



# Hitting the diffraction limit: first results of the AGPM-VORTEX project

Olivier Absil

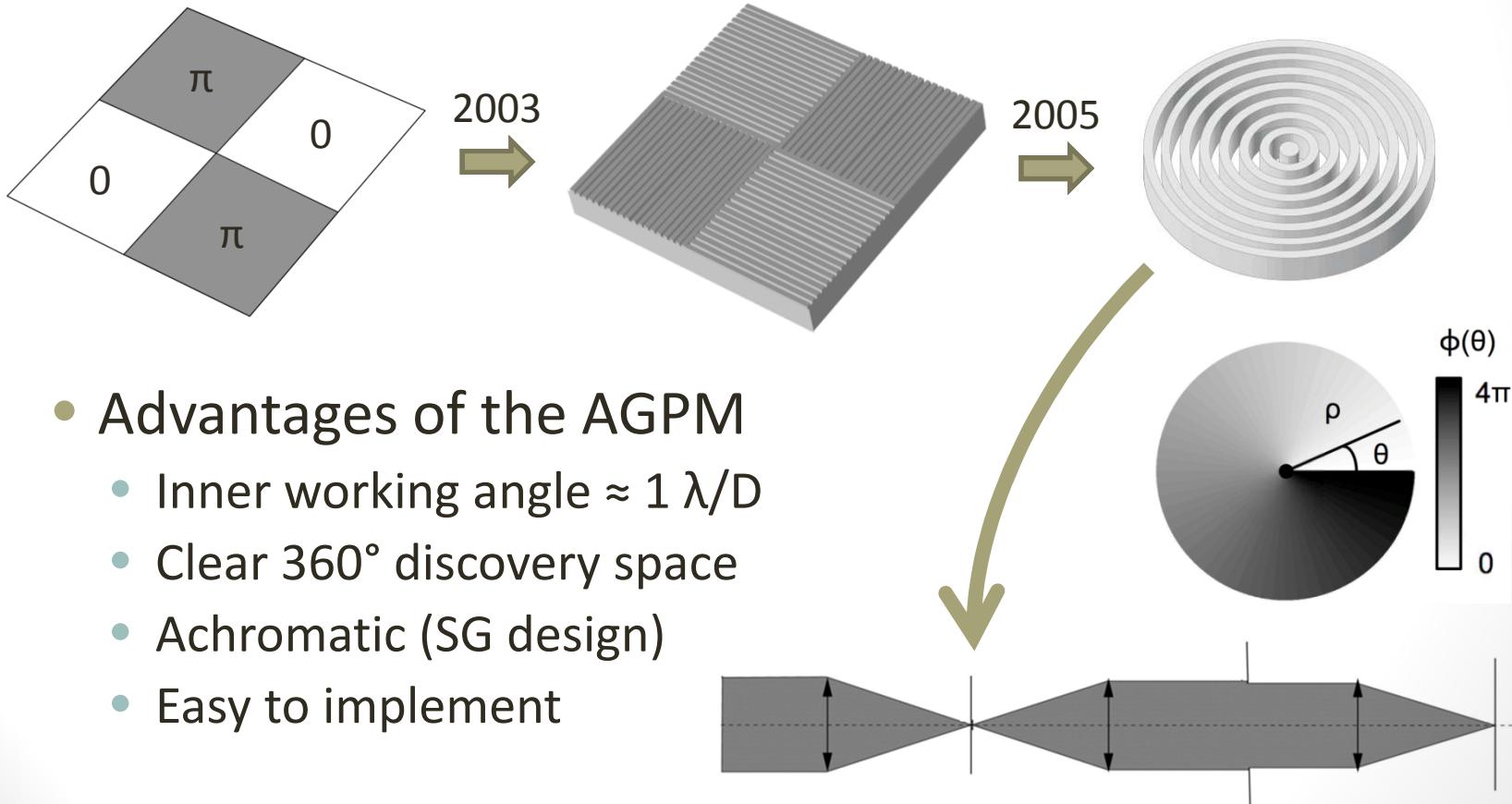
FNRS Research Associate  
Université de Liège

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# The birth of a concept

- FQPM → sub-wavelength gratings → Annular Groove PM

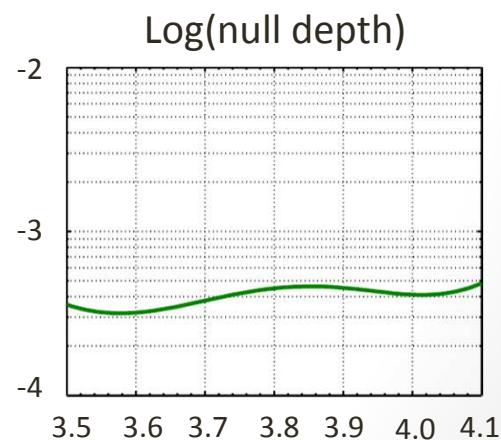
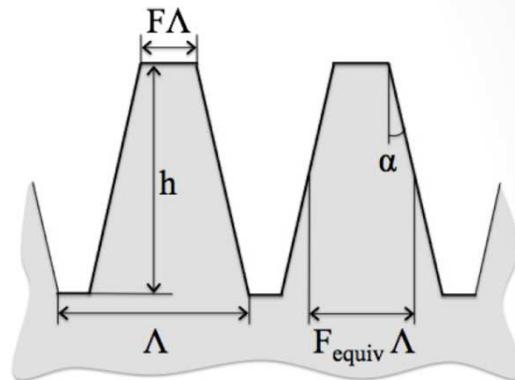
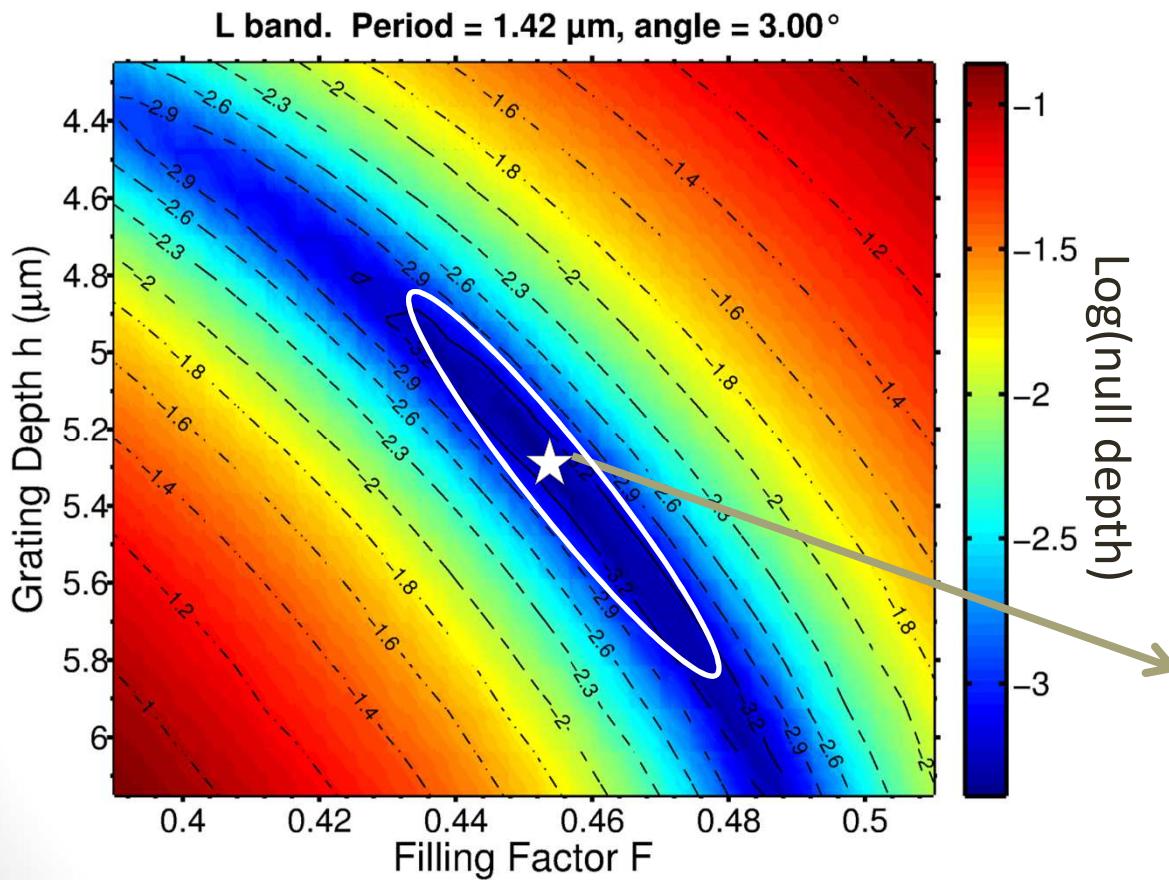


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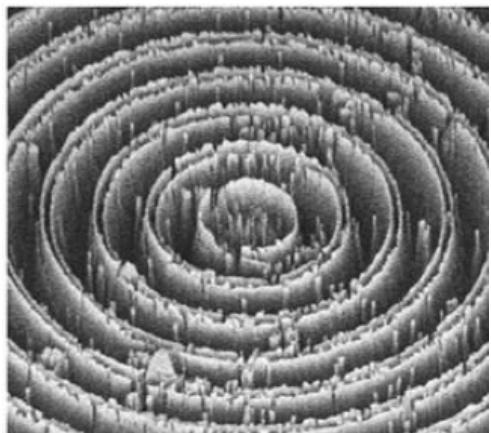
# Grating design/optimization



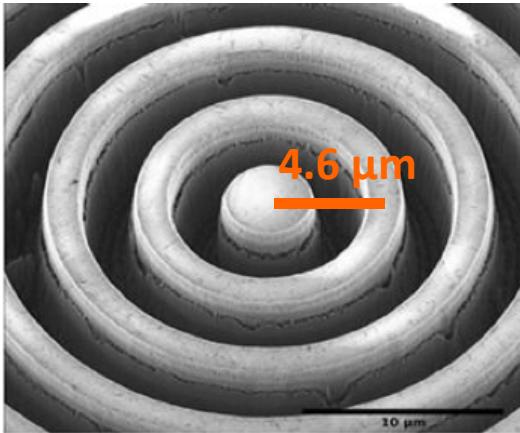
# Etching on CVD diamond

- Nanoimprint lithography + dry plasma etching
  - N band (grating period = 4.6  $\mu\text{m}$ )
  - L band (grating period = 1.4  $\mu\text{m}$ )

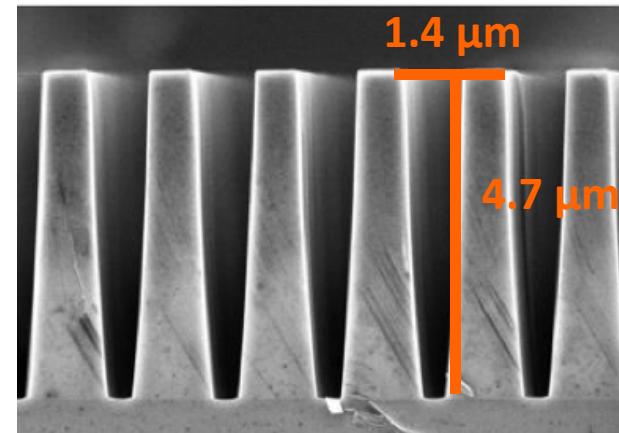
N band (Nov 2009)



N band (Feb 2012)



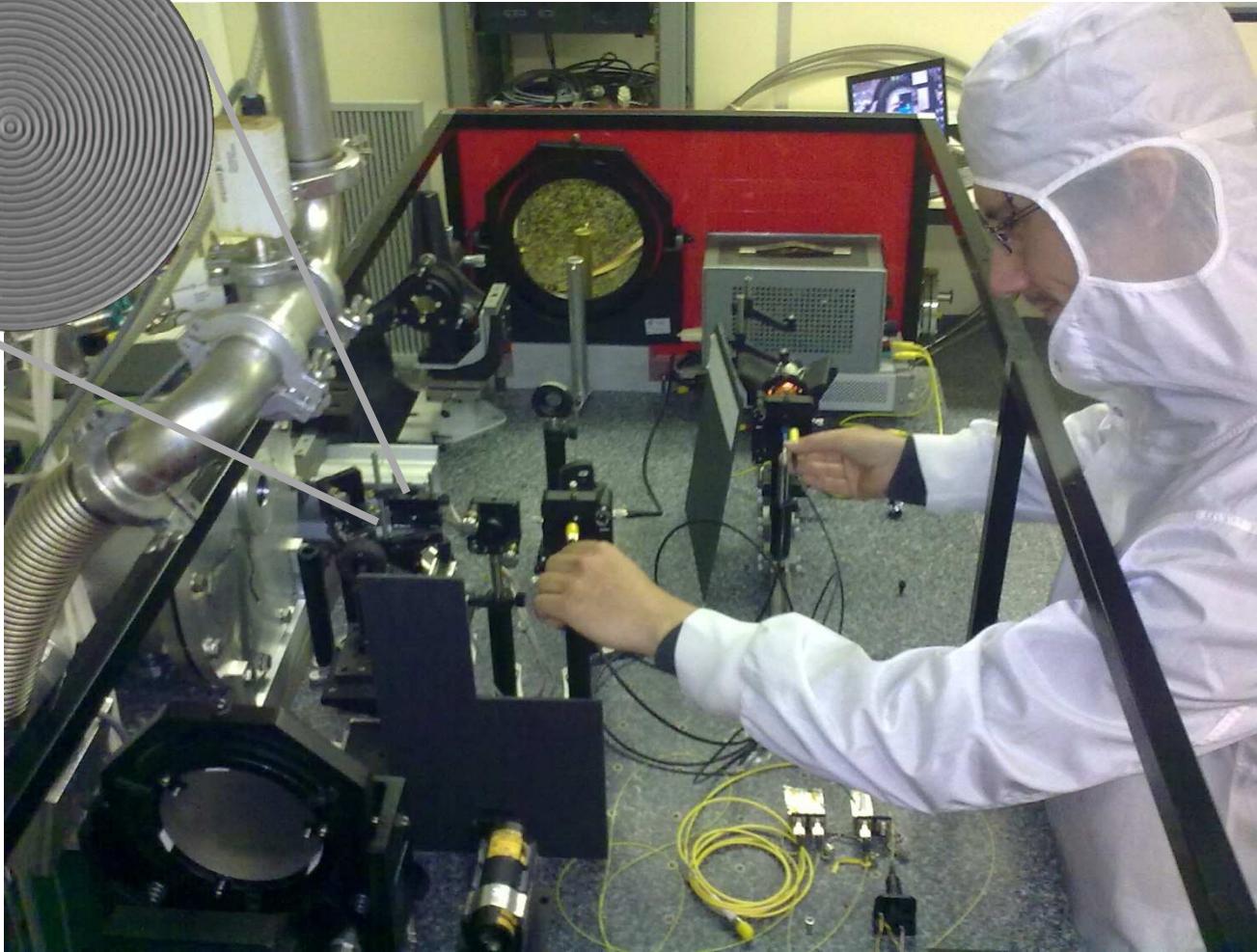
L band (Sep 2012)



- Parameters close to optimal ... need to test!

# Setting up the bench

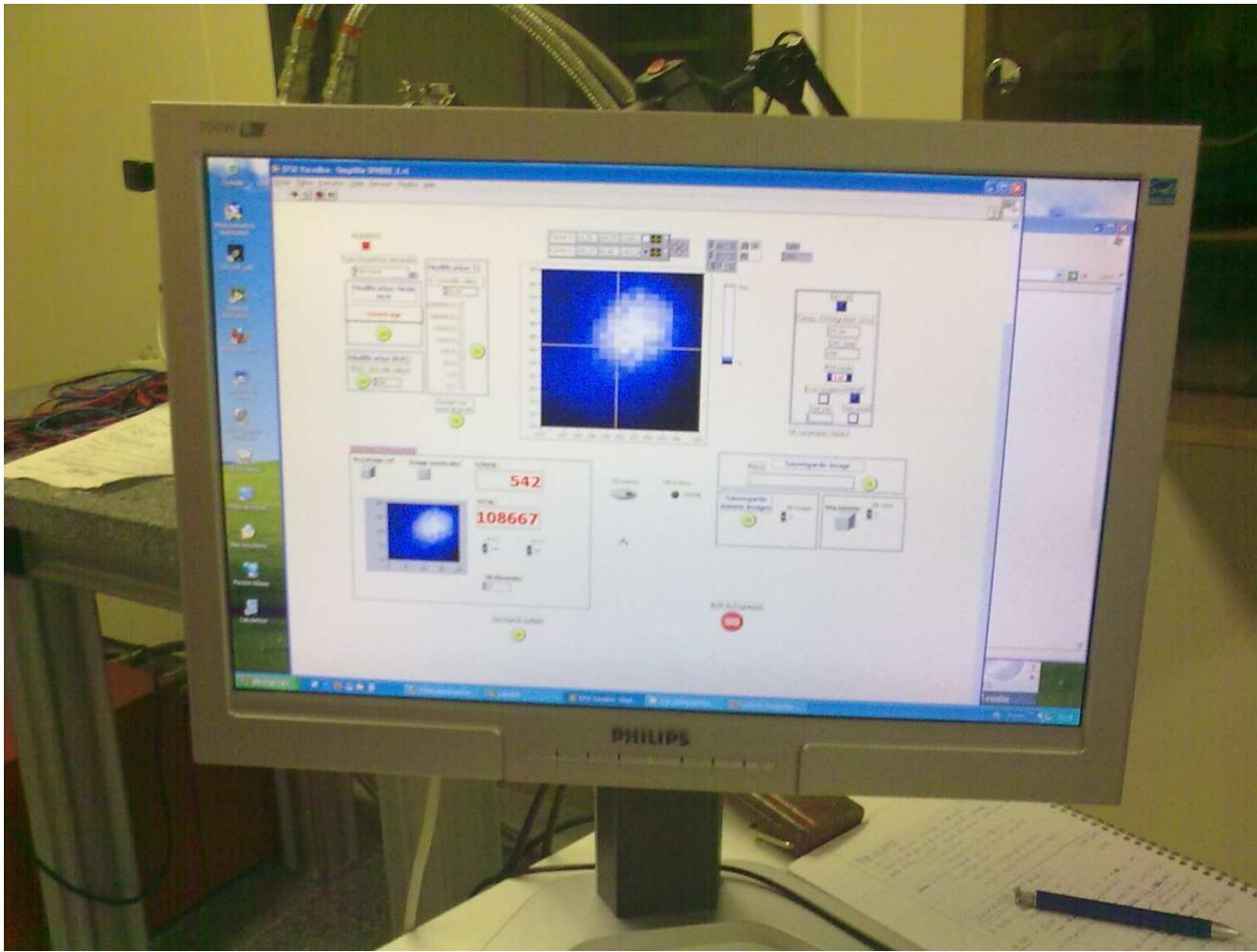
“Ycadire” @ Paris-Meudon



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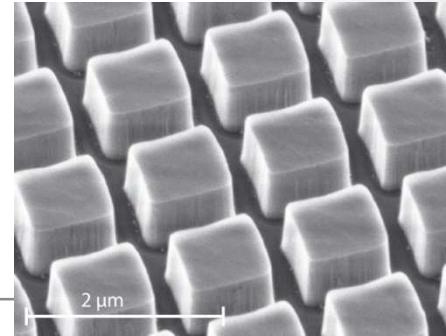
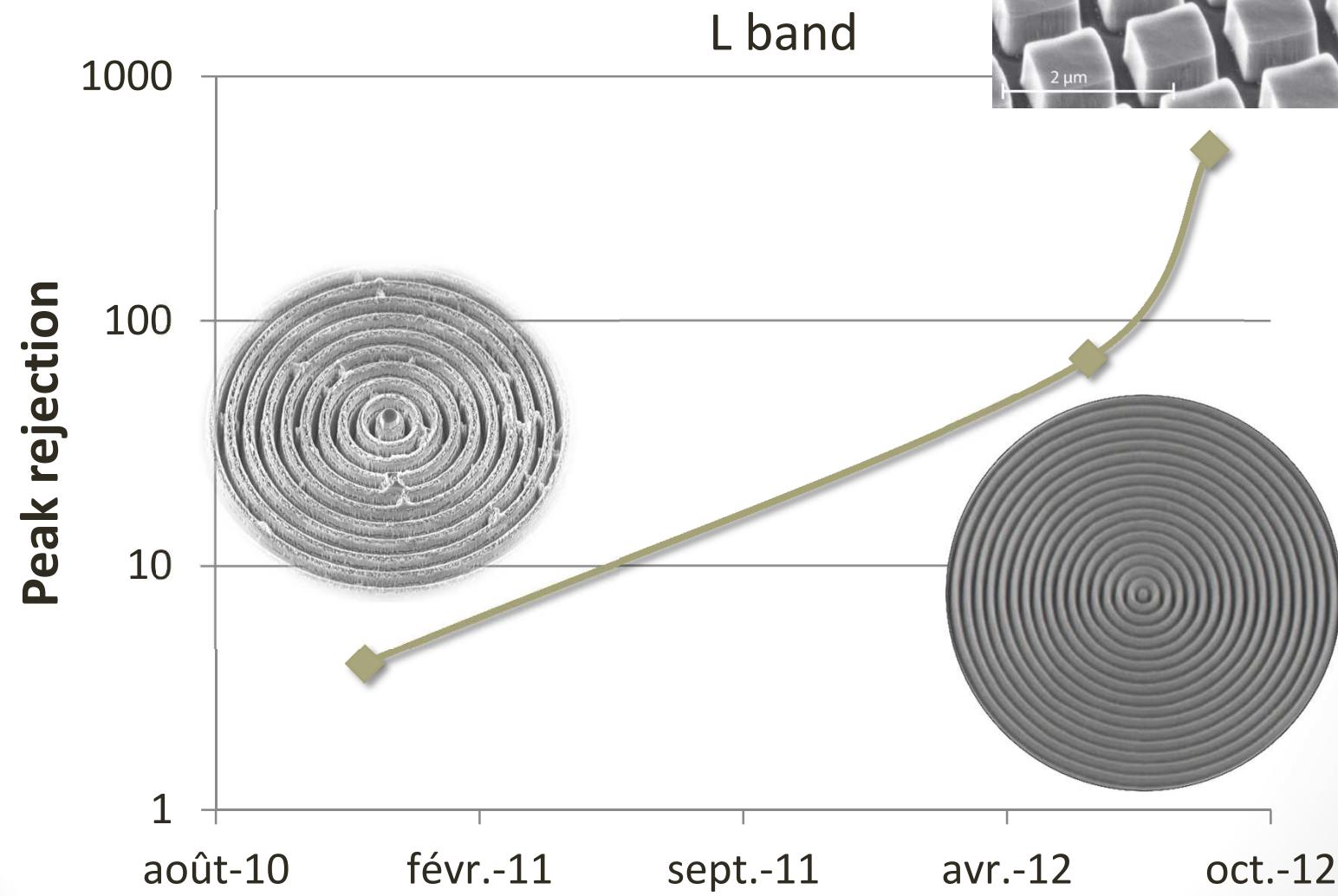
# Anguish...



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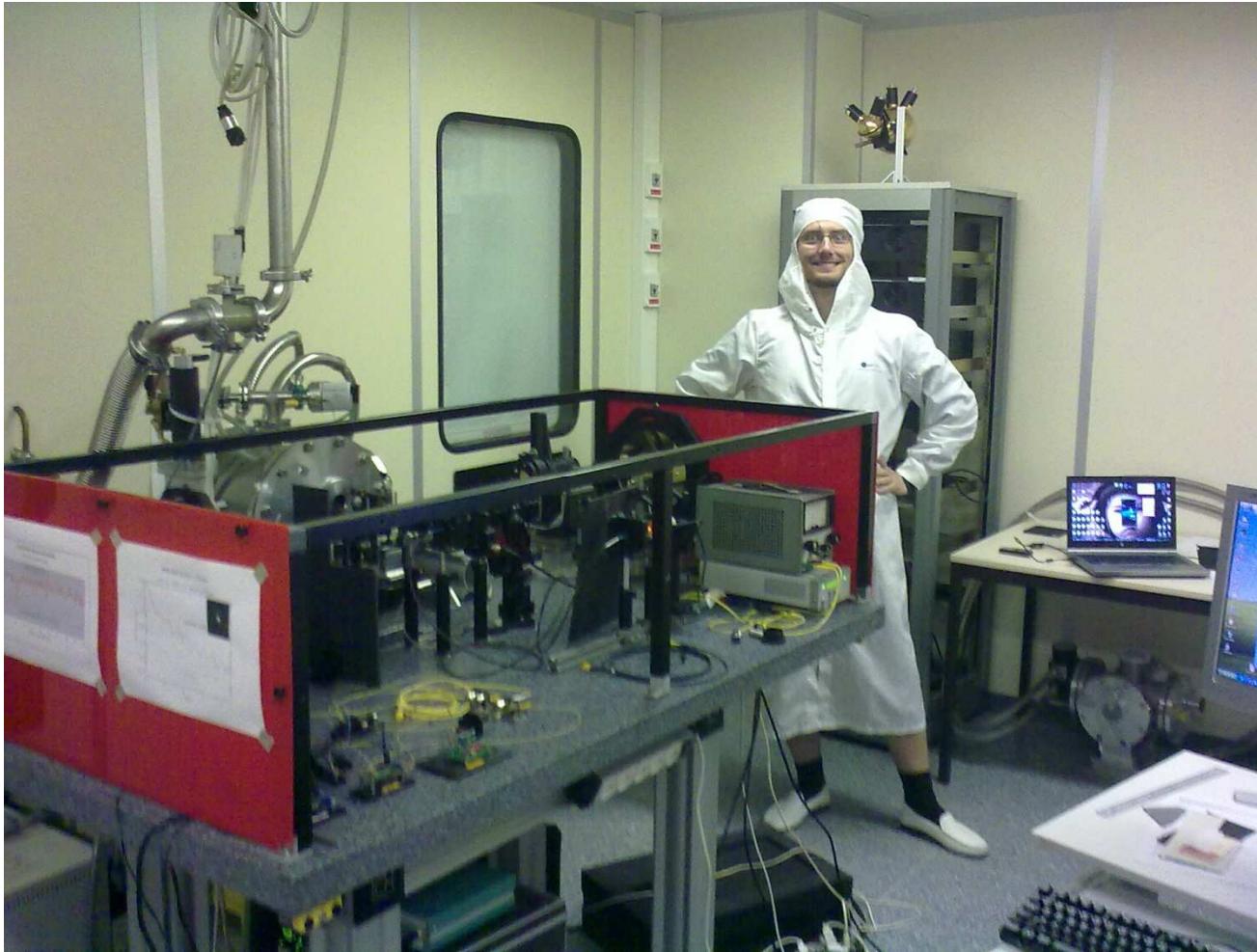
# High performance



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# Bliss!



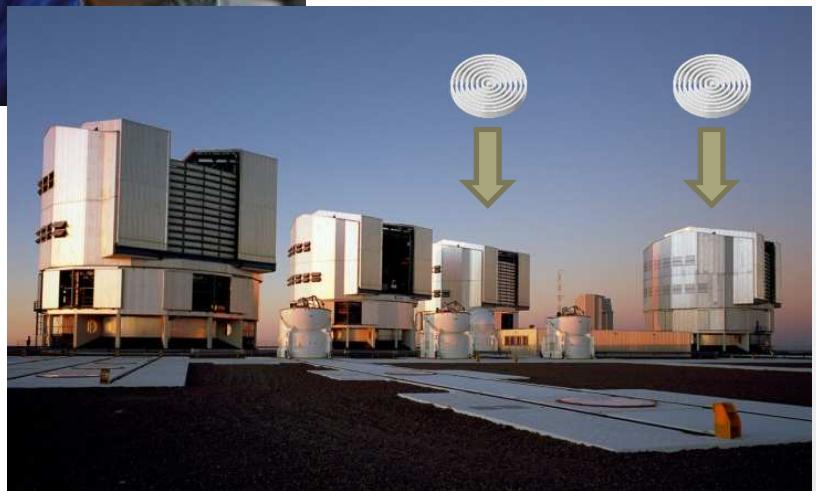
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# Installation at VLT

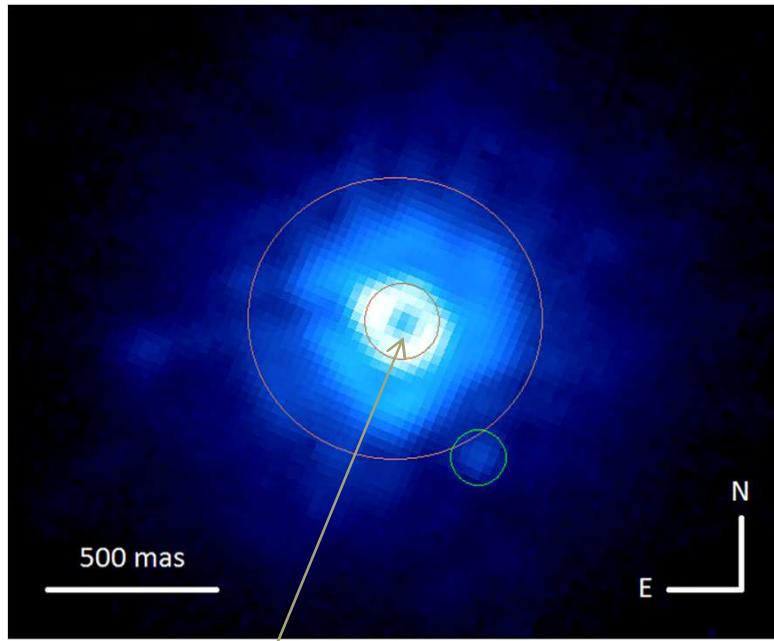
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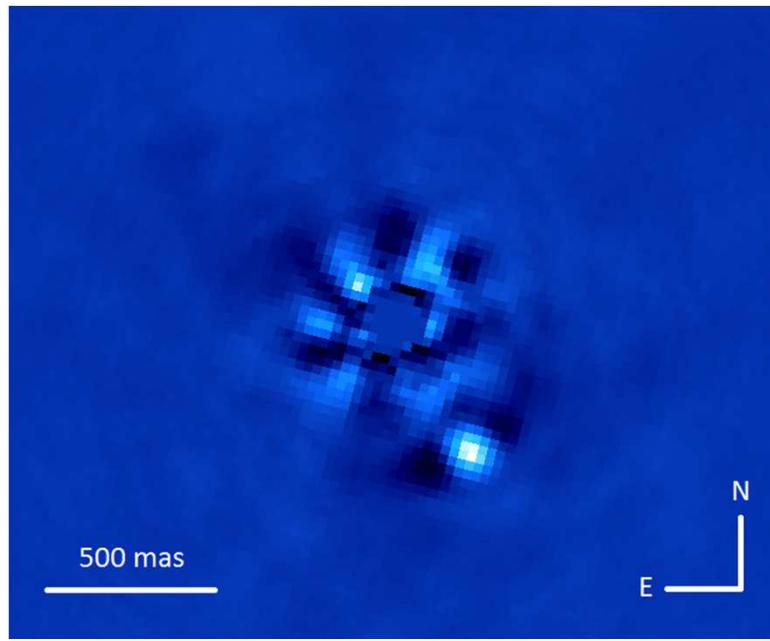
# NACO: science demonstration

Raw image of  $\beta$  Pic



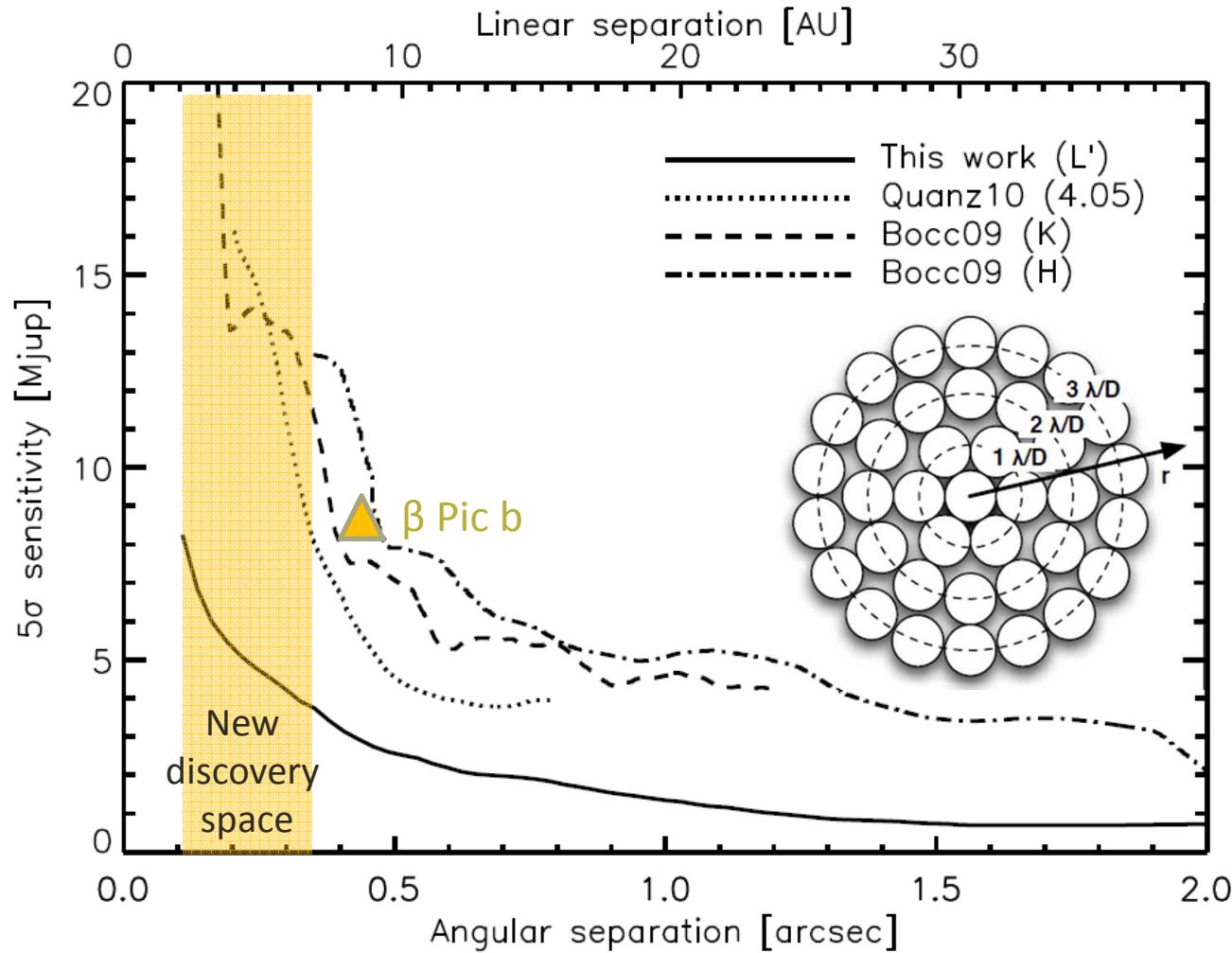
Peak rejection  $\sim 50:1$

Post-processed image



Absil et al. 2013

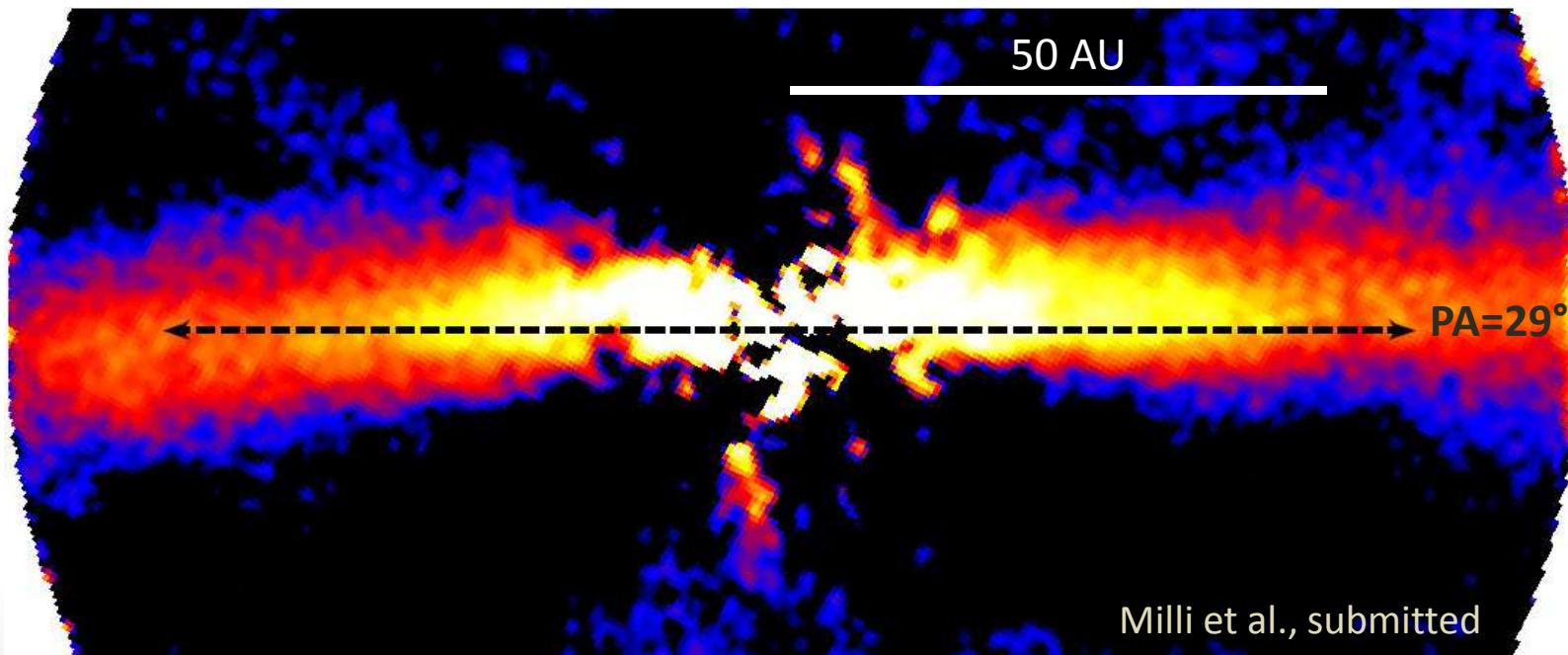
# Sensitivity to inner planets



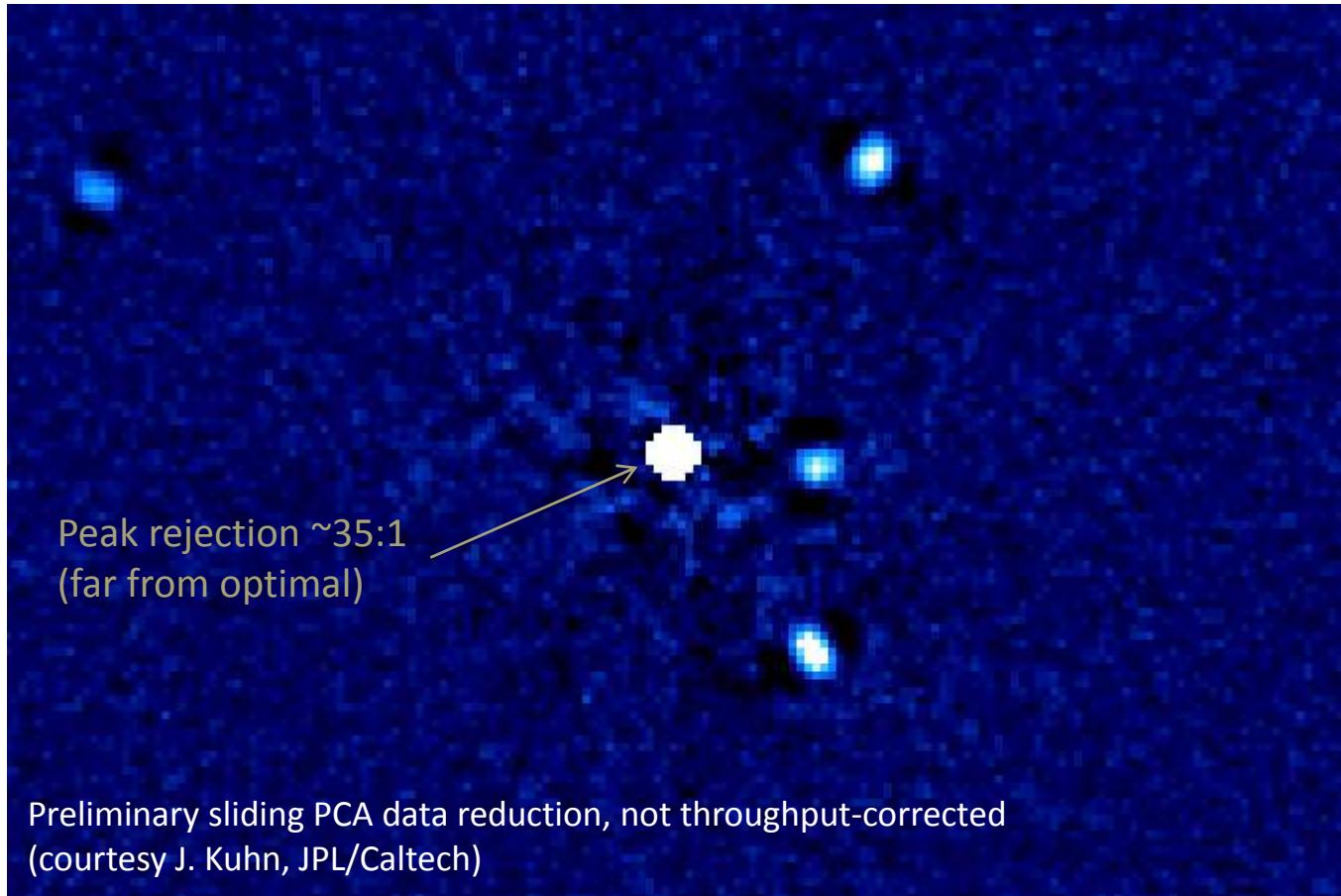
Absil et al. 2013

# The $\beta$ Pic disk at L band

- Warped, inner component
  - Disk detected down to 10 AU (0.4'')
- Spine offset and bowed (anisotropic scattering)



# First light with LBT/LMIRCam



# The VORTEX project



- **WP1: Exploitation of 1<sup>st</sup> generation AGPMs**
  - Install, test and optimize AGPMs on 10m-class telescopes
  - Perform the observations / analyze the data
- **WP2: Development of 2<sup>nd</sup> generation AGPMs**
  - Better L, M, N band AGPMs
  - Shorter wavelengths (K, H, ... where's the limit?)
  - Beyond topological charge = 2
- **WP3: Test and validation of new ideas**
  - Exploitation of photon orbital angular momentum (OAM)
  - Post-vortex speckle cancellation techniques
  - Optimal apodization

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# A vortex in your instrument?

- We're currently baking more L-band AGPMs

