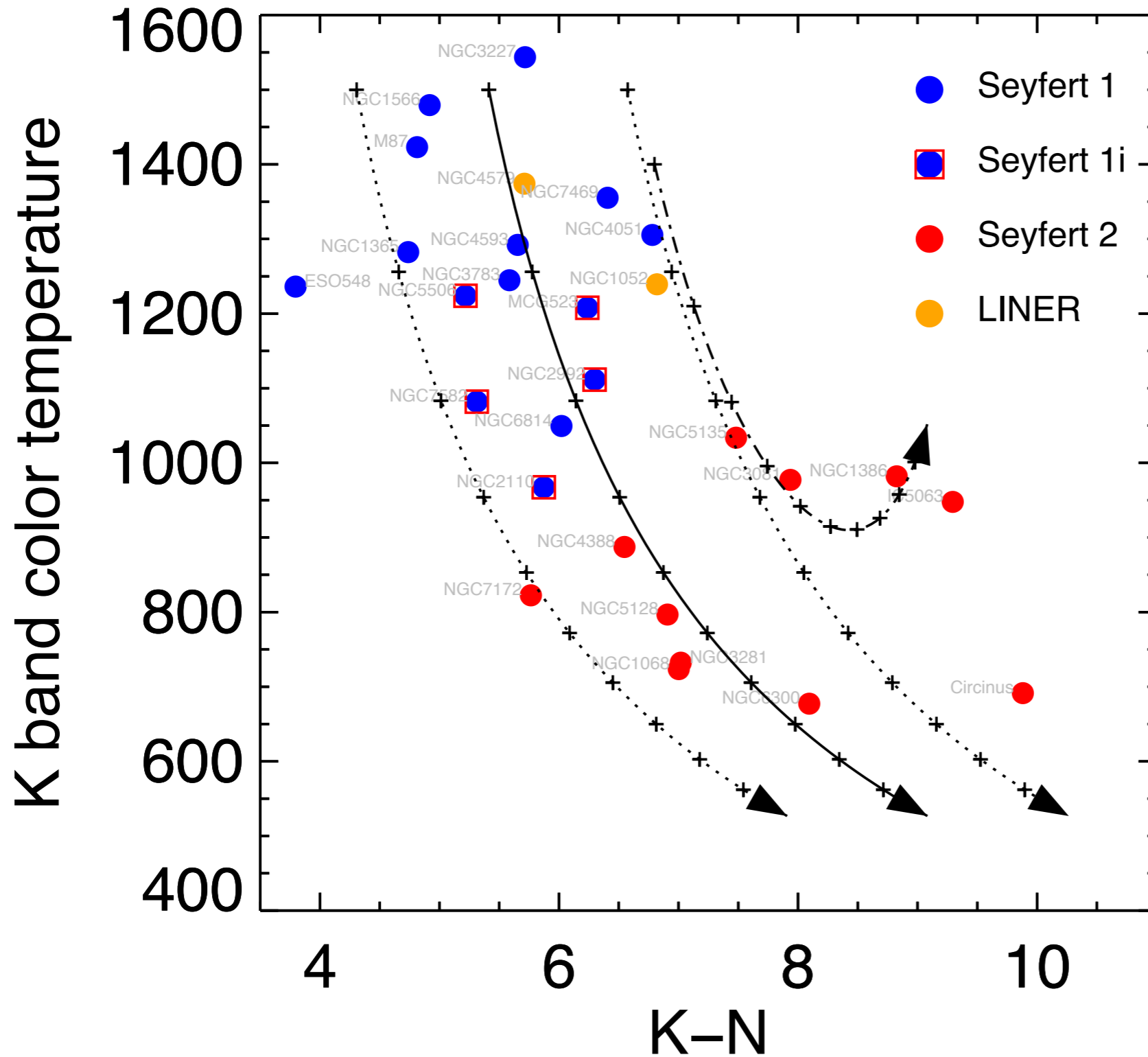
ISSN 0004-6361 • 578 - Part I • L1-L6/A1-A68/C1 • June 2015

Spectral decomposition

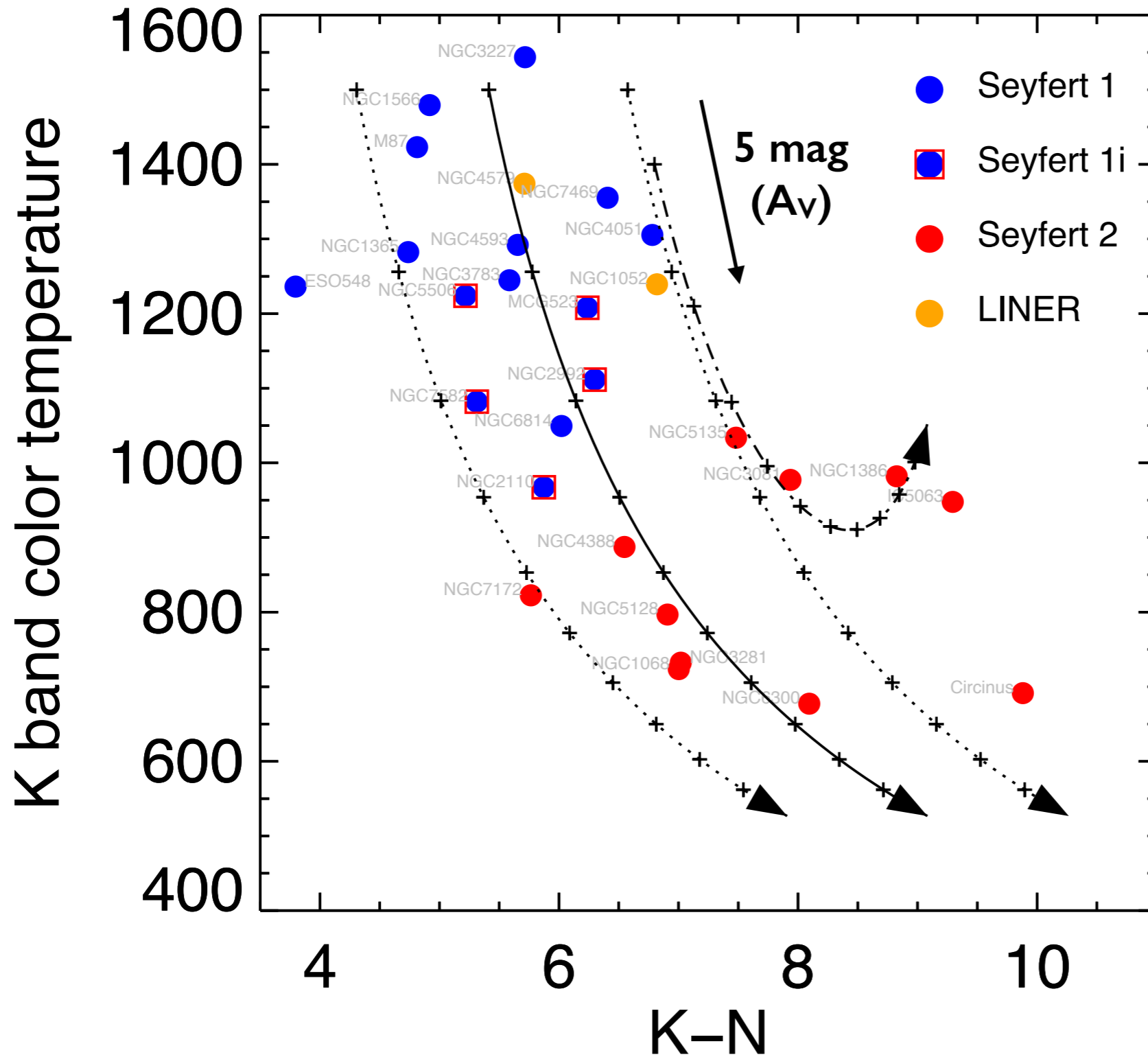
NGC 1386 (Sy 2)



A robust way to estimate the obscuration

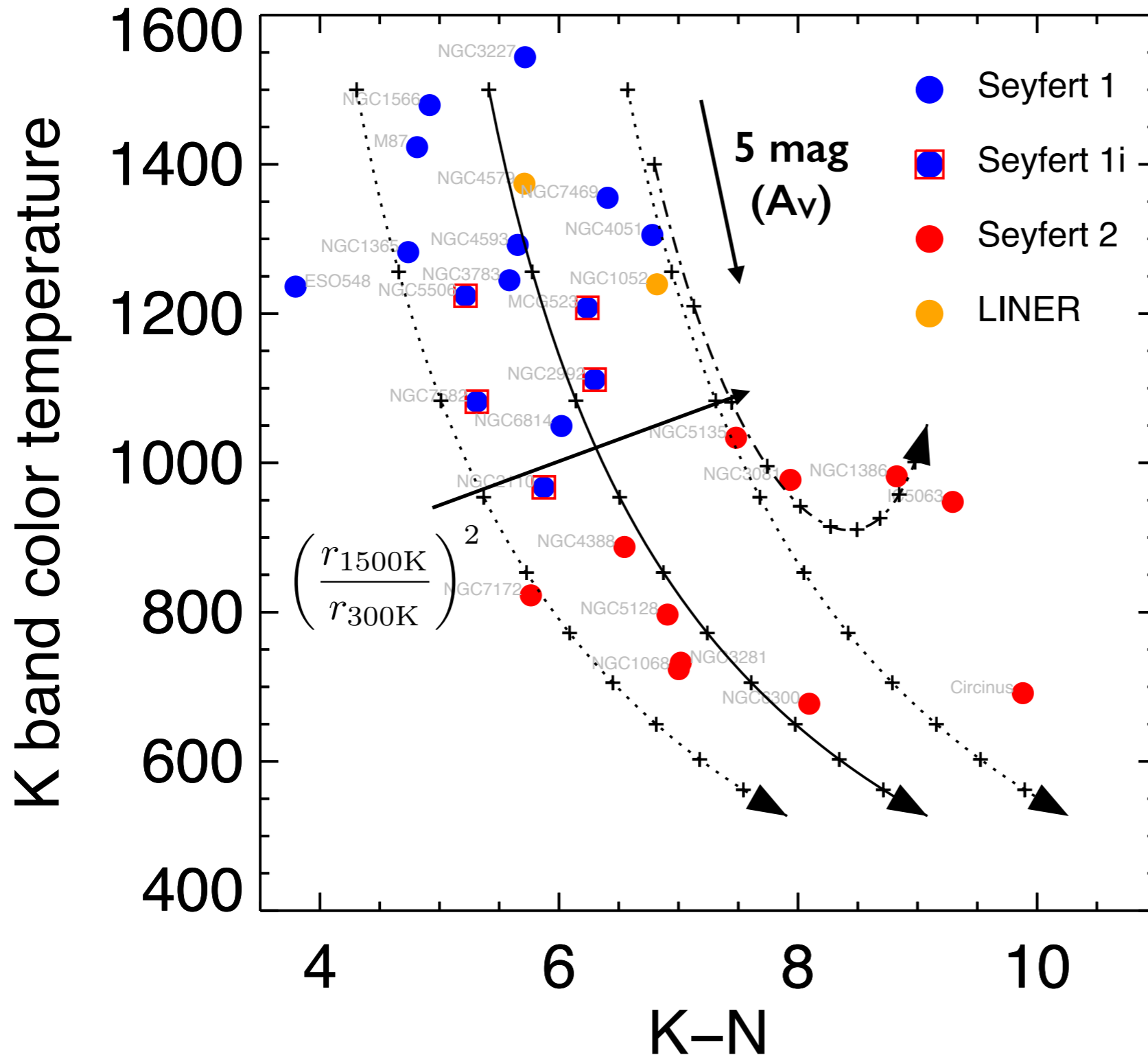


A robust way to estimate the obscuration



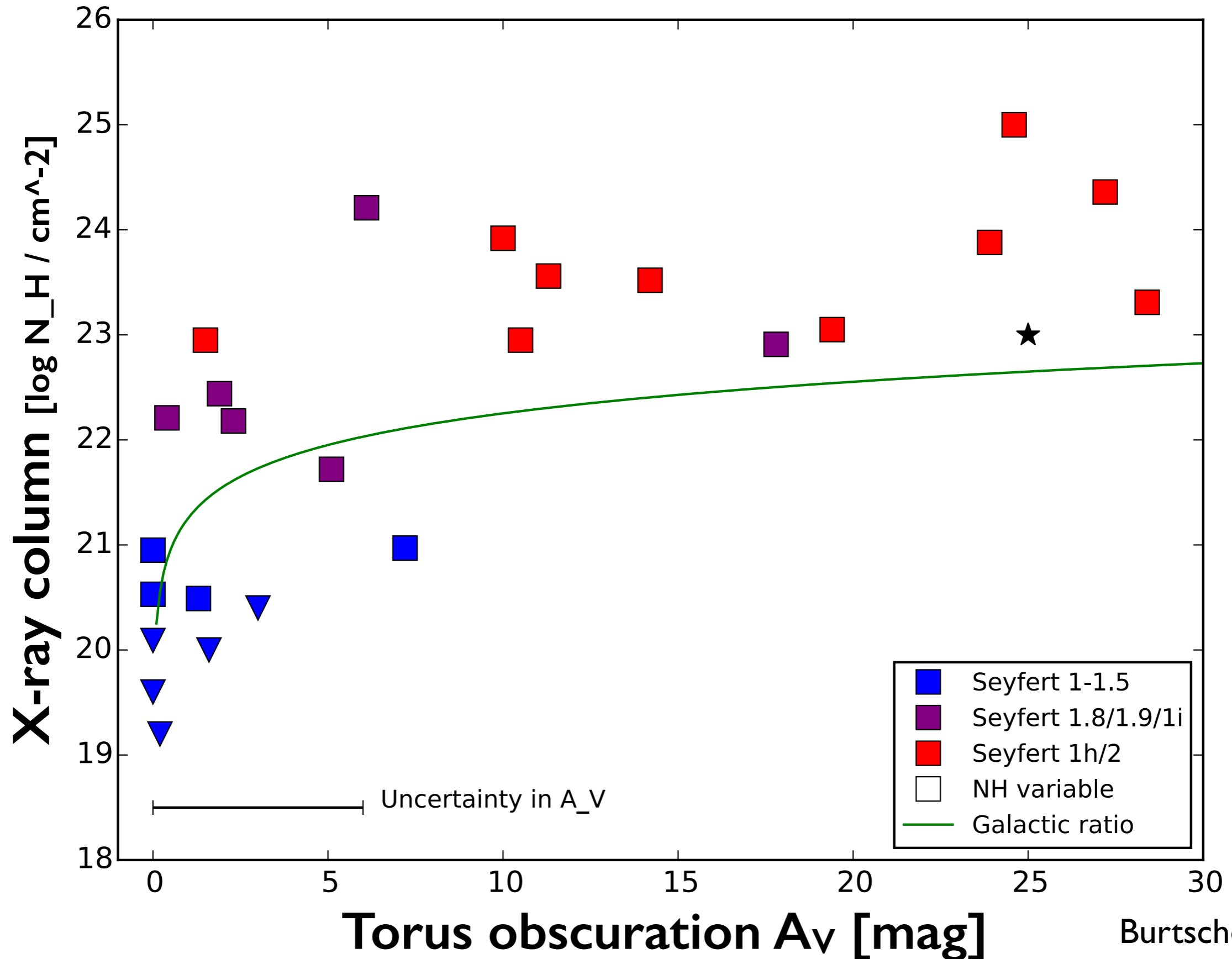
- **color temperature consistent with optical appearance** → obscuration

A robust way to estimate the obscuration

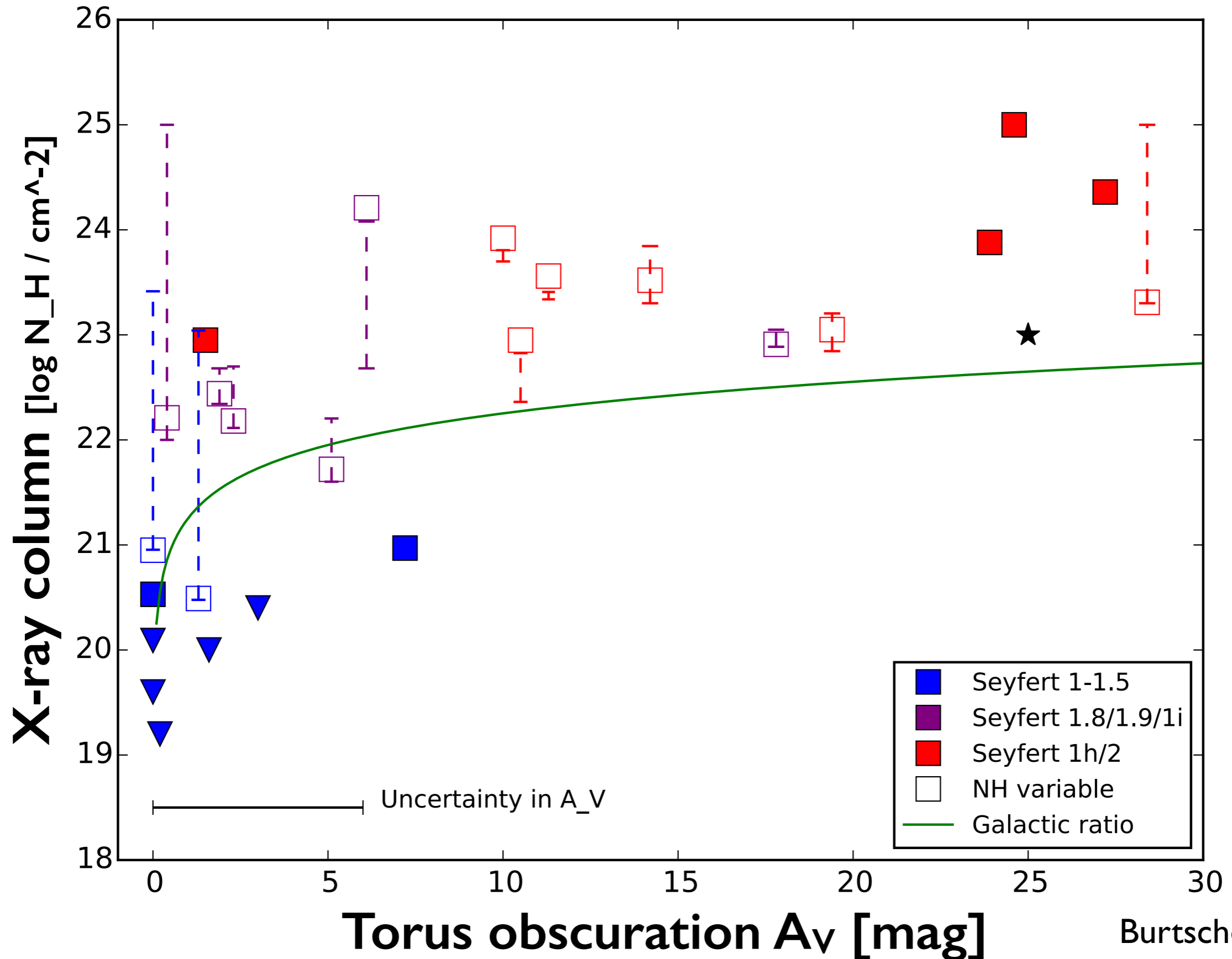


- **color temperature** consistent with optical appearance → obscuration
- **normalization** consistent with observed radii of hot/warm dust

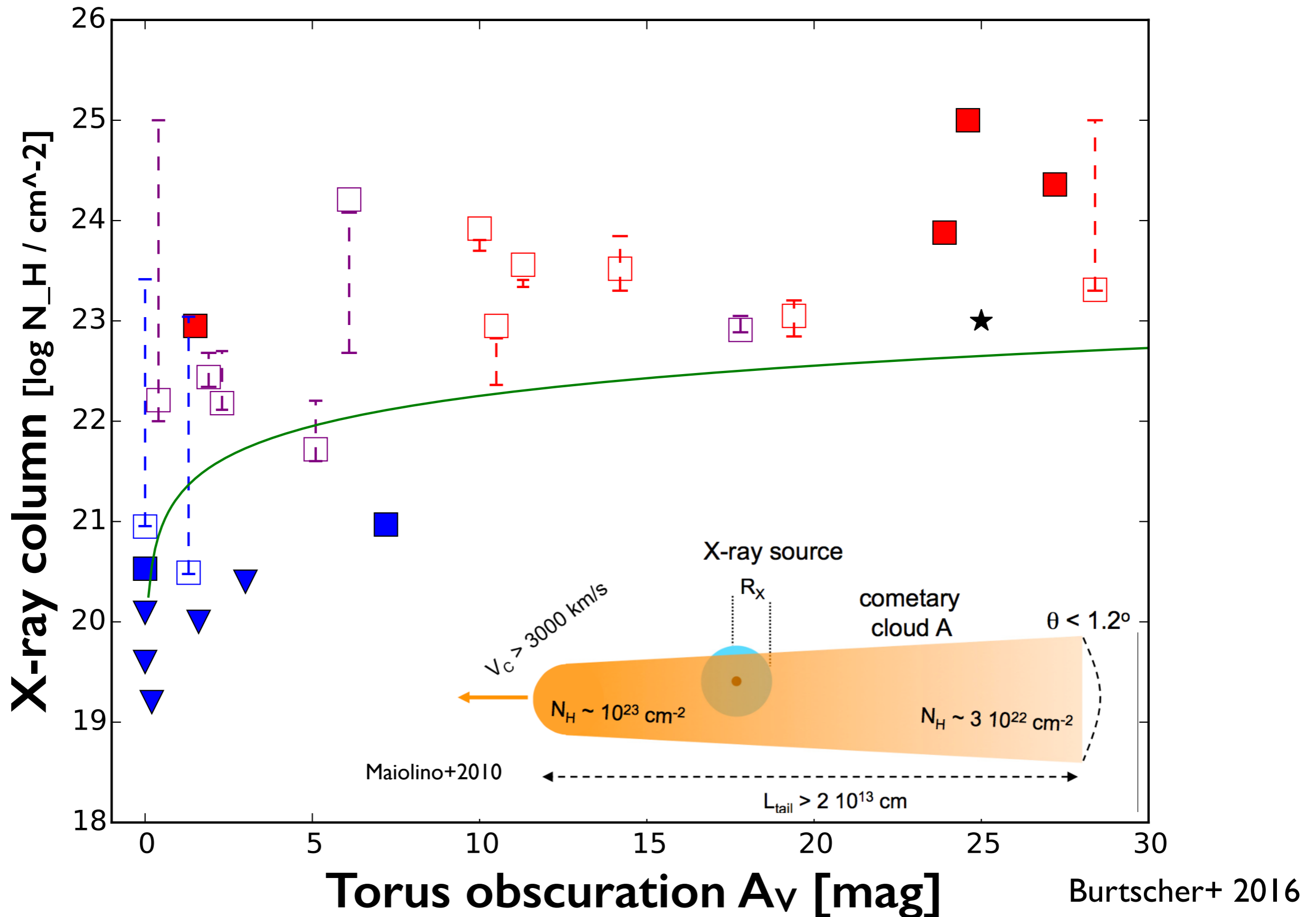
Obscuration: Torus vs. BLR



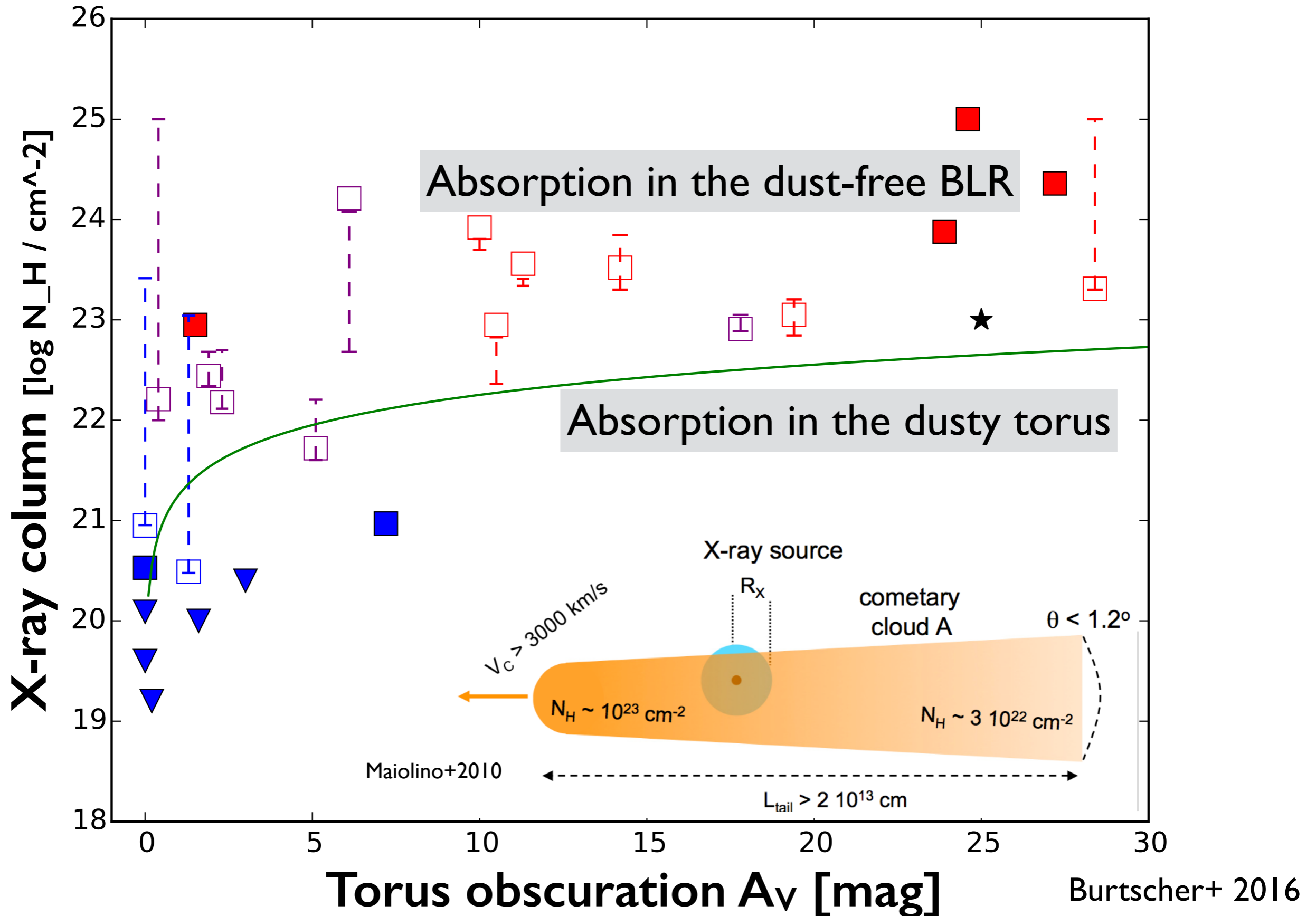
Obscuration: Torus vs. BLR



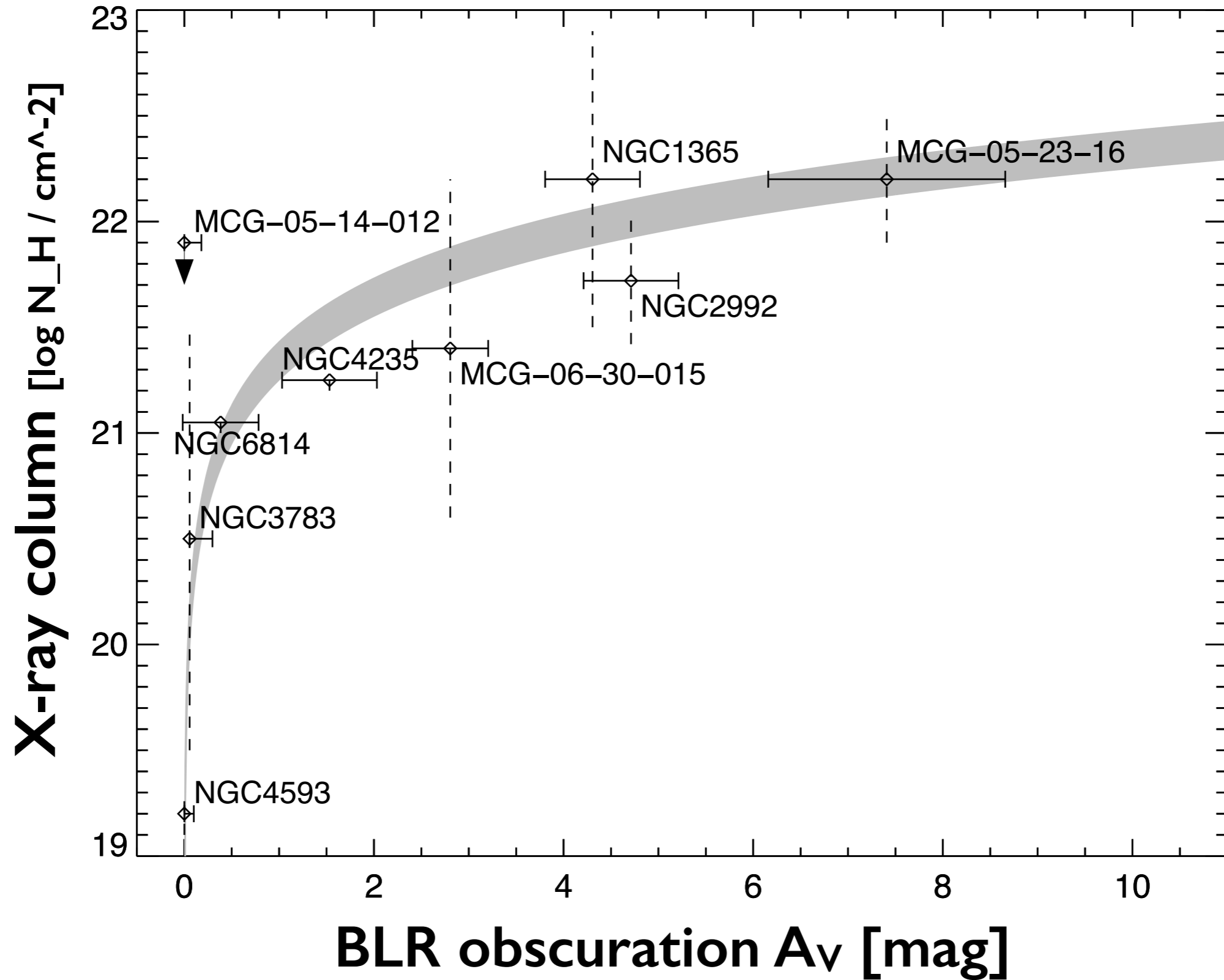
Obscuration: Torus vs. BLR



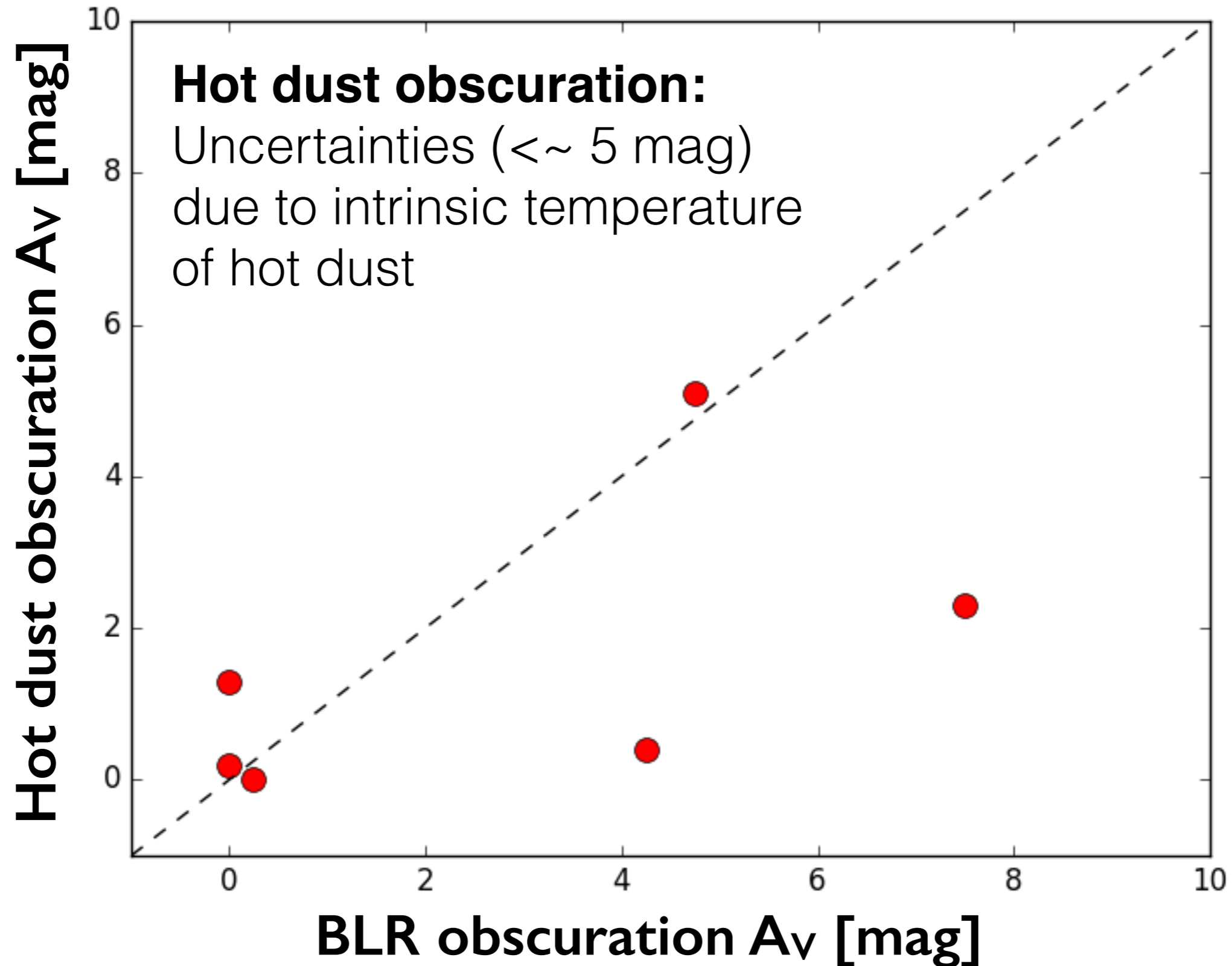
Obscuration: Torus vs. BLR



Obscuration: Torus vs. BLR



Obscuration: Torus vs. BLR



Outlook

for local, high-resolution AGN studies



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- **GRAVITY; MATISSE: 2nd generation VLTI instrument offering phases (imaging), higher resolution (*L* band, and *N*) and more efficiency (4 beams)**

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for local, high-resolution AGN studies

- GRAVITY; MATISSE: 2nd generation VLTI instrument offering phases (imaging), higher resolution (L band, and N) and more efficiency (4 beams)
- E-ELT/METIS

Outlook

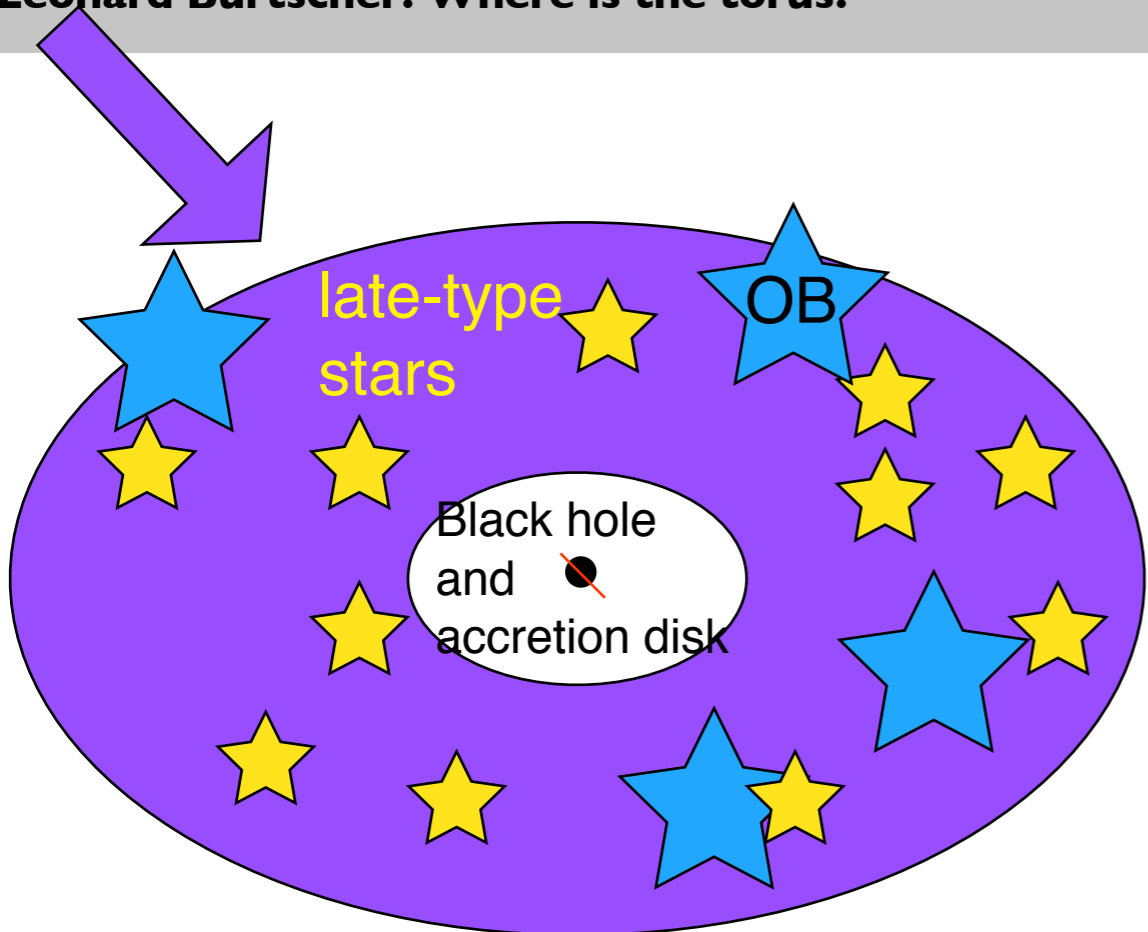
for local, high-resolution AGN studies

- GRAVITY; MATISSE: 2nd generation VLTI instrument offering phases (imaging), higher resolution (L band, and N) and more efficiency (4 beams)
- E-ELT/METIS
 - resolve large-scale „torus“ component found with MIDI: determine the kinematics of the wind launching region

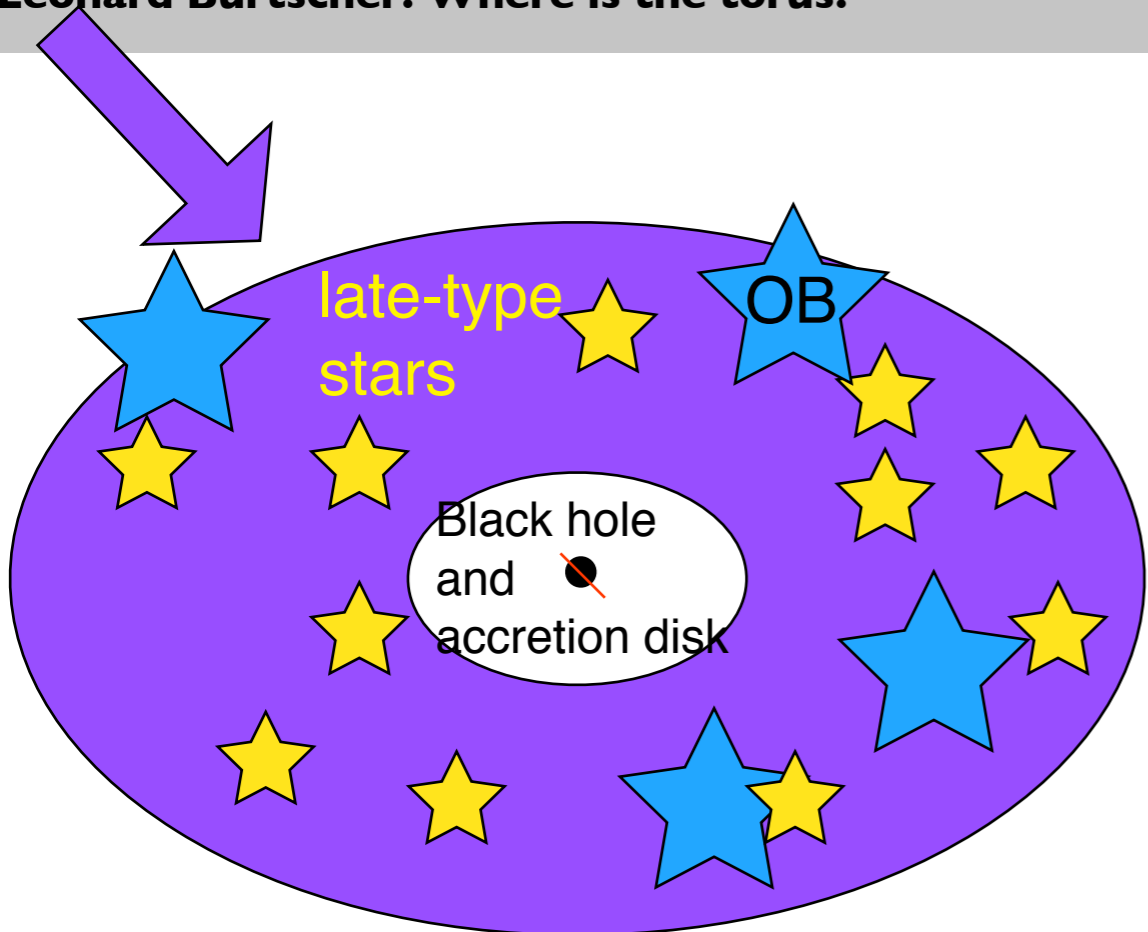
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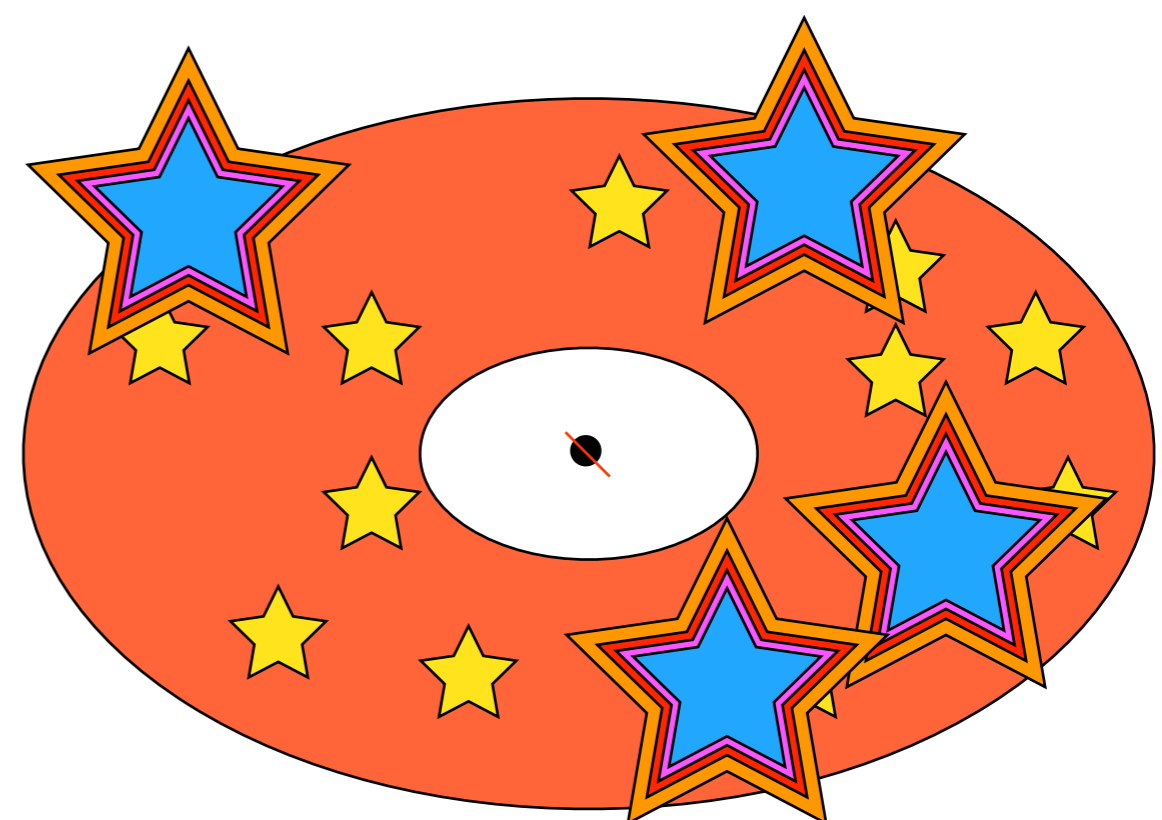
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- E-ELT/METIS
 - resolve large-scale „torus“ component found with MIDI: determine the kinematics of the wind launching region
 - resolve stellar populations very close to nearby AGNs



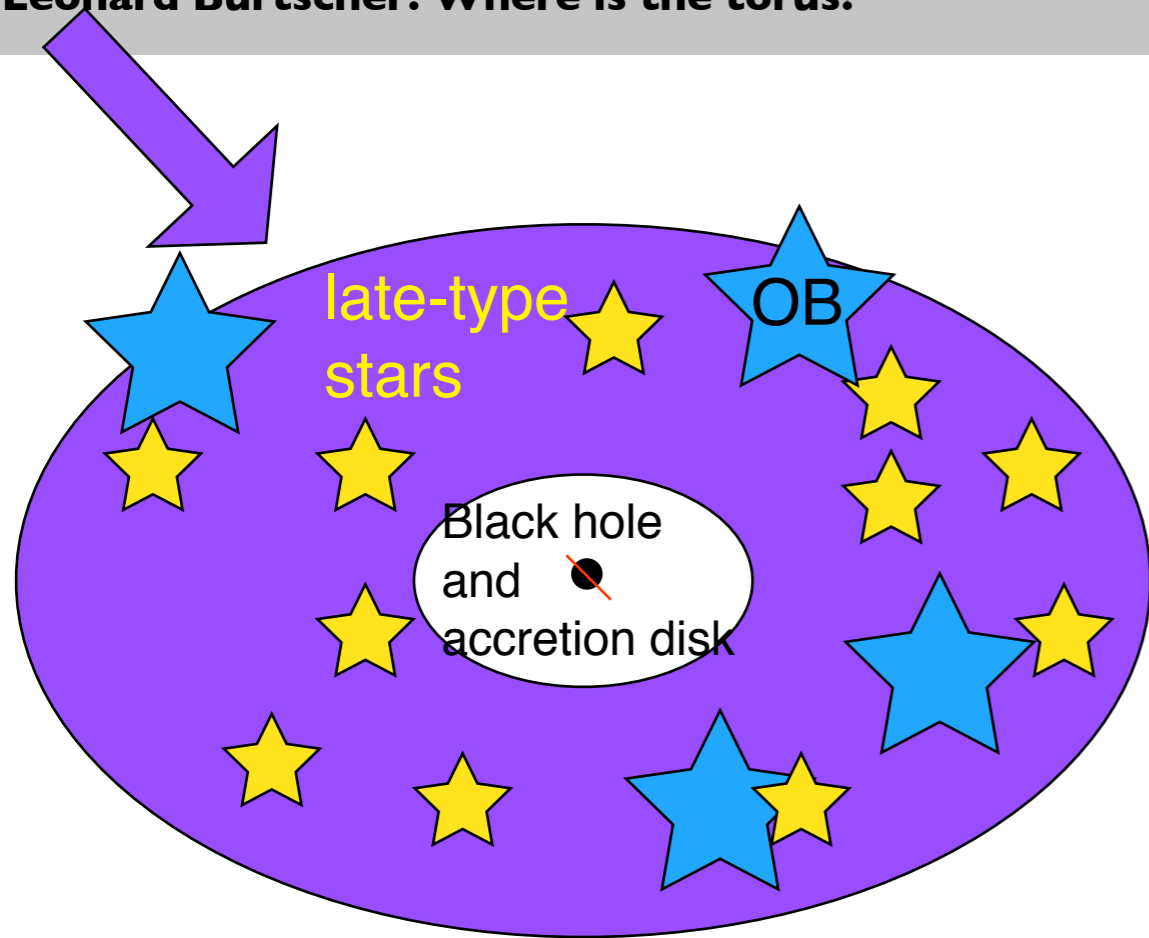
(1) inflow + starburst



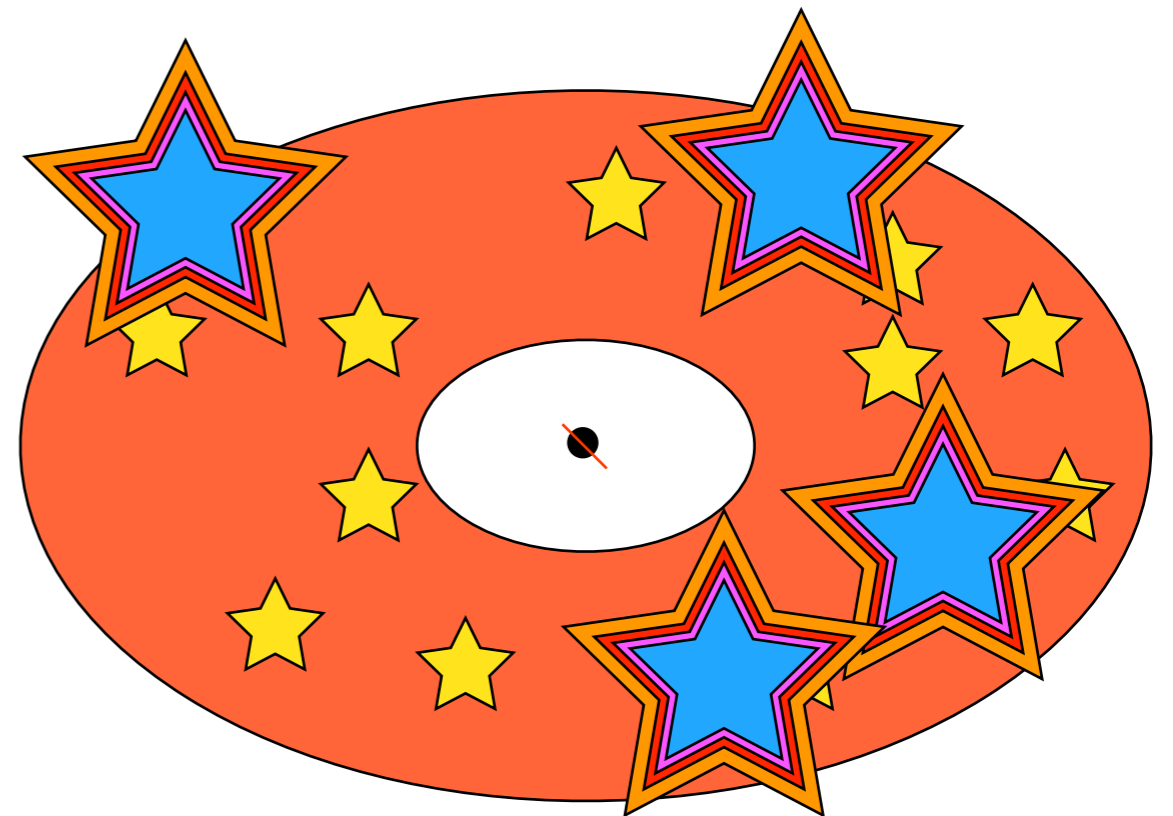
(1) inflow + starburst



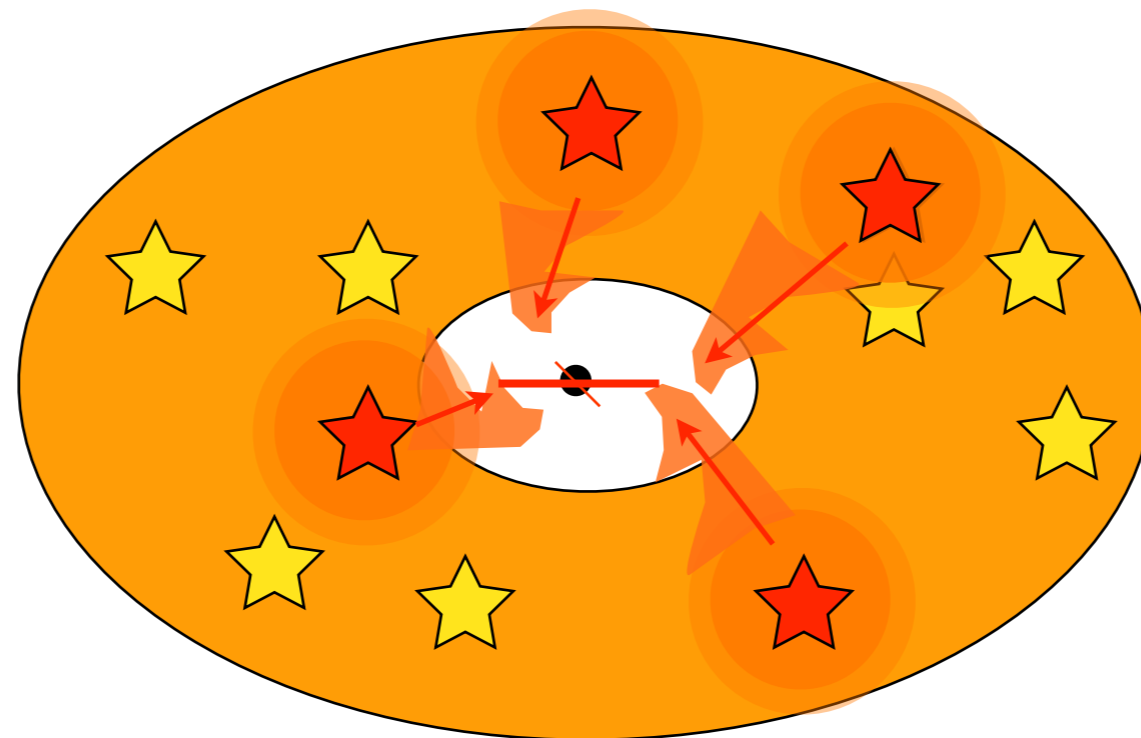
(2) supernovae + turbulence



(1) inflow + starburst

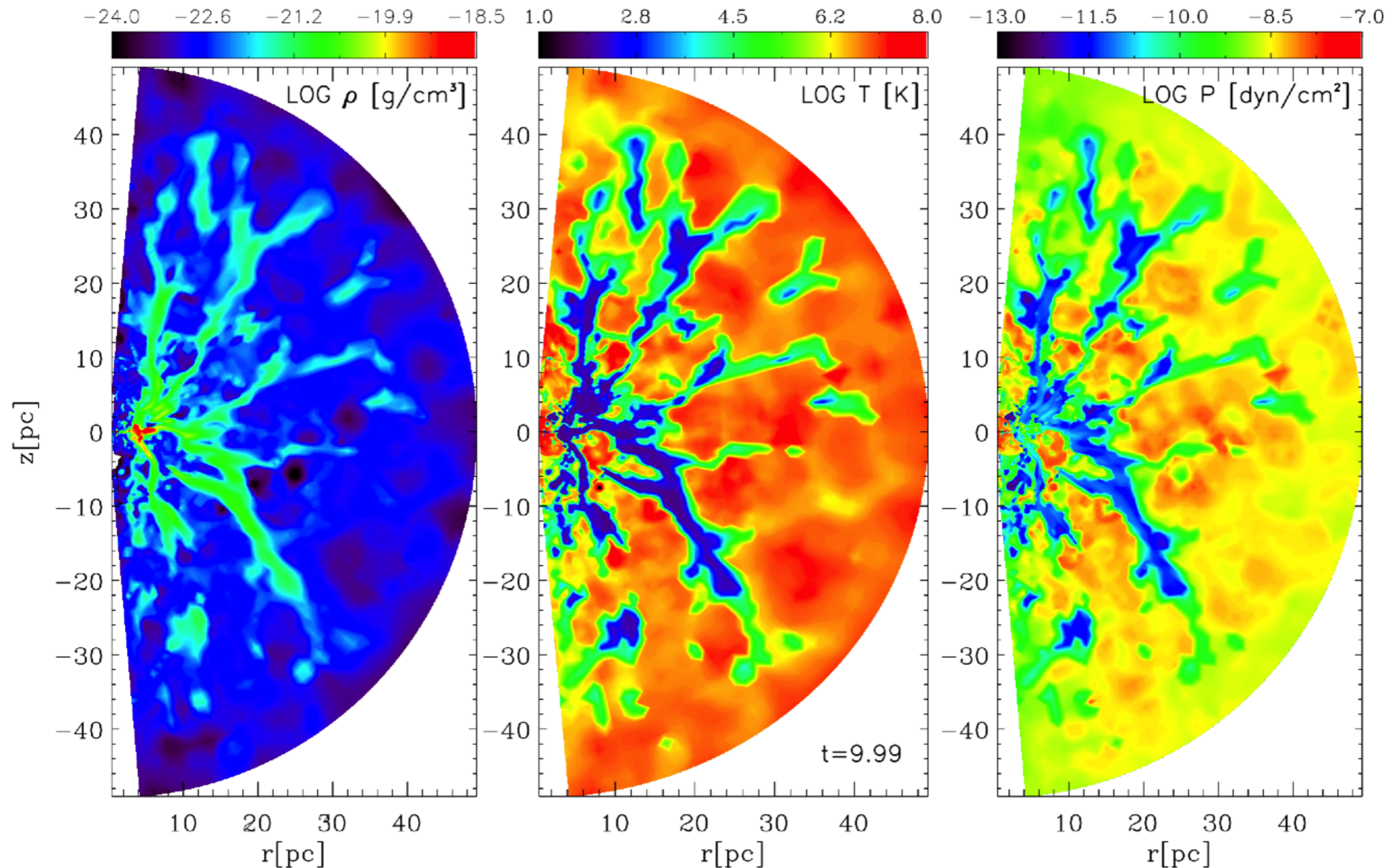


(2) supernovae + turbulence



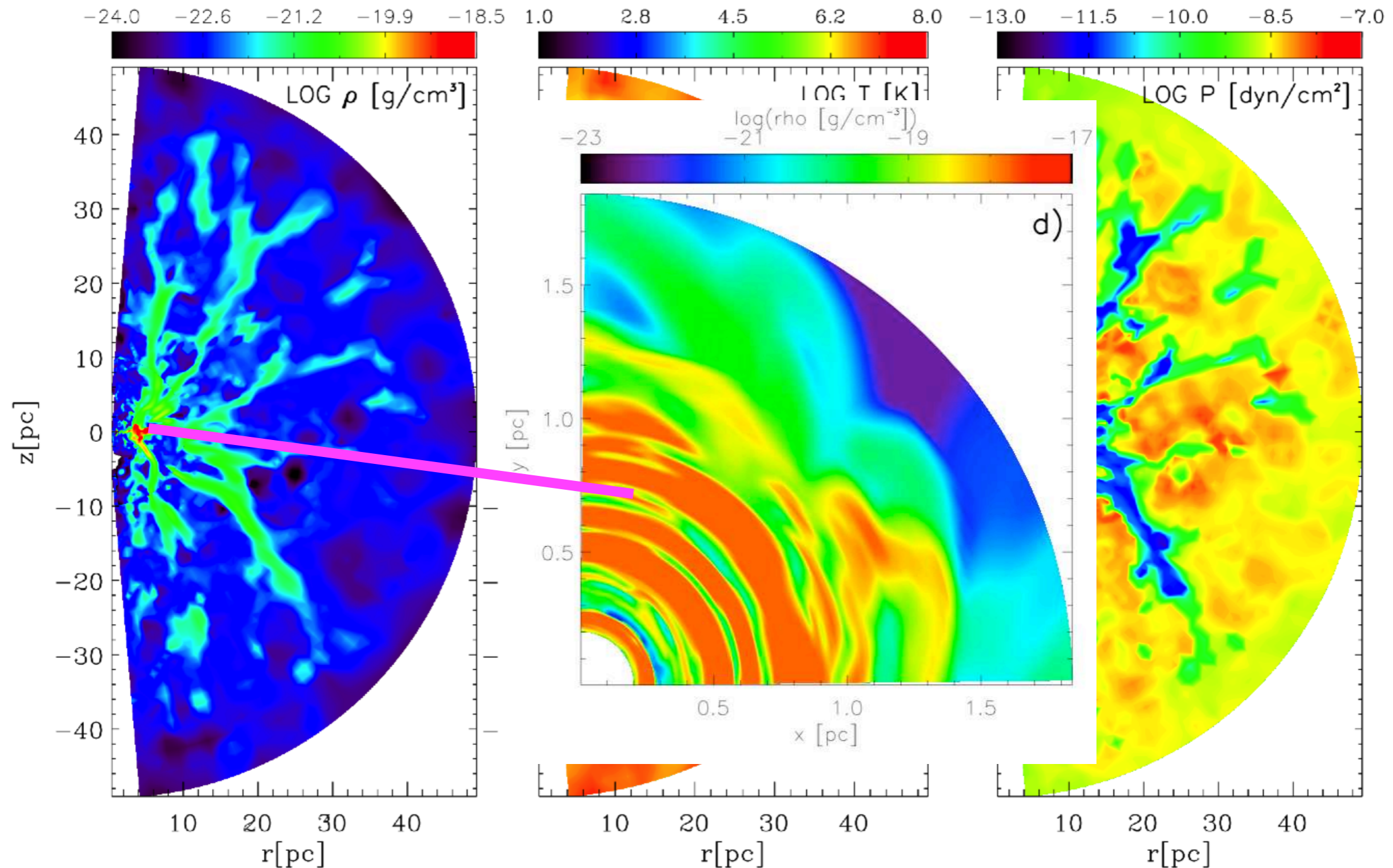
(3) AGB stars and stellar winds

Physical torus models

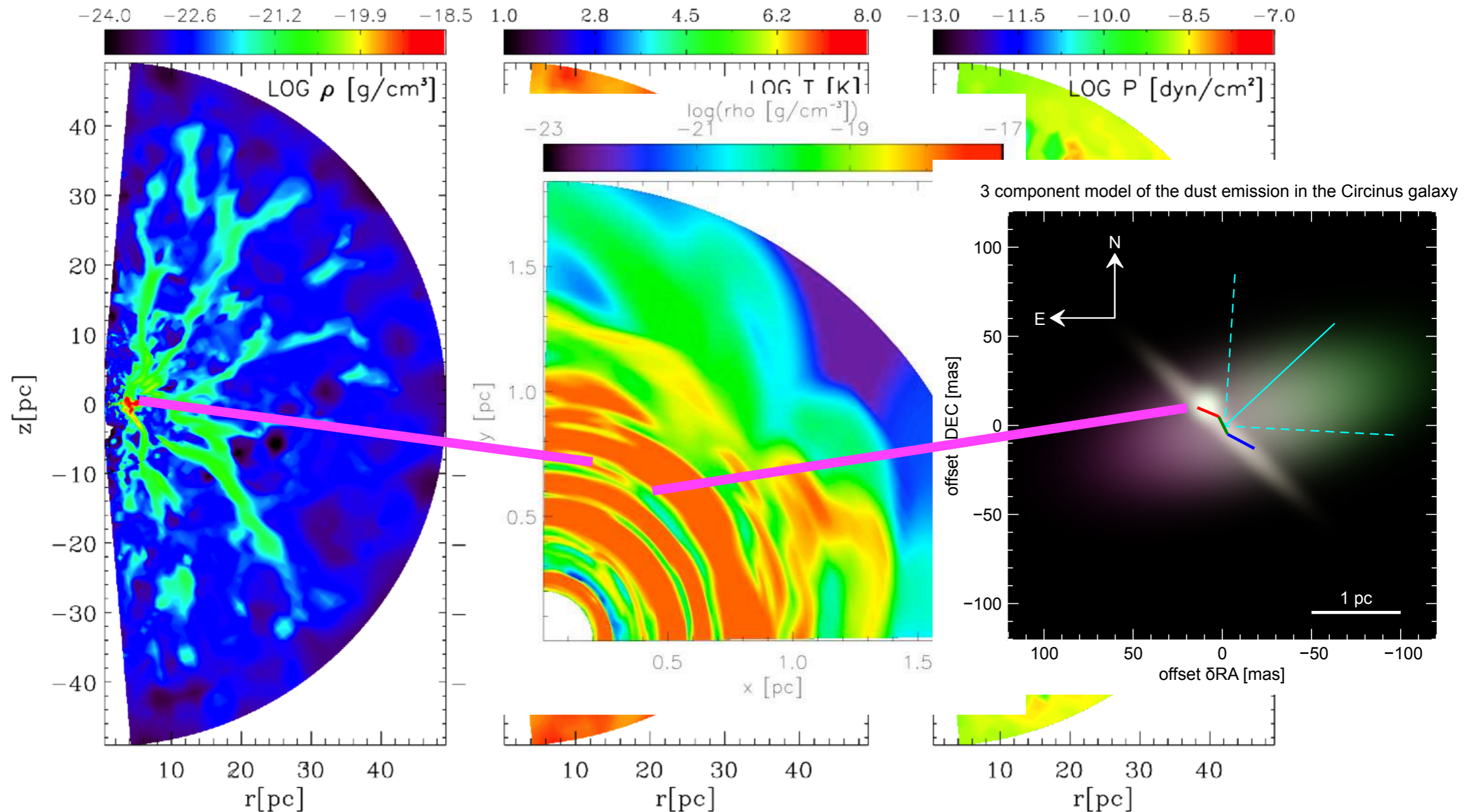


Schartmann+ 2009

Physical torus models



Physical torus models



Schartmann+ 2009