

Numerical Study of Balearic Meteotsunami Generation
and Propagation under Synthetic Gravity Wave Forcing
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Abstract

We use a high resolution nested ocean modelling system forced by synthetic atmospheric gravity waves to investigate Balearic meteotsunami generation, amplification and propagation properties. We determine how meteotsunami amplitude outside and inside of the Balearic port of Ciutadella depends on forcing gravity wave direction, speed and trajectory. We quantify the contributions of Mallorca shelves and Menorca Channel for different gravity wave forcing angles and speeds. The Channel is demonstrated to be the key build-up region determining meteotsunami amplitude in Ciutadella while northern and southern Mallorca shelves serve mostly as barotropic wave guides but do not significantly contribute to seiche amplitude in Ciutadella. This fact seriously reduces early-warning alert times in cases of locally generated pressure perturbations. We track meteotsunami propagation paths in the Menorca Channel for several forcing velocities and show that the Channel bathymetry serves as a focusing lens for meteotsunami waves whose paths are constrained by the forcing direction.

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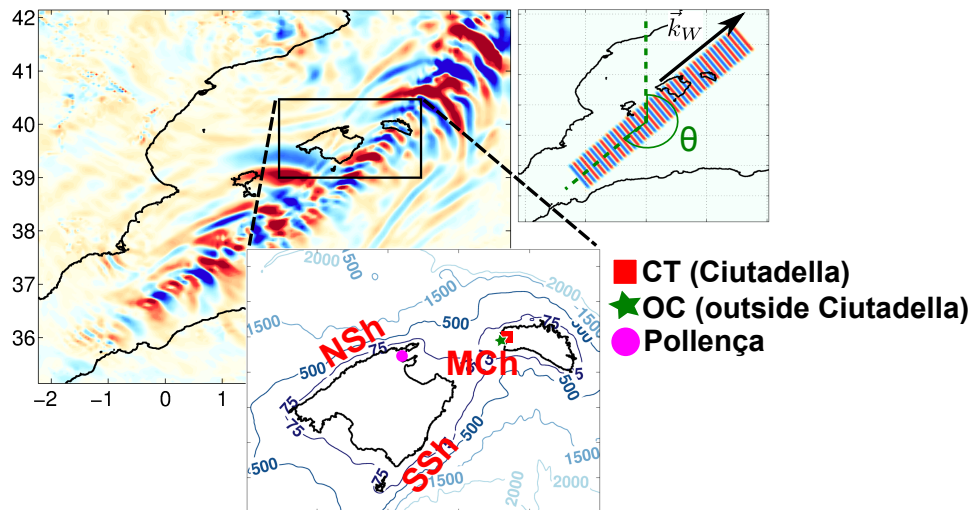


Figure 1: Background left: BRIFS operational WRF model high-pass filtered air pressure wave on the 11 June 2015 (color scale is truncated at ± 0.3 hPa). Black box, zoomed-up in the inset: ROMS parent domain with the Balearic bathymetry contours in meters with marked northern shelf (NSh), southern shelf (SSh) and Menorca Channel (MCh). Tiny red square in the inset shows ROMS child domain over the Ciutadella harbour. Green star shows the location of interest outside the Ciutadella harbour (denoted as OC in the rest of the paper). Purple circle shows location of the town of Pollença, Mallorca. Top right: synthetic air pressure snapshot for the $\theta = 230^\circ$ incident angle.

We show that faster meteotsunamis propagate over deeper ocean regions, as required by Proudman resonance. We estimate meteotsunami speed under sub- and supercritical forcing and derive a first order estimate of its magnitude. We show that meteotsunamis, generated by supercritical gravity waves, propagate with a velocity which is equal to an arithmetic mean of the forcing velocity and local barotropic ocean wave speed.

Keywords: meteotsunami modelling, meteotsunami propagation, Proudman resonance, meteotsunami amplification

1. Introduction

Meteotsunamis, i.e. tsunamis of meteorological origin [1, 2], are ocean waves in the tsunami frequency band generated over open ocean by the high frequency air pressure modulations of atmospheric gravity waves, convective pressure jumps or other kinds of atmospheric instabilities [1, 3]. Their amplification mechanisms include Proudman resonance (the matching of the air pressure disturbance velocity U and local ocean barotropic velocity $c_b = \sqrt{gH}$), topographic amplifications over continental shelves and harbour resonances as ocean waves enter narrow bays and inlets, oscillating at frequencies close to the resonant frequencies of these partially enclosed basins. These processes have been thoroughly explained elsewhere, e.g. [1, 4, 5]. Meteotsunamis have been observed all over the world oceans and their destructive port oscillations can also be found in the Mediterranean, for instance in the Balearic port of Ciutadella (see red square in the inset to Figure 1) [3, 6, 7, 8, 9, 10]. Here they are known as 'rissagas', but other examples can be found along the Sicilian and Croatian Adriatic coasts [11]. Meteotsunamis can even occur sequentially along the trajectory of the same synoptic system [4, 12].

Meteotsunami research efforts have been growing in the past decades and our understanding of the underlying processes has substantially improved. Nevertheless open issues remain. Even though Proudman resonance is known to play an important role in meteotsunami-related atmosphere-ocean energy transfer [13, 14, 15, 16], we still lack sufficient understanding of how a meteotsunami amplification is influenced by the resonant atmosphere-ocean interactions or by the local bathymetry, as was noted recently in [2]. In this paper we revisit this issue by presenting rissaga generation and propagation studies under synthetic forcing by atmospheric gravity waves, similar to the wave train modelled for June 11 2015, shown in Figure 1. Similar waves have previously been observed over the region [17, 18, 19]. Here we study the details of the respective contributions of the Mallorca shelves and Menorca Channel (see Figure 1 for location) to Ciutadella rissagas under gravity wave forcing conditions, and

31 the role of Proudman resonance in determining the meteotsunami propagation
32 paths over the shelves and the Menorca Channel. We conclude with an anal-
33 ysis of meteotsunami propagation speeds in the Menorca Channel during sub-
34 and supercritical ($U > c_{b0}$) forcing regimes, where $c_{b0} \approx 27 - 28 \text{ m s}^{-1}$ is the
35 Menorca Channel barotropic speed. The paper is organized as follows: Section
36 2.1 provides a brief description of the ocean modelling system. Section 2.2 de-
37 scribes synthetic atmospheric forcing used in the presented sensitivity analysis.
38 Rissaga sensitivity on forcing speed and direction is described in Section 3.1,
39 respective roles of the shelves and the Channel are presented in Section 3.2 and
40 propagation speed analysis is contained in Section 3.3. This is followed by the
41 conclusions in Section 4.

42 **2. Modelling Setup**

43 *2.1. ROMS Model Configuration*

44 The ocean modelling system used in this study is similar to that presented
45 in [3], which is also the oceanic component of the Balearic Islands Rissaga
46 Forecasting System (BRIFS) run operationally on a daily basis at SOCIB. It
47 is based on a double grid configuration of the ROMS model [20] with a 10-m
48 resolution grid around Ciutadella Inlet (hereafter child model) nested in a 1-km
49 resolution grid encompassing Mallorca and Menorca Islands (hereafter parent
50 model).

51 Both modeling domains are depicted in Figure 1. ROMS is a 3D free-surface,
52 split-explicit primitive equation model with Boussinesq and hydrostatic approx-
53 imations. Due to the 2-dimensional nature of the processes at play, the model
54 uses homogeneous temperature and salinity as initial conditions. A quadratic
55 parameterization is used for bottom drag. At the open lateral boundaries, Chap-
56 man and Flather conditions are used for free surface and 2D momentum respec-
57 tively. While no external boundary input is used for the larger domain, the
58 free surface and vertically integrated velocities of the parent model provide the
59 lateral boundary conditions for the child model with a 2-minute temporal res-

60 olution. BRIFS features a full WRF atmospheric model component to provide
 61 realistic high resolution atmospheric forcing for ROMS (see Figure 1 for an
 62 illustration). Here, the simulations were forced by the 2-minute resolution syn-
 63 thetic atmospheric pressure fields, similar to realistic WRF outputs, to analyse
 64 in detail the impact of the gravity wave properties on ocean model response.

65 2.2. Synthetic Atmospheric Boundary Conditions

66 Pressure wave forcing fields, with parameters similar to existing observa-
 67 tions ([17, 18], and SOCIB June 11 2015 mean sea level pressure observa-
 68 tions, available online via SOCIB THREDDS server at [http://thredds.socib.es-](http://thredds.socib.es/thredds/catalog.html)
 69 [/thredds/catalog.html](http://thredds.socib.es/thredds/catalog.html)), were generated as follows. Each respective numerical
 70 experiment featured an atmospheric pressure wave travelling at a definite phase
 71 velocity U , incident angle θ and lateral arc width of 1° . Frequency ν_W was kept
 72 constant at 10^{-3} s^{-1} for all waves. This frequency implies a 16.7-minute pe-
 73 riod, corresponding to a Ciutadella harbour amplification factor of 2.6, see e.g.
 74 [5]. Note that the ground state eigenperiod of the Ciutadella harbour equals
 75 10.5 minutes with a harbour amplification factor of 9.0. This paper does not
 76 address the sensitivity of the rissaga to the forcing frequency, which scales the
 77 final harbour resonance, but focuses on the generation, amplification and prop-
 78 agation processes over the shelves before entering Ciutadella inlet. The angle
 79 θ is stated throughout the paper using the nautical notation: $\theta = 180^\circ$ - wave
 80 propagates from the south, $\theta = 270^\circ$ - wave propagates from the west. For each
 81 numerical experiment a separate pressure wave was generated for angles from
 82 $\theta = 180^\circ$ to 270° in steps of 10° and with phase velocities from $U = 21 \text{ m/s}$ to
 83 36 m/s in steps of 1 m/s . The atmospheric pressure wave is then generated with
 84 $p(\vec{r}, t) = p_0 \cdot R(t, t_R) \cdot \cos(\vec{k}_W \cdot \vec{r} - \omega_W t)$, where p_0 is the pressure wave amplitude
 85 (set to 3 hPa , matching past observed values and implying pressure change rates
 86 of about 0.7 hPa/min), $R(t, t_R)$ is a pressure amplitude ramp function which
 87 linearly rises from 0 to 1 as t grows from 0 to ramp time t_R (set to $t_R = 16$
 88 minutes). The quantity $\vec{k}_W = (2\pi R_\oplus / \lambda_W) [\cos \theta, \sin \theta]$ is the wave-vector, R_\oplus is
 89 the mean Earth radius, $\lambda_W = U / \nu_W$ is the pressure wave wavelength, $\vec{r} = [\lambda, \phi]$

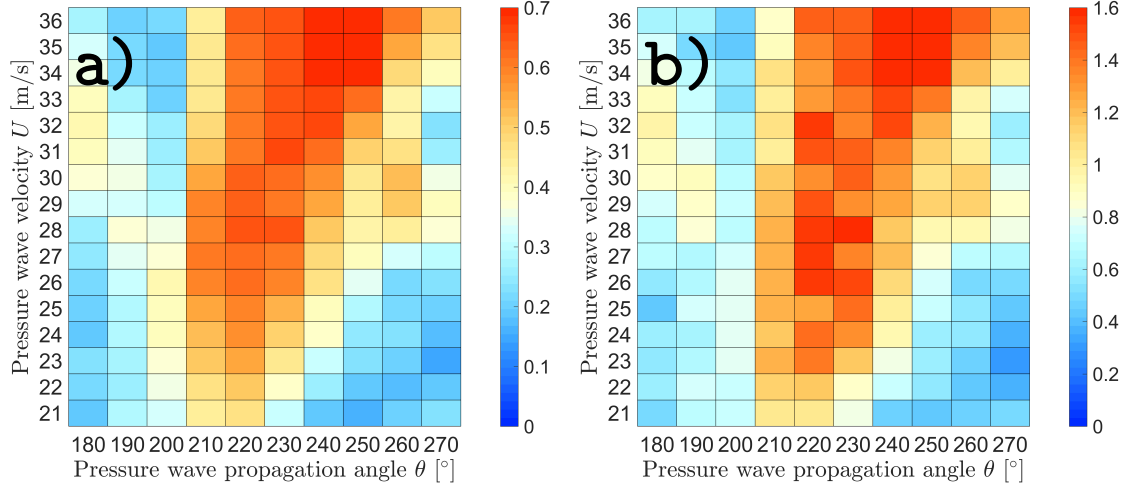


Figure 2: a) Dependence of the maximum generated SSH anomaly [m] at the OC point on the forcing gravity wave incident angle θ and its speed U . Colorbar depicts the elevation scale in meters. b) The same for the CT point inside Ciutadella harbour (see Figure 1 for location). Note the different color scales in both panels.

90 is the geo-referenced location on the Earth surface (λ and ϕ being longitude
 91 and latitude), $\omega_W = 2\pi\nu_W$ is the wave angular frequency and t is time. The
 92 pressure wave is then cropped in space and time to have the appropriate lateral
 93 width (0.5° to 1° arc degree) and duration (set to 6 hours - note that the gravity
 94 wave shown in Figure 1 endured for over 9 hours).

95 3. Results and Discussion

96 3.1. Rissaga Sensitivity to Pressure Wave Direction and Speed: the influence of 97 Proudman resonance.

98 To test the reliability of the modelling system's physics we first ran a separate
 99 24-hour simulation for each of the 160 forcing scenarios described in Section 2.2
 100 to study how the SSH anomaly outside and inside Ciutadella harbour depends
 101 on two key parameters: pressure wave propagation angle θ and velocity U .
 102 These results are presented in Figure 2. Note that they are consistent with

103 those from [6] even though our ocean model is forced by wave trains instead of
 104 a pressure step function as in [6] or [21]. Several features are present in Figure
 105 2. First, both matrices, outside (left panel in Figure 2) and inside (right panel
 106 in Figure 2) Ciutadella harbour, exhibit the absence of significant rissagas for
 107 angles below $\theta = 210^\circ$ and above $\theta = 260^\circ$. Second, there is a wide interval
 108 of velocities, from 23 m/s to 36 m/s, which generate substantial rissagas for
 109 forcing angles $210^\circ \leq \theta \leq 250^\circ$. Third, there are two areas with maximum SSH
 110 anomalies, one at $\theta = 230^\circ$ and $U = 28$ m/s, and another at $\theta = 240^\circ - 250^\circ$
 111 and $U = 34 - 36$ m/s. And finally, the SSH anomaly pattern at the CT point,
 112 extracted from the ROMS child domain and depicted in the right panel in
 113 Figure 2, exhibits the same overall features as the one at OC point, taken from
 114 the ROMS parent model and depicted in the left panel in Figure 2, although
 115 some additional modulations (i.e. intensification of the $(\theta = 230^\circ, 28$ m/s)
 116 maximum inside the harbour) are present.

117 In what follows we proceed to explain these features. To understand the
 118 absence of significant rissagas for angles below $\theta = 210^\circ$ and above $\theta = 260^\circ$,
 119 the left panel of Figure 3 is instrumental. The $\theta = 210^\circ$ and $\theta = 260^\circ$ incidence
 120 angles, plotted over the shallow water wave speed \sqrt{gH} , illustrate that the
 121 range of angles leading to significant rissagas in our simulations is linked with
 122 the orientation of Mallorca shelves and Menorca Channel. Within this cone
 123 of forcing directions, a large range of shallow water velocities is found along
 124 the propagating path of any given atmospheric disturbance towards Ciutadella.
 125 One of these velocities will match the forcing velocity giving rise to efficient
 126 transfer of energy from atmosphere to ocean through Proudman resonance.

127 The right panel of Figure 3 shows the Proudman length $L_P(\theta, U)$ dependence
 128 at the OC location on the forcing gravity wave incident angle θ and its speed
 129 U . Proudman length was computed in a similar fashion as suggested in [22]. It
 130 is defined as the total accumulated length (in kilometers) of the gravity wave
 131 path (propagating over the OC point at propagation angle θ) over which the
 132 Froude number $Fr = U/\sqrt{gH}$ is within 5 percent of its critical value, namely
 133 where $|Fr - 1| < 0.05$.

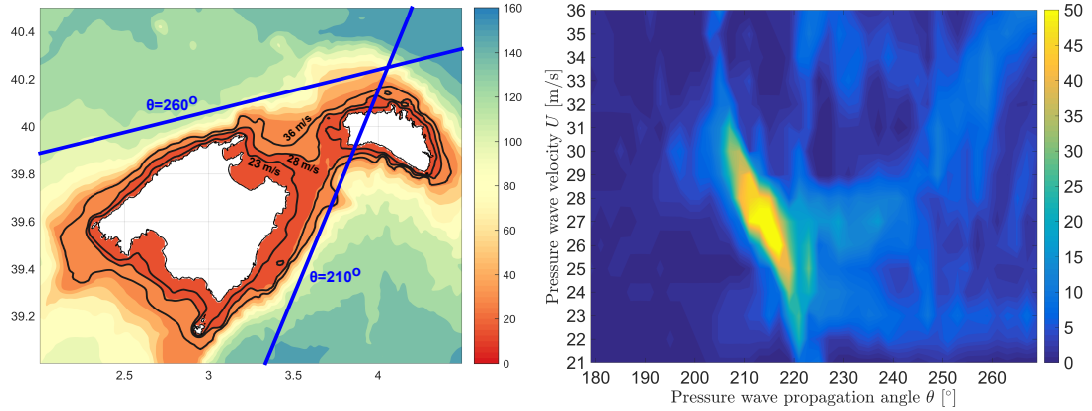


Figure 3: Left: Shallow water wave speed \sqrt{gH} [m/s] over the model domain. Black contour lines denote 23 m/s (innermost line), 28 m/s and 36 m/s (outermost line) speed isolines. The blue lines denote directions of gravity wave propagations at angles $\theta = 210^\circ$ and $\theta = 260^\circ$. Right: Proudman length $L_P(\theta, U)$ matrix [km] computed for the OC point (see Figure 1 for exact location).

134 The Proudman length matrix indicates that gravity waves, which propagate
 135 through the OC point in straight lines at angles $210^\circ - 220^\circ$ and speeds 26 – 28
 136 m/s, are the ones which travel to the largest extent over suitable ocean depths
 137 for the Proudman resonance to occur. Propagation through the OC point at
 138 these angles can only take place over the Southern Mallorcan shelf and the
 139 Menorca Channel, as can be seen from the left panel in Figure 3. The local SSH
 140 anomaly maximum in Figure 2, which occurs at $\theta = 220 - 230^\circ$ and $U = 26 - 28$
 141 m/s, therefore seems a direct consequence of the Proudman resonance over the
 142 Southern Mallorcan shelf and the Menorca Channel. Since it is computed along
 143 straight lines, the Proudman length criterion is however insufficient to explain
 144 the wider pattern of high values of the SSH anomalies at OC and CT points
 145 (see Figure 2) between angles $\theta = 210^\circ$ and $\theta = 260^\circ$. To explain this wide
 146 range one must take into account how the broader shape of Balearic bathymetry
 147 impacts the propagation of the barotropic waves. As will be shown below in
 148 more detail, Balearic bathymetry acts as a focusing lens for meteotsunami waves
 149 while the shelves act as the guides for coastal trapped edge waves. This is due

150 to the fact that free waves in inhomogeneous media propagate in accordance
151 with the Snell's law of refraction, which effectively states that travelling waves
152 deflect toward regions with lower propagation velocity. Consequently the waves,
153 generated by the forcing angles between $\theta = 210^\circ$ and $\theta = 260^\circ$, which would
154 otherwise perhaps have missed Ciutadella, will be guided along the isobaths
155 into the Channel and towards Ciutadella - this is also explicitly shown below
156 in Section 3.3. Outside the $\theta = 210^\circ - 260^\circ$ region, the ocean is too deep for
157 the Proudman resonance to occur since velocities of atmospheric disturbances
158 rarely surpass 40 m/s.

159 As for the SSH anomaly peak around $U = 34 - 36$ m/s and $240^\circ - 260^\circ$:
160 our simulations indicate that gravity waves at these speeds and angles generate
161 energetic coastally trapped edge waves and excite Channel eigenmodes with
162 periods of the order of a few hours (not shown in the paper). Edge waves
163 propagate northeastward along the Mallorca shelves, enter the Channel along
164 the isobaths and superimpose with the eigenmodes adding a final contribution
165 to the SSH anomalies visible at OC and CT points. Note that the SSH matrix
166 in [6], where a single pressure jump (modelled after the 2006 rissaga event) was
167 employed for synthetic experiments, indicates no such response maximum at
168 $U = 34 - 36$ m/s and $240^\circ - 260^\circ$ values. This further hints that the secondary
169 maximum, visible in our Figure 2 at $U = 34 - 36$ m/s and $240^\circ - 260^\circ$, might be
170 a consequence of forcing with gravity wave-trains of sufficient duration to excite
171 edge waves which have enough energy to contribute to Ciutadella rissagas.

172 Finally, we attribute the modulation in amplitude at point CT, as compared
173 to the point OC, to a modulation of the frequency of the first sea level oscillations
174 reaching the entrance of Ciutadella inlet. Preliminary calculations shows these
175 oscillations to be comprised between 13 and 17 minutes (not shown in the paper),
176 directly affecting the amplification inside the inlet [5]. No particular distribution
177 pattern was found over the range of forcing wave speeds and directions. These
178 aspects remains out of the scope of the present paper and further investigation
179 is needed to better understand these processes.

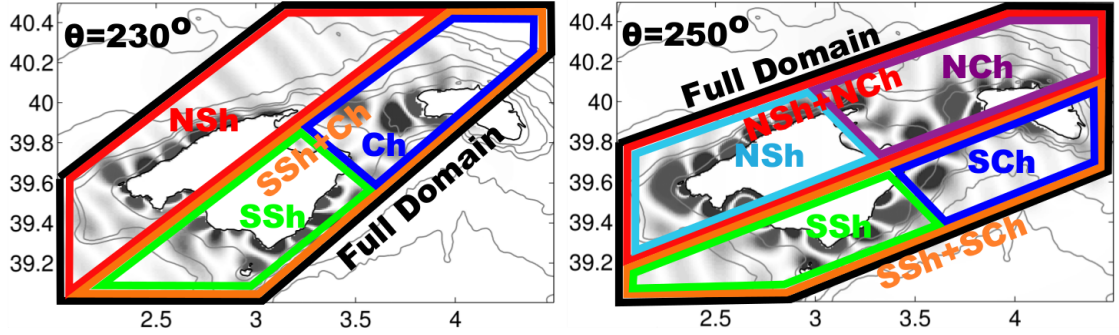


Figure 4: Gravity wave forcing scenarios. Each polygon depicts the domain of the passage of the synthetic gravity wave for $\theta = 230^\circ$ (left panel) and $\theta = 250^\circ$ (right panel). Polygon name abbreviations are as follows: Full Domain - full forcing domain, NSh - northern shelf, SSh - southern shelf, Ch - entire Channel, NCh - northern part of the Channel, SCh - southern part of the Channel.

180 *3.2. The role of the Mallorca shelves and Menorca Channel*

181 To quantify the shelves' and Channel's contributions to the meteotsunami
 182 amplification prior to the impact in Ciutadella harbour, a number of scenarios
 183 were run with the aim of exciting different subregions of the domain separately
 184 and exclusively. These synthetic forcing scenarios are depicted in Figure 4.
 185 During each forcing scenario the gravity wave was allowed to propagate only
 186 over the respective part of the domain. Gravity wave amplitudes rose linearly in
 187 time from zero to 3 hPa for 8 model timesteps and they were spatially smoothed
 188 at the relevant subdomain boundaries over the two outermost modeling cells to
 189 reduce sharp atmospheric pressure gradients. We chose angles and speeds which
 190 were shown to generate maximum rissaga in Ciutadella (see Figure 2), namely
 191 $\theta = 230^\circ$ forcing (Figure 4, left panel) with gravity wave speeds $U = 28$ m/s and
 192 $U = 35$ m/s, and $\theta = 250^\circ$ forcing (Figure 4, right panel) with speeds $U = 35$
 193 m/s. Note that the $\theta = 230^\circ$ forcing is aligned roughly with the isobaths of the
 194 southern shelf, while $\theta = 250^\circ$ forcing is aligned with the isobaths of the northern
 195 shelf, see Figure 3. Results of these simulations are shown in Table 3.2. For
 196 each scenario, the maximum sea level anomaly is also provided as the percentage
 197 of the maximum anomaly obtained when exciting the whole continental shelf

198 domain.

199 During the ($\theta = 230^\circ, 28 \text{ m/s}$) forcing scenario the Channel excitation is
200 the most important one - we can generate 93% of the full forcing (Full Domain)
201 response by forcing the Channel alone (Ch scenario). The difference is minimal
202 between Ch and SSh+Ch forcing scenarios, indicating that the contribution of
203 the southern Mallorca shelf is minimal in the build-up of the SSH anomaly.
204 Moreover, exciting only the shelf, either northern or southern, does not lead
205 to significant SSH anomalies in Ciutadella. The same qualitative results are
206 obtained with faster waves in the ($\theta = 230^\circ, 35 \text{ m/s}$) forcing scenario. The
207 Channel is still of paramount importance (97%). The northern shelf has greater
208 importance - we generate 28% of the full response by exciting the northern shelf
209 and 24% of the full response by exciting the southern one, but neither of them
210 gives rise to significant rissagas when excited separately. While the Proudman
211 resonance over the steeper northern shelf becomes more important for faster
212 gravity waves, the Channel is still the key region where critical amplification
213 takes place. The impact of deeper waters off the northern shelf is also visible for
214 the ($\theta = 250^\circ, 35 \text{ m/s}$) forcing scenario, where the excitation over the northern
215 shelf and the northern part of the Channel (NSh+NCh) generates 81% of the
216 full response, while forcing over the northern part of the Channel alone (NCh)
217 generates 73% of the full response. The contribution of the northern shelf is
218 still low but relatively larger than in previous cases with $\theta = 230^\circ$ forcing.
219 Southern shelf and southern part of the Channel (SSh+SCh) generate about
220 31% of the full response while southern part of the Channel (SCh) generates
221 30% of the full response. The main conclusion of these experiments is that the
222 key region that must be atmospherically excited to produce a strong rissaga is
223 almost exclusively the Channel itself.

224 As mentioned already in Section 3.1, the shelves therefore serve merely as
225 wave guides for barotropic meteotsunami waves into the Channel but do not con-
226 tribute significantly to the magnitude of the final SSH oscillation in Ciutadella.
227 From a practical operational point of view this means that pressure anomalies,
228 generated very locally only over the Channel, may lead to substantial rissagas.

θ, U	Full Domain	SSh+Ch	SSh	Ch	NSh		
230°, 28 m/s	1.64 m	1.56 m (95%)	0.28 m (17%)	1.54 m (93%)	0.17 m (10%)		
230°, 35 m/s	1.56 m	1.56 m (100%)	0.37 m (24%)	1.51 m (97%)	0.44 m (28%)		
θ, U	Full Domain	SSh+SCh	SSh	SCh	NSh+NCh	NSh	NCh
250°, 35 m/s	1.83 m	0.56 m (31%)	0.40 m (22%)	0.54 m (30%)	1.49 m (81%)	0.33 m (18%)	1.34 m (73%)

Table 1: Maximum SSH anomaly in Ciutadella during different forcing scenarios, shown in Figure 4. The elevations are in meters, percentages below each value are proportions to the maximum anomaly from Full Domain forcing scenario.

229 Such was, for example, the event on the August 18th 2014: a passage of
230 a single convective nucleus generated a very local pressure oscillation of 2.5
231 hPa in Ciutadella (light blue line in Figure 5) which was entirely absent in
232 the measurements in Pollença and other points of the island of Mallorca, as
233 shown in Figure 5, but nevertheless generated a peak-to-peak 1.45 m rissaga in
234 Ciutadella (not shown in this paper, see [23]). As shown by thermal IR satellite
235 image analysis in [23], the nucleus appeared in southern Mallorca and travelled
236 across Mallorca and the Menorca Channel to reach Ciutadella at the time of the
237 pressure oscillation which is also the time of the main sea level oscillation (not
238 shown). With the available data there is no way of proving that the convective
239 nucleus was generating the same kind of pressure oscillation along all its track
240 but the constant appearance (size, shape and intensity) from the satellite images
241 suggests that no important changes were occurring. Pressure oscillation linked
242 to a highly localized convective nucleus has induced a significant marine response
243 as the nucleus crossed the Menorca Channel at $\theta = 230 - 240^\circ$ direction with
244 $U = 27 - 28$ m/s speed.

245 This kind of evolution presents a serious challenge for the current Balearic
246 observation system since it dramatically reduces effective warning times and

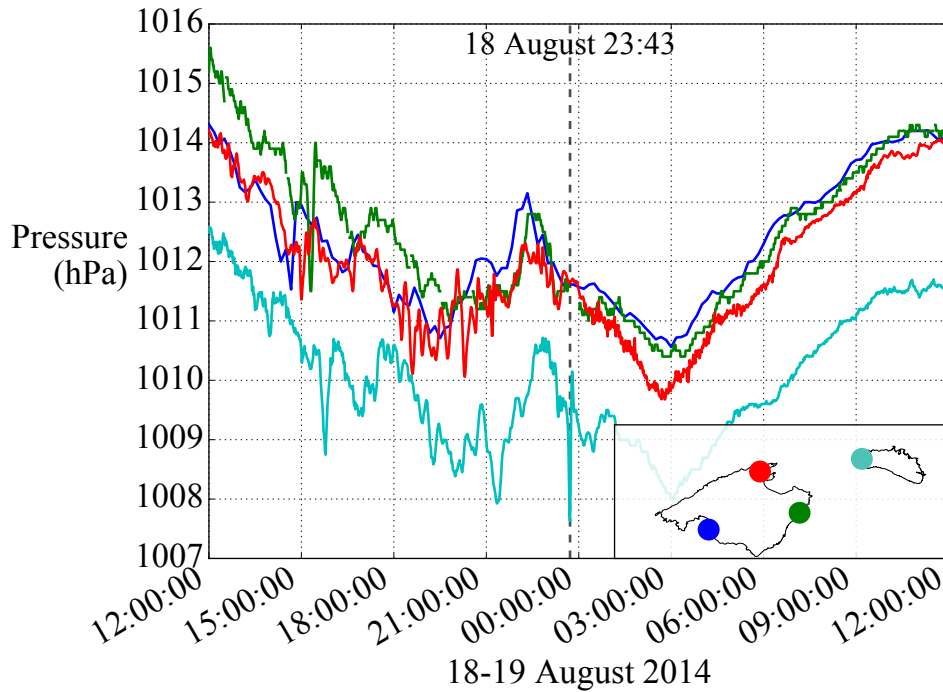


Figure 5: Air pressure [hPa] during a single and highly localized pressure oscillation occurring at Ciutadella (light blue line), generating a 1.45 m rissaga in the harbour. The oscillation was completely absent at Pollença (red line) and other pressure sensor locations on Mallorca (green and blue lines). See colored circles in the inset map for sensor locations.

247 indicates that, when too sparse, air pressure observations over Mallorca are not
 248 always an infallible criterion for *dismissing* the rissaga occurrence. Until a denser
 249 grid of (inland and offshore) air pressure stations is facilitated, the atmospheric
 250 model capability of timely forecasting the appearance of an atmospheric distur-
 251 bance, however difficult this may be, could be of importance to anticipating a
 252 meteotsunami arrival.

253 *3.3. Propagation Paths and Speeds in the Menorca Channel: Forced and Free*
 254 *Waves.*

255 To further understand the interplay between the atmospheric forcing and
 256 ocean bathymetry, we investigated how meteotsunamis propagate over the Menorca

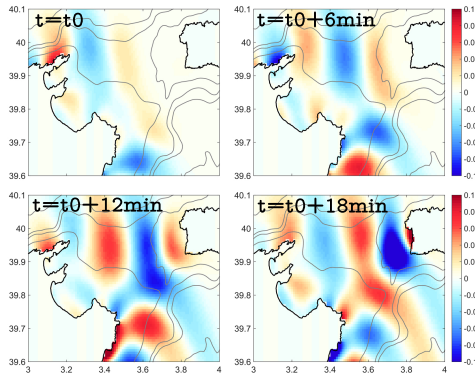


Figure 6: Meteotsunami focusing on Menorca Channel bathymetry. Panels show snapshots of SSH anomaly field [m] as meteotsunamis, arriving from northern and southern shelf, converge over the Channel towards Ciutadella. The colorscale has been limited to ± 0.1 m to enhance visibility. Plots depict the $U = 28$ m/s gravity wave forcing scenario.

257 Channel at forcing velocities $U = 22, 26, 28, 30, 34$ m/s and gravity wave inci-
 258 dent angle $\theta = 230^\circ$. Figure 6 shows that the Channel bathymetry acts as a
 259 focusing lens for the two meteotsunami waves coming from the northern and
 260 southern shelf. Waves from the southern shelf propagate along its general di-
 261 rection and traverse the Channel straight towards Ciutadella. Waves from the
 262 northern shelf are on the other hand deflected from the along-shelf direction by
 263 the shallow Channel bathymetry, turning them towards Ciutadella. Figure 6
 264 depicts four SSH anomaly snapshots where focusing and amplification of northern
 265 and southern meteotsunami branches are visible. Meteotsunami tracks converge
 266 along the Channel isobaths and merge prior to entering the Ciutadella harbour.

267 Proudman resonance influence on the propagation path can be evaluated
 268 by tracking the maximum SSH anomaly trajectories from both (northern and
 269 southern) meteotsunami branches, shown in Figure 6, as they travel across the
 270 Channel. These tracks are shown in Figure 7. As expected, the resonance plays a
 271 significant role: faster meteotsunamis traverse the Channel over a larger mean
 272 depth $\langle H \rangle_L$ which was estimated as a length-averaged line integral of ocean
 273 depth over the respective trajectory $\langle H \rangle_L = \Lambda^{-1} \int_L H(l) dl$, where $\Lambda = \int_L dl$ is
 274 the trajectory length and $H(l)$ is the model bathymetry value at location l along

275 the corresponding path L (colored curves in Figure 7), obtained at forcing speed
 276 U . A meteotsunami, generated by the $U = 22$ m/s gravity wave, traverses the
 277 Channel over the ocean region with a mean depth of 69.8 m. A meteotsunami,
 278 generated by the $U = 34$ m/s gravity wave, on the other hand propagates over
 279 the Channel at a mean depth of 76.4 m, see $\langle H \rangle_L$ in the legend of Figure 7.

280 Meteotsunami speeds were estimated from the maximum SSH anomaly tra-
 281 jectories across the Menorca Channel. These results are presented in Figure 8
 282 which depicts how the mean modelled meteotsunami speed (dark blue circles)
 283 varies with the atmospheric gravity wave speed and barotropic speed over its
 284 propagation path. As expected, in the subcritical forcing regime $U < c_{b0}$ the
 285 sea surface displacement $\eta(x, t)$ also propagates with the speed U , since it can
 286 be expressed as

$$287 \quad \eta(x, t) = -\frac{\delta p}{\rho g} \cdot \frac{x - Ut}{1 - (U/c_{b0})^2}, \quad (1)$$

288 where δp is the amplitude of the pressure perturbation, ρ is sea water density,
 289 g is acceleration due to gravity, and x and t are spatial and time coordinates,
 290 respectively [24].

291 If however the atmospheric gravity wave is supercritical ($U > c_{b0}$) then
 292 modelled meteotsunamis propagate faster than barotropic free waves but slower
 293 than the atmospheric gravity wave (see dark blue circles in the $U > c_{b0}$ region of
 294 Figure 8). A first order explanation of this behaviour is as follows. It has been
 295 shown [24] that when a supercritical storm, simulated as a Gaussian packet
 296 pressure perturbation of velocity $U > c_{b0}$, crosses a topographic obstacle, it
 297 generates a forced ocean wave, propagating with the storm velocity U , but also
 298 transmitted and reflected barotropic transient wake packets propagating with
 299 the local barotropic velocities. A supercritical air pressure wave train, crossing
 300 the Menorca Channel, will consequently radiate forced and transient barotropic
 301 waves, most notably a forced supercritical ocean wave, propagating as $\sin(x -$
 302 $Ut)$, x being distance along the direction of the gravity wave, and a continuous
 303 superposition of transient barotropic wake waves, generated continuously as the
 304 atmospheric gravity wave passes ocean regions of depths $H(x)$, each propagating

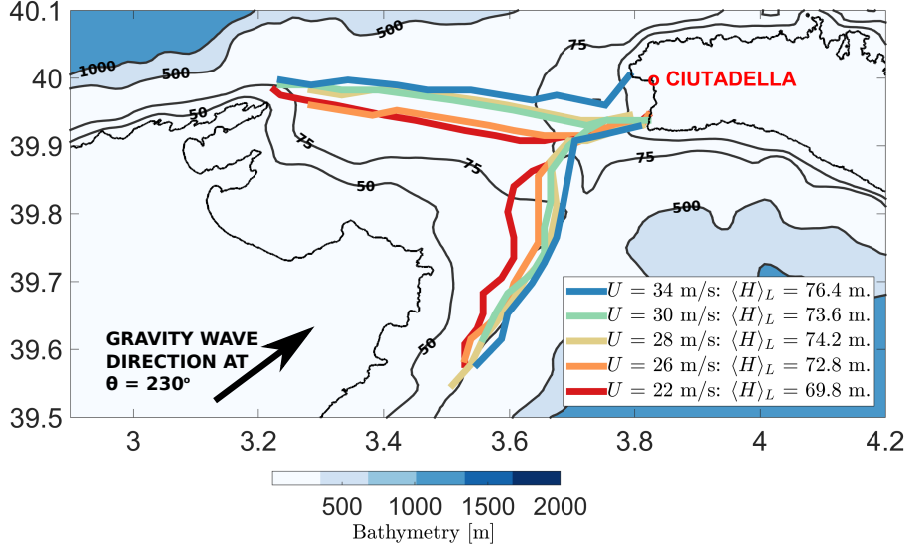


Figure 7: Maximum SSH anomaly trajectories across the Menorca Channel (colored curves) for different gravity wave speeds U at $\theta = 230^\circ$. Quantity $\langle H \rangle_L$ in the Figure legend denotes average depth along each respective trajectory L . Blue shades with dark grey contours denote the model bathymetry [in meters]. Lateral angular width of the meteotsunami forcing was 1° .

roughly along the forcing direction as $\sin(x - \sqrt{gH(x)}t)$.

To get an estimate of the meteotsunami velocity along the forcing direction, we approximate the superposition of wake waves with a single wave $\sin(x - \langle c_b \rangle_L t)$, travelling at the mean barotropic speed $\langle c_b \rangle_L$ along the path of the meteotsunami propagation. Mean barotropic speed is estimated as a distance-averaged line integral of \sqrt{gH} over the meteotsunami propagation contour L (these propagation and integration contours are shown as colored curves in Figure 7):

$$\langle c_b \rangle_L = \Lambda^{-1} \int_L \sqrt{gH(l)} dl. \quad (2)$$

The superposition of the forced and barotropic waves then simplifies to

$$\sin(x - Ut) + \sin(x - \langle c_b \rangle_L t) = \quad (3)$$

$$= 2 \cos \left[\frac{U - \langle c_b \rangle_L}{2} t \right] \sin \left[x - \frac{U + \langle c_b \rangle_L}{2} t \right].$$

316 The cosine term above denotes temporal modulation of the superposition am-
 317 plitude (and becomes a unit constant at the Proudman resonance where $U =$
 318 $\langle c_b \rangle_L$). The sine term above implies that the meteotsunami propagates with a
 319 velocity which is to a first order equal to the average of the atmospheric wave
 320 velocity and mean barotropic velocity along its path:

$$\tilde{c}_L = \frac{U + \langle c_b \rangle_L}{2}. \quad (4)$$

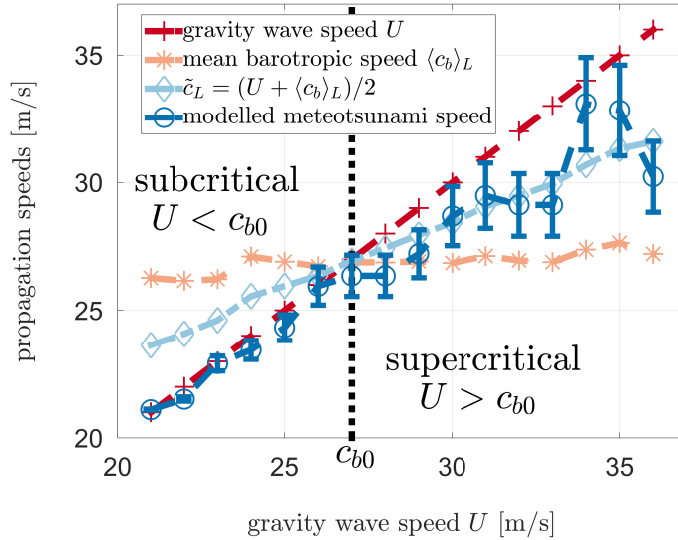


Figure 8: Mean propagation speeds along the maximum SSH anomaly trajectories for different gravity wave speeds U at $\theta = 230^\circ$. Red color: air pressure gravity wave speeds U . Beige color: mean barotropic speeds computed along the trajectory of meteotsunamis, generated by the gravity waves of velocities U . Light blue color: $\tilde{c}_L = (U + \langle c_b \rangle_L)/2$, the average of both speeds. Dark blue color: the modelled meteotsunami propagation speeds along the forcing direction, obtained by directly tracking the meteotsunamis in the simulation results. Vertical dashed black line marks Menorca Channel barotropic speed $c_{b0} \approx 27 - 28 \text{ m s}^{-1}$, delimiting sub- and supercritical forcing regimes.

321 Note that at the Proudman resonance \tilde{c}_L equals U as it should. Equation

322 (4) explains why the modelled meteotsunami speed values (dark blue circles
323 in Figure 8) are faster than barotropic (beige symbols in Figure 8) and slower
324 than forcing speeds (red symbols in Figure 8). Error bars in Figure 8 are our
325 estimates of the errors in tracked meteotsunami speeds, rising from 1 percent at
326 21 m s^{-1} to 5 percent at 36 m s^{-1} . This drop in accuracy is due to the fact that
327 faster waves tend to exhibit spatially spread front-like maxima patterns in the
328 Channel as they approach Menorca (not shown), making the possible location,
329 and hence the computed propagation speed of the meteotsunami somewhat
330 smeared in space. It is nevertheless clear that supercritical meteotsunamis travel
331 with velocities below the forcing speed U and above the local barotropic speed,
332 as indicated by the derived relationship (4), which arises out of the wave nature
333 of the forcing as the propagation speed of the superimposed forced waves and
334 free barotropic wake waves. (An interesting case of an actual occurrence of the
335 superposition of forced and free waves, leading to the great Adriatic surge of
336 1978, has been described in [25].)

337 4. Conclusions

338 An array of numerically generated atmospheric gravity wave trains was used
339 to investigate Balearic meteotsunami generation, amplification and propagation.
340 A sensitivity study of meteotsunami amplitude inside and outside Ciutadella
341 harbour was performed in a similar fashion to [6], indicating that a wide range
342 of gravity wave speeds ($23 - 36 \text{ m/s}$) and angles of propagation ($210 - 250^\circ$)
343 may lead to substantial meteotsunami generation. The above speed and angle
344 ranges were accounted for by the regional orientation and shape of the Balearic
345 bathymetry which determine where the Proudman resonance may occur.

346 To quantify shelf and Channel contributions to meteotsunami intensity, we
347 ran a set of simulations with forcings constrained to localized and isolated sub-
348 regions of the modelling domain. These simulations indicate that in order to
349 generate a significant meteotsunami, the Channel has to be atmospherically ex-
350 cited. Excitations over the northern and southern Mallorca shelves alone are

351 not capable of generating destructive meteotsunamis. The shelves serve mostly
 352 as waveguides, leading meteotsunami waves along the Mallorca coast and into
 353 the Channel where they get amplified by the atmospheric forcing over the Chan-
 354 nel, if such forcing is present. This means that very locally induced pressure
 355 oscillations, arising from atmospheric instabilities over the Menorca Channel
 356 alone, can lead to significant meteotsunamis as well. Such events have indeed
 357 occurred and one such observation is provided. Such atmospheric scenarios, rare
 358 as they may be, seriously reduce lead times for any early warning observation
 359 system in Mallorca. Numerical modelling capabilities to forecast (or at least
 360 indicate a significant possibility of) meteotsunami occurrences at least a day
 361 in advance could therefore be important for the prevention and mitigation of
 362 the consequences of such events. This provides further motivation for comple-
 363 menting existing observational networks and meteotsunami alert systems with
 364 numerical modelling systems.

365 If the gravity waves propagate over the Balearic region from the southeast,
 366 then the Channel bathymetry is shown to act as a focusing lens for the corre-
 367 sponding meteotsunami waves. These waves propagate along the Channel iso-
 368 baths and focus just prior to the impact with Ciutadella harbour. It is further
 369 demonstrated that maximum SSH anomalies, generated by faster atmospheric
 370 gravity waves, propagate over deeper ocean regions, as one would expect from
 371 the Proudman resonance.

372 We have further investigated meteotsunami propagation under sub- and su-
 373 percritical atmospheric forcing. Subcritical atmospheric gravity waves generate
 374 meteotsunamis which propagate over the Menorca Channel with velocity U .
 375 Meteotsunamis created by supercritical atmospheric gravity waves on the other
 376 hand propagate with velocity $\tilde{c}_L = (U + \langle c_b \rangle_L)/2$ which is, to the first order,
 377 an average of the gravity wave speed U and the barotropic speed of the wake
 378 wave, as shown in Figure 8, implying that, unlike subcritical meteotsunamis,
 379 these are not free barotropic but rather forced waves arising from an interplay of
 380 influences by forcing and the bathymetry. The relationship $\tilde{c}_L = (U + \langle c_b \rangle_L)/2$
 381 arises out of the wave nature of the atmospheric forcing and turns out to be

382 a consequence of the superposition between supercritically forced ocean waves
383 and barotropic ocean wake waves.

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387 *Author contributions.* C. T. and A. K. analysed the observational data upon
388 which synthetic gravity wave forcing was generated. B. M. maintained the
389 modelling environment and performed the simulations. M. L., B. M. and C. T.
390 created the figures. M. L. and B. M. prepared synthetic atmospheric boundary
391 conditions for ROMS, performed post-processing analyses and wrote the paper.
392 A. J. provided the analysis for the August 19th 2014 event. C.T., A.K., B.M.,
393 M.L. and J.T. devised the research plan. All authors contributed to the final
394 version of the manuscript. M. L. and B. M. contributed equally to this research.

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