Spatio-Temporal Analysis of Equatorial Ionospheric Scintillations in the Frame of Absolute GNSS Positioning Algorithms

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Introduction

Global Navigation Satellite Systems (**GNSS**) allow users to determine their positions anywhere on Earth at any time.

GNSS positioning techniques support many high-precision positioning applications for civil, military, scientific and industrial purposes.

Objectives

The global objective of this piece of research is to develop new **strategies** based on the combined use of **Multi-GNSS** in order to improve the **performances** of GNSS positioning during **lonos-**

Conclusions

This first part of our **Spatio-Temporal Analysis** highlights the increase of the **positive Global Spatial Autocorrelation** in the ionospheric scintillation data (S4). This result supports the second part of the approach including the following persectives and leading to Spatio-Temporal information for the design of the stochastic model in the frame of absolute GNSS positioning:

Recent GNSS developements and modernisation offer new powerful possiblities in **Earth Science**.



One of the most important **sources of errors** highly limiting the **accuracy** and **reliability** of high-precision GNSS positioning techniques is the **lonosphere**.

The **lonosphere** is a region of the upper atmosphere ionized **by solar radiations**. Among other effects, the lonosphere is responsible for the **diffraction** of GNSS signals producing **« lonospheric Scintillations »** and involving **major errors** in positioning algorithms. pheric Scintillations.

The **Stochastic Model** definition in the frame of absolute positioning applications is frequently reduced to the computation of diagonal (**variances**) elements assuming that non-diagonal elements (**covariances**) are negligible. This assumption could be challenged during lonospheric Scintillations and the design of a more complete and **advanced stochastic model** could improve positoning performances during such circumstances.

 $\Sigma = \begin{pmatrix} 1 & 0 & \dots & 0 \\ 0 & 1 & \dots & 0 \\ \dots & \dots & \dots & \dots \\ 0 & 0 & \dots & 1 \end{pmatrix} \implies \Sigma = \begin{pmatrix} \sigma_1^2 & 0 & \dots & 0 \\ 0 & \sigma_2^2 & \dots & 0 \\ \dots & \dots & \dots & \dots \\ 0 & 0 & \dots & \sigma_n^2 \end{pmatrix} \implies \Sigma = \begin{pmatrix} \sigma_1^2 & \sigma_{12} & \dots & \sigma_{1n} \\ \sigma_{21} & \sigma_2^2 & \dots & \sigma_{2n} \\ \dots & \dots & \dots & \dots \\ \sigma_{n1} & \sigma_{n2} & \dots & \sigma_n^2 \end{pmatrix}$

The combined use of **Multi-GNSS** (GPS, GLONASS, Galileo, etc.) constitutes a powerful tool for **Space Weather** monitoring. Based on a network of **ISMR** (Ionospheric Scintillation Monitoring Receivers) stations in **South America**, we are currently completing a **Spatio-Temporal Analysis** of **Ionospheric Scintillation** data with the aim of measuring and exploiting **spatial** and **temporal correlation** in GNSS positioning **algorithms**. Correlograms for the determination of the scale of spatial structures Local Spatial Autocorrelation tests — *« Hot Spot Analysis »* Interpolation Model (spatial covariance function) Time Correlation detection Spatio-Temporal Information vs Stochastic Model

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The use of multiple ISMR stations in a delimited area provides denser data set by combining the different lonospheric Pierce Point from several stations





Ionosphere

The **lonosphere** is a region of the upper atmosphere, from **about 85 to 600 km altitude**. The lonosphere is a **plasma**, ionized by solar radiations, whose the **free electron density** disturbs the propagation of radio signals.

The ionospheric plasma density produces a **refraction** effect on GNSS signals which mainly results in a variation of the signal propagation speed and a **bias** in the measurements achieved by a GNSS receiver.

The ionospheric plasma **inhomogeneities** are responsible for the **diffraction** of the GNSS signals involving rapid fluctuations (**scintillations**) of the **phase** (**Phi6o**) and amplitude (**S4**).



Spatio-Temporal Analysis

Spatial Autocorrelation



Ionospheric Scintillation

0.25f

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The time variation of the **global mean S4 parameter** computed from all the ISMR stations is a good indicator of the presence and intensity level of **lonospheric Scintillations**. Based on the observation of S4 and Phi6o measurements (ISMR), the effects of ionospheric scintillations seem to be colocated in the same portion(s) of the sky at a given time with eventually a time lag due to relative displacements of the ionospheric irregularities and satellites.

Global Spatial Autocorrelation (GSAC) can be detected by satistical hypothesis tests of the well-known Moran's I and Geary's C indices. During the scintillation event (ooh-ogh UTC), both indices are reaching the limits of their respective confidence intervals indicating the presence of a significative positive Global Spatial Autocorrelation in the data during this period.

Additional Spatial Analysis tests at specific times will give advanced information about the scale and locations of spatial structures (« hot

GNSS

Ionospheric Scintillations are highly limiting for absolute GNSS positioning (SPP/PPP) due to several severe impacts.



